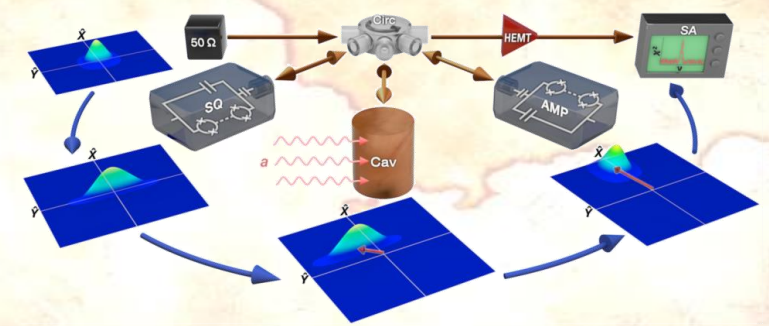
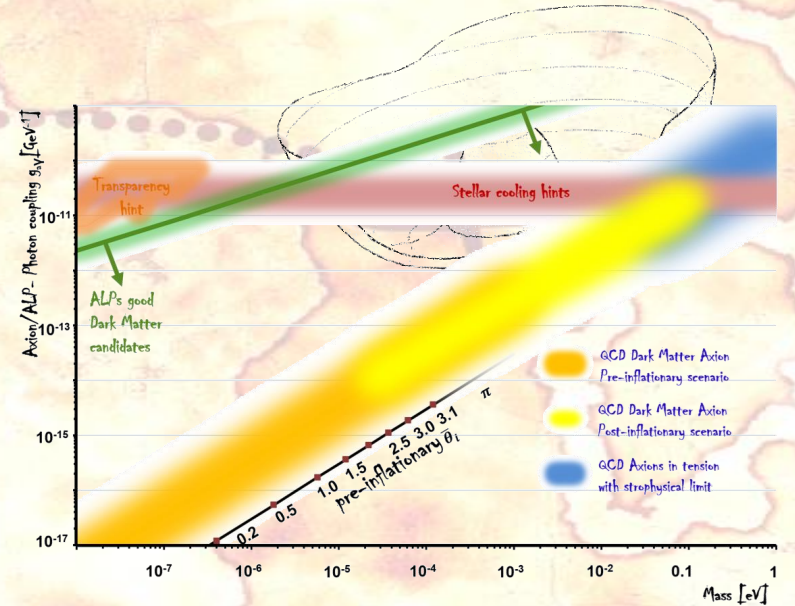
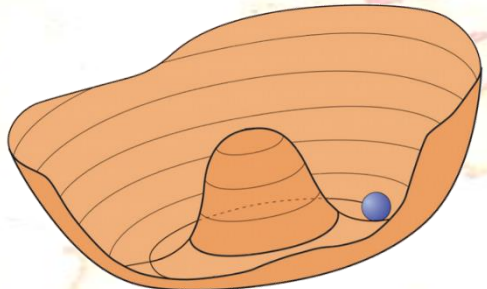


Navigating through the Axion & ALP's landscape:

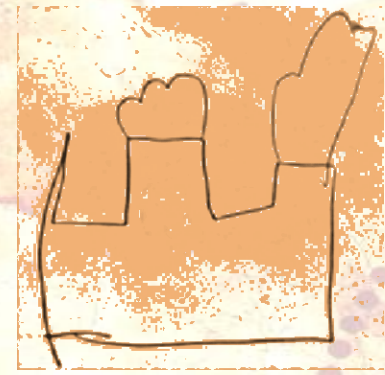


Béla Majorovits

- Wanted !!
Axion profile
- Navigating the axiverse:
explore the vast range!
- The bounty hunters:
Chose your arms!
- Surveying terra incognita



WANTED



the QCD Axion

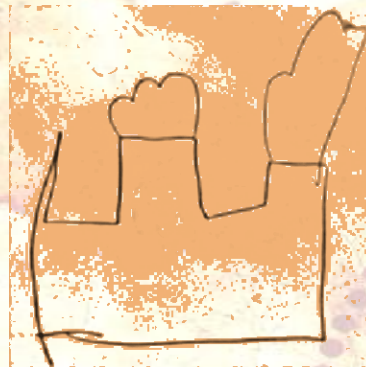
Guilty for solving the **strong CP problem!**

$$\bar{\Theta} \cdot \frac{\alpha_s}{8\pi} G_{\mu\nu a} \tilde{G}_a^{\mu\nu} \in \mathcal{L}_{\text{QCD}}$$

$$\bar{\Theta} = \Theta - \arg \det M_q \quad -\pi < \bar{\Theta} < \pi$$

Random phase from Θ -vacuum phases from Yukawa coupling: CKM matrix

WANTED



the QCD Axion

Guilty for solving the **strong CP problem!**

Expect nEDM: $d_n \sim \bar{\theta} \cdot 10^{-16} \text{ e cm}$

nEDM limit: $d_n < 3 \cdot 10^{-26} \text{ e cm}$

Phys. Rev. Lett. 124, 081803 (2020)

$$\rightarrow \bar{\theta} = \theta - \arg \det M_q < 10^{-10}$$

Random phase
from θ -vacuum

phases from Yukawa coupling:
CKM matrix

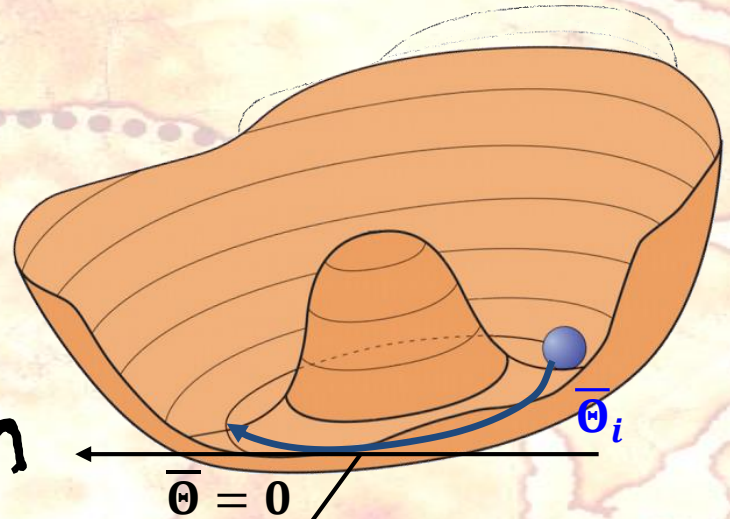
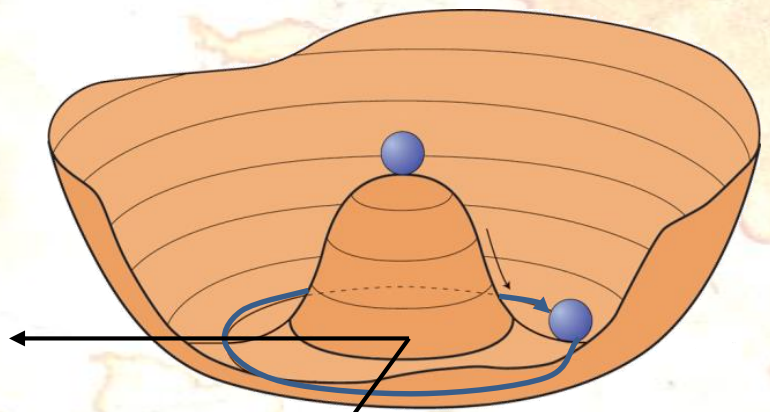


WANTED

Peccei Quinn symmetry breaking of $U(1)$

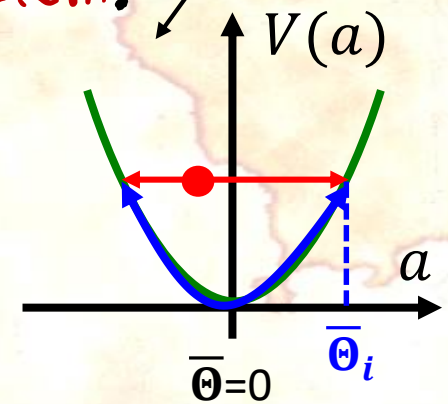
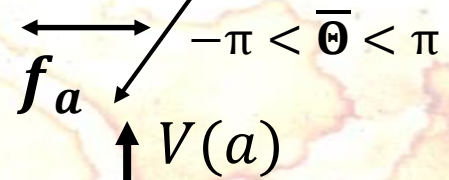
QCD:

Explicit symmetry breaking



the QCD Axion

Guilty for solving the **strong CP problem!**

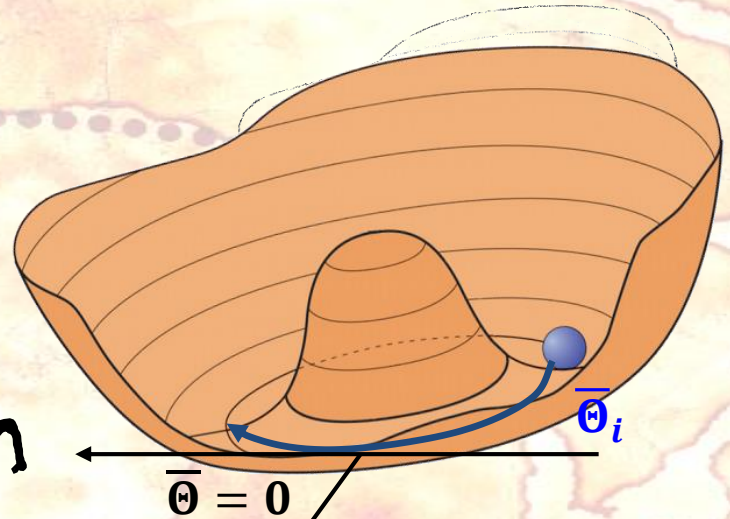
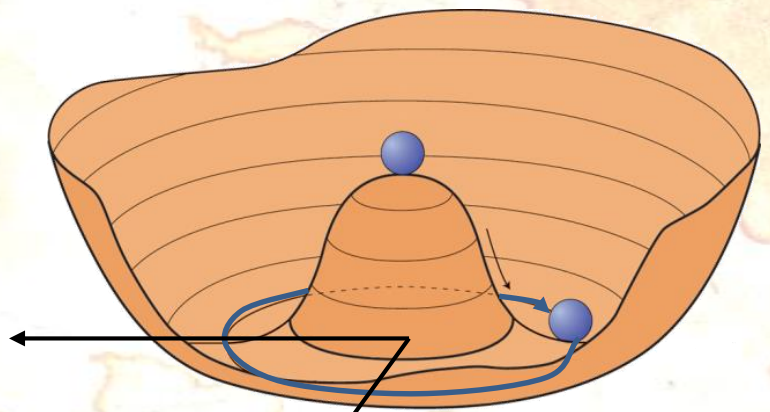


$$m_a \sim 5.7 \mu\text{eV} \frac{10^{12} \text{ GeV}}{f_a}$$

WANTED

QCD:

Explicit symmetry breaking



the QCD Axion

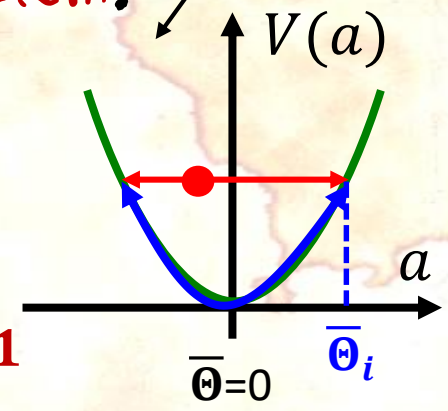
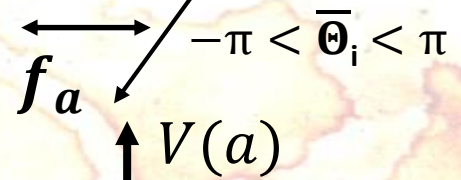
Guilty for solving the strong CP problem!

master of disguise!

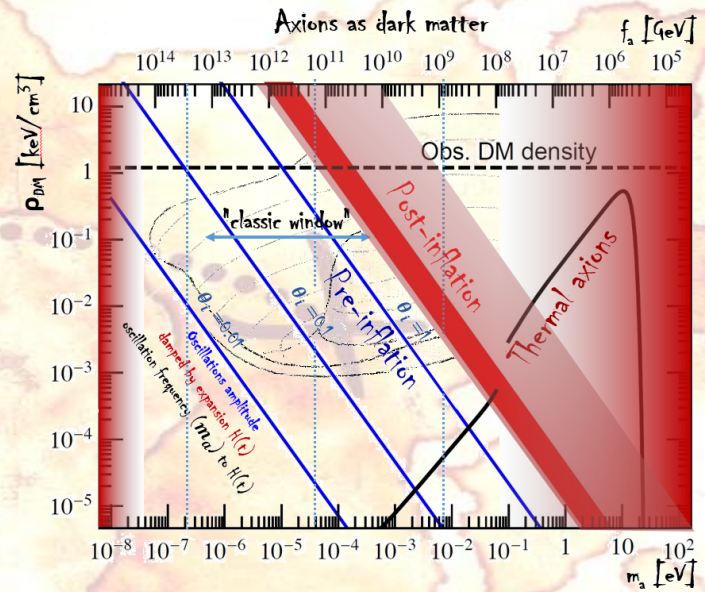
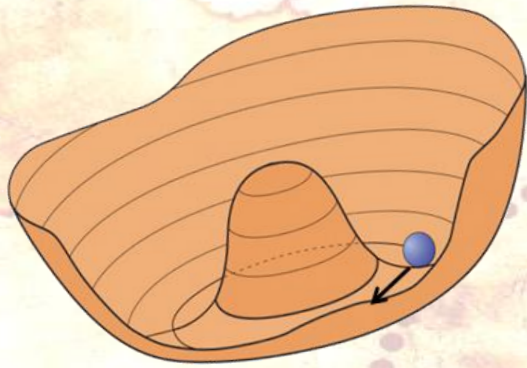
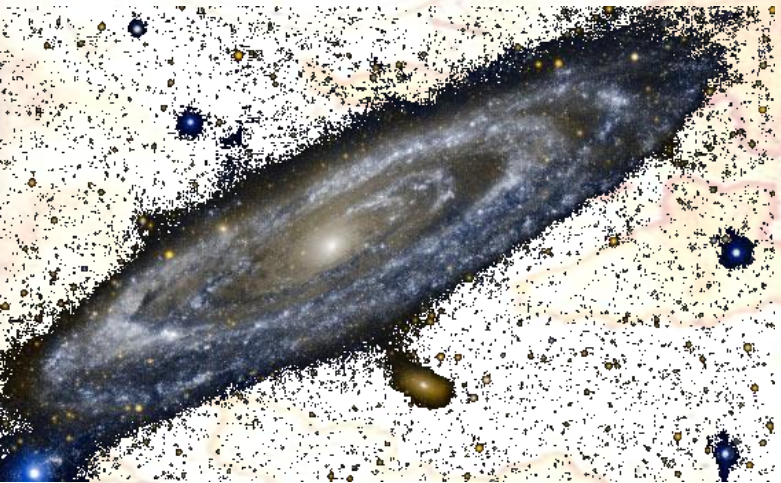
$$m_a \sim 5.7 \mu\text{eV} \frac{10^{12} \text{ GeV}}{f_a}$$

Coupling to photons suppressed:

$$g_{a\gamma} \propto f_a^{-1}$$



WANTED



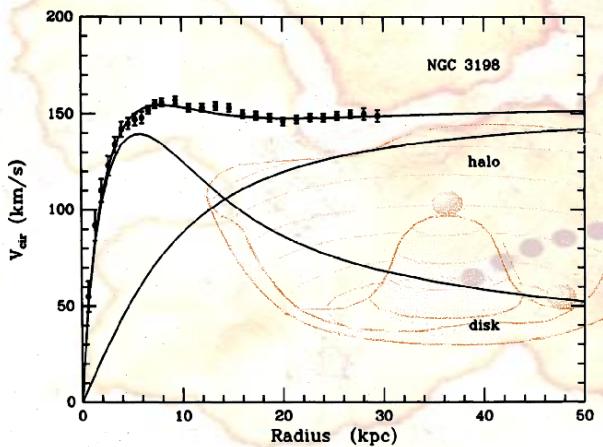
the QCD Axion

Guilty for solving the **strong CP problem!**
master of disguise!

$$m_a \sim 5.7 \mu\text{eV} \frac{10^{12} \text{ GeV}}{f_a}$$

$$g_{a\gamma} \propto f_a^{-1}$$

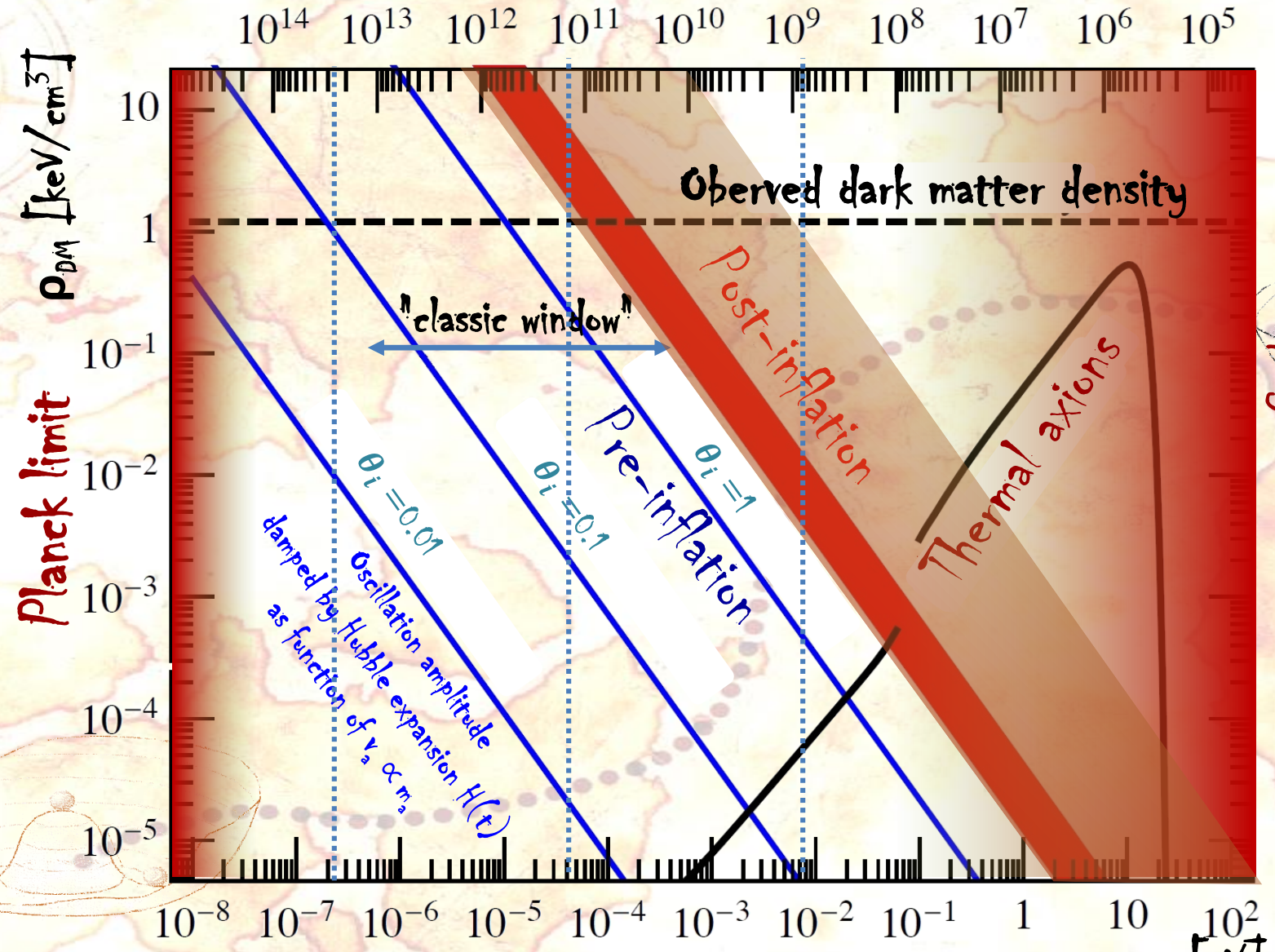
DISTRIBUTION OF DARK MATTER IN NGC 3198



REWARD: Solve the **DM crisis?**

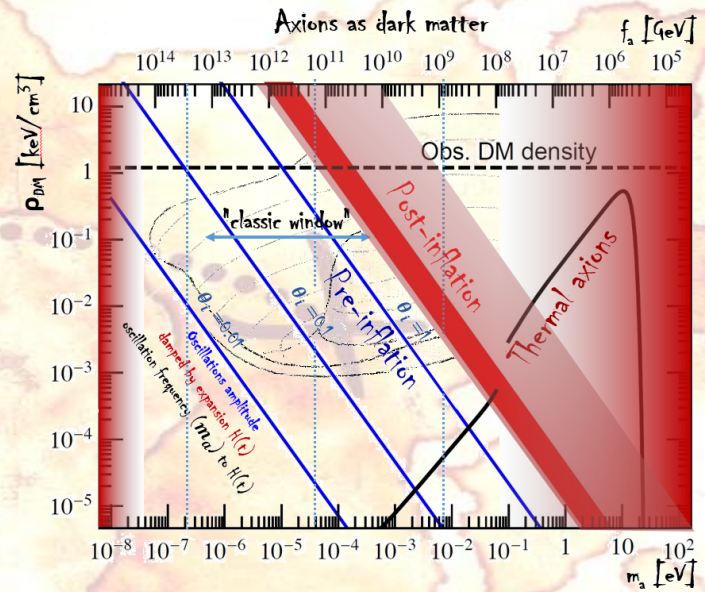
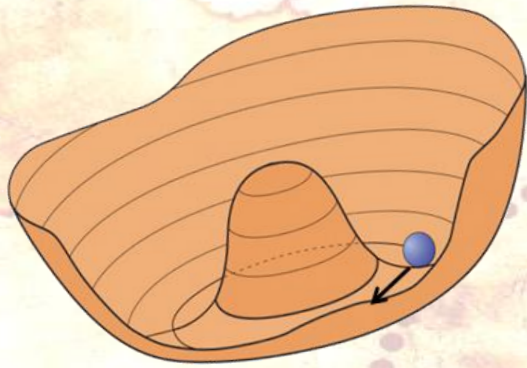
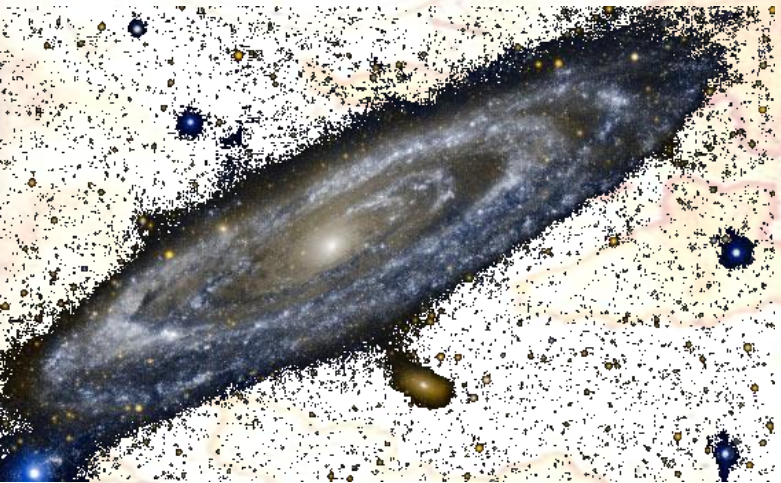
Axions as dark matter

f_a [GeV]



REWARD: Solve the **DM** crisis?

WANTED



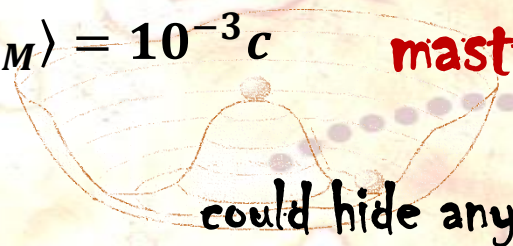
the QCD Axion

Galactic dark matter:

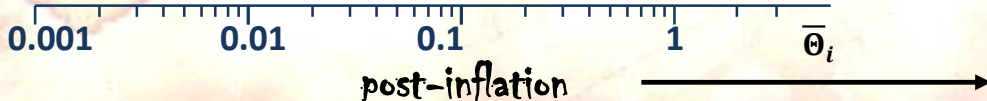
Guilty for solving the **strong CP problem!**

$$\langle v_{DM} \rangle = 10^{-3} c$$

master of disguise: (nearly) invisible & wave-like!

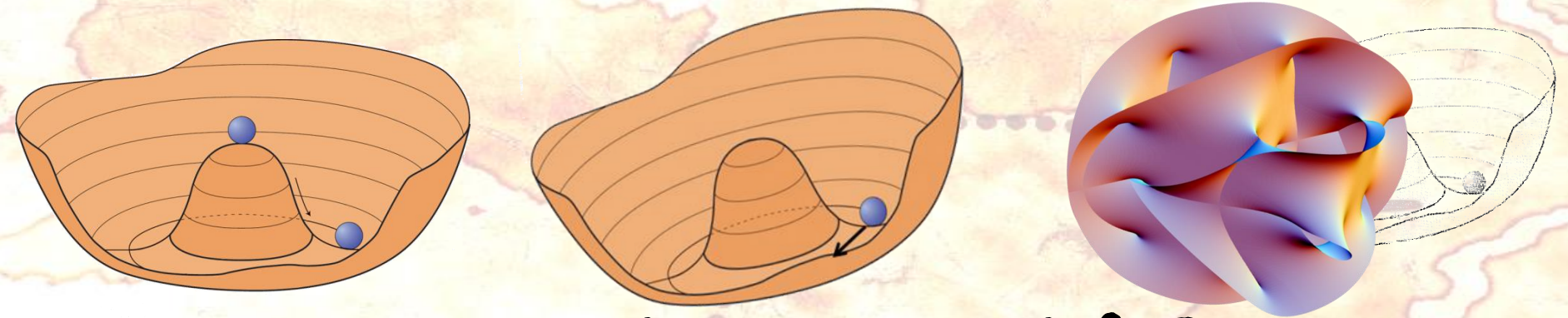


could hide anywhere
pre-inflation





WANTED



the QCD Axion & ALPs

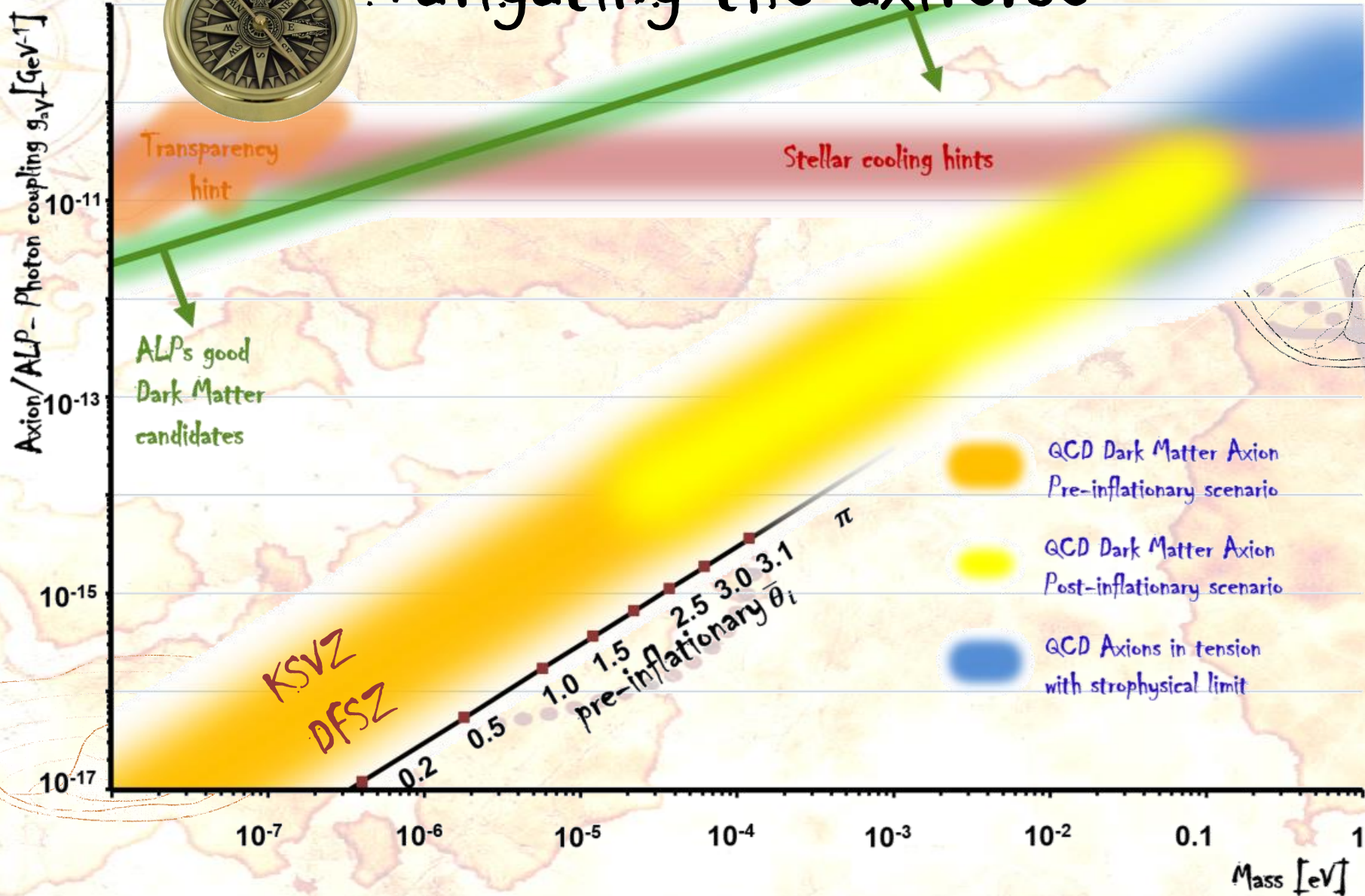
Guilty for solving the **strong CP** problem!

Axions: **masters of disguise!**

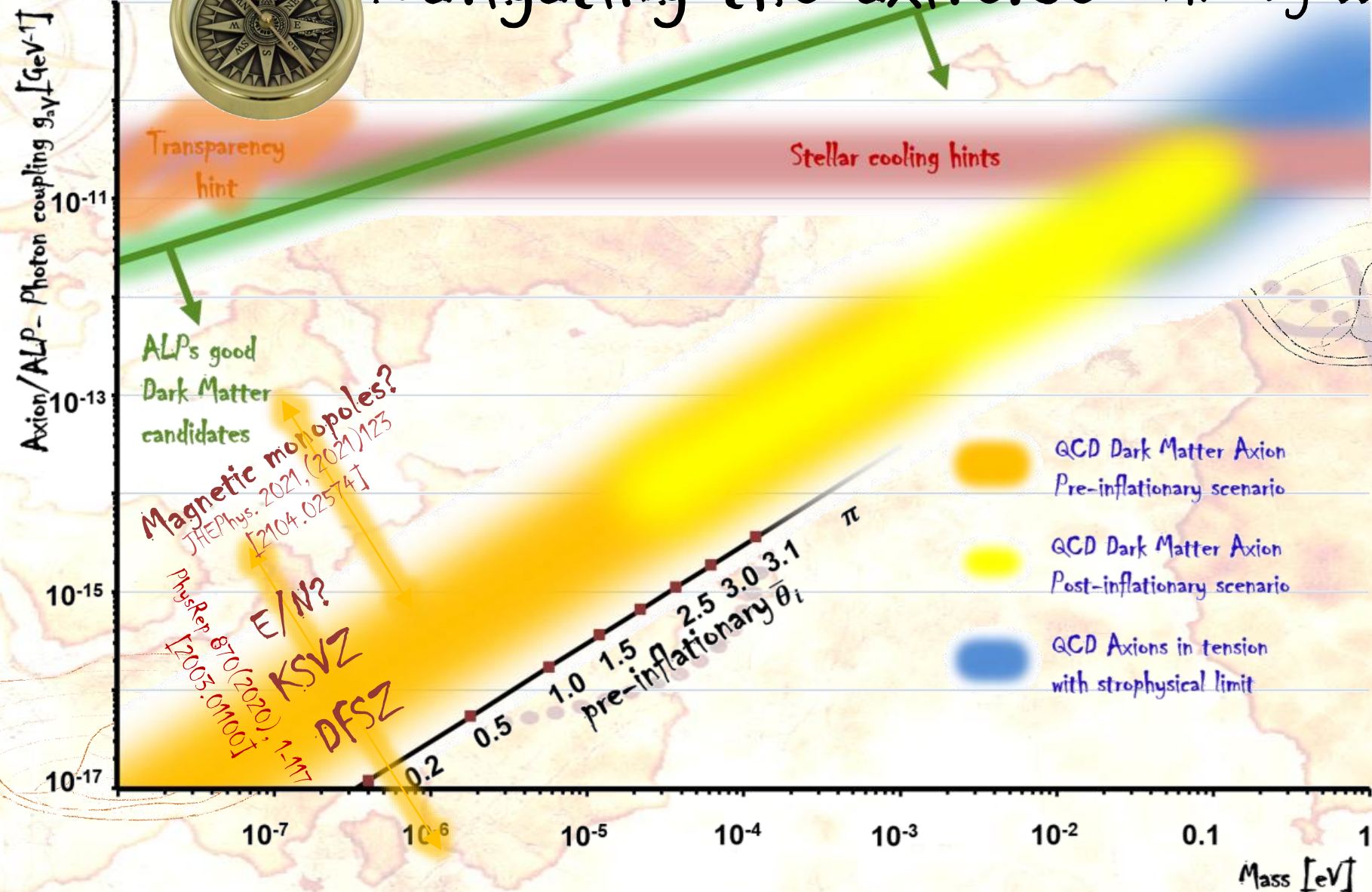
Prime suspect for cause of **DM** crisis!

Compactification of dimensions \rightarrow **Axiverse!**

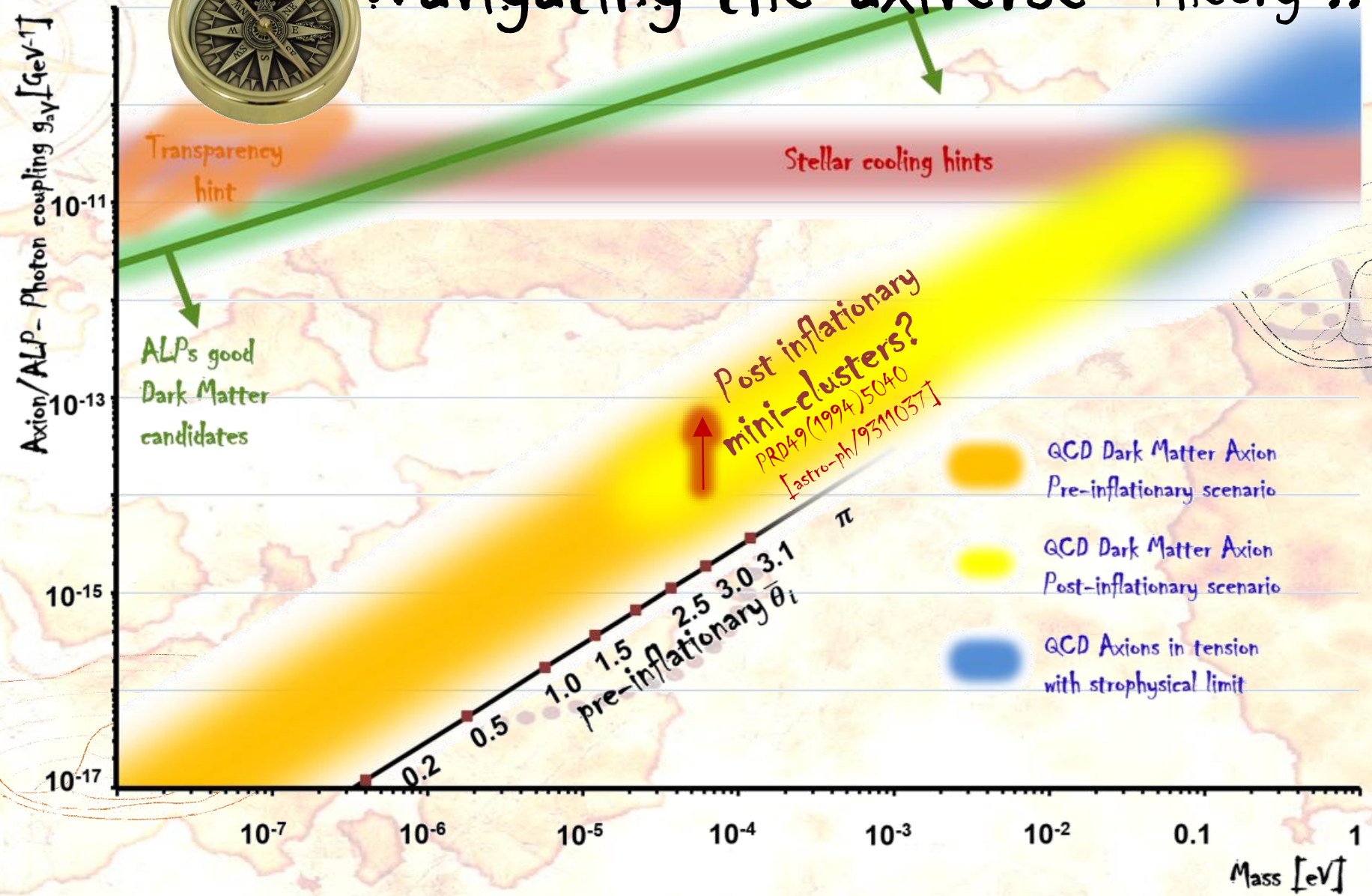
Navigating the axiverse



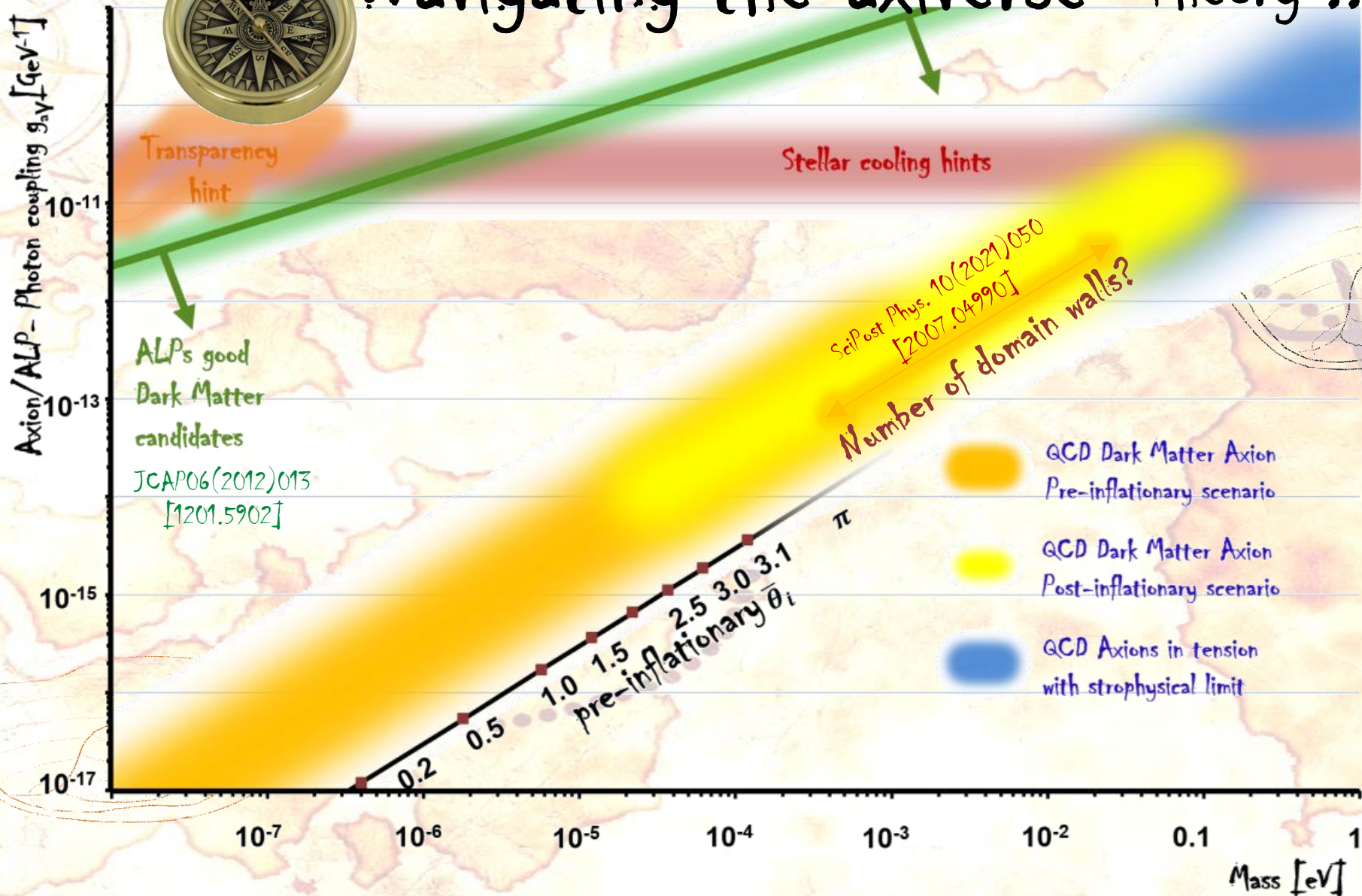
Navigating the axiverse Theory !!



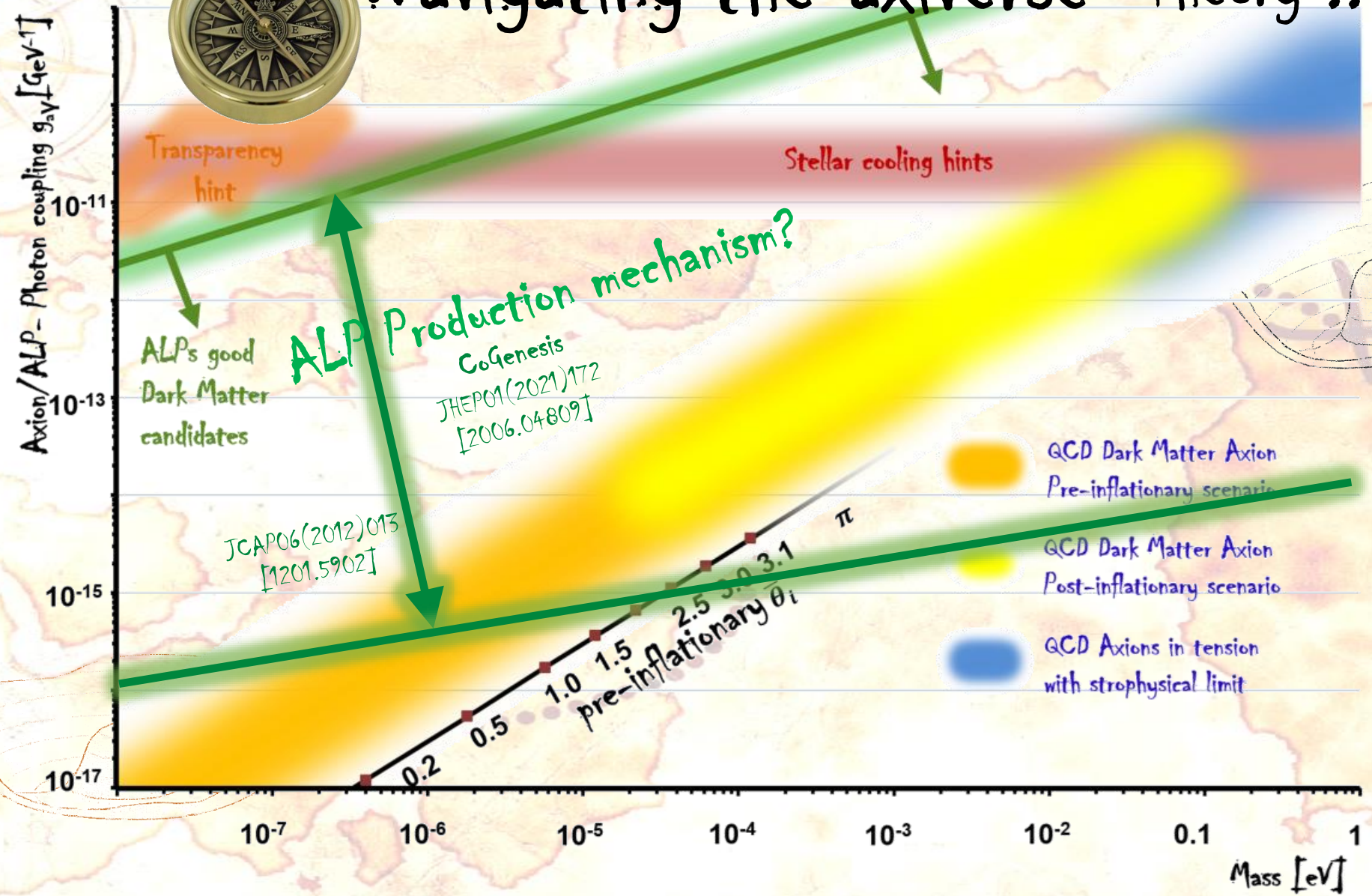
Navigating the axiverse Theory !!



Navigating the axiverse Theory !!



Navigating the axiverse Theory !!



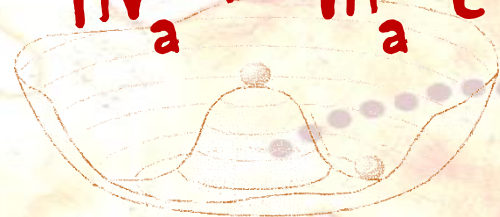
The bounty hunters' most important arms: The inverse Primakoff effect



$$\nabla \times B - \dot{E} = J + g_{a\gamma} B \dot{a}$$

Power $\propto B^2$

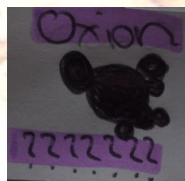
$$h\nu_a = m_a c^2$$



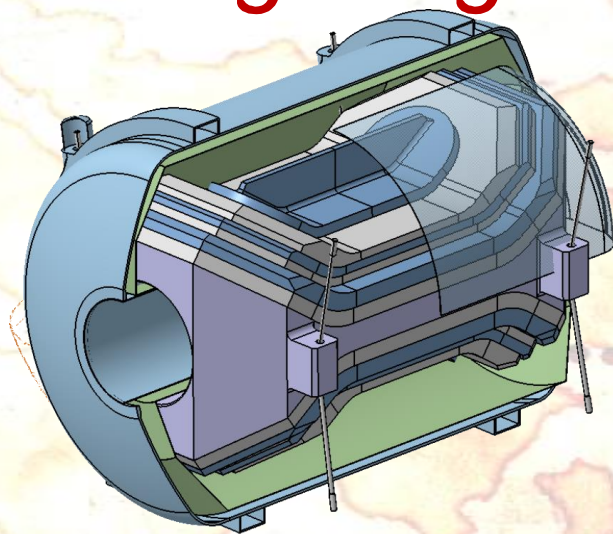
Axion in static B-field
-> E-field oscillation!

B-field

The bounty hunters' most important arms: The inverse Primakoff effect



Strong magnets!



Axion in static B -field
 \rightarrow E -field oscillation!

$$\nabla \times B - \dot{E} = J + g_{a\gamma} B \dot{a}$$

B -field

Power $\propto B^2$

The bounty hunters' most important arms: Axion induced E-field oscillations:

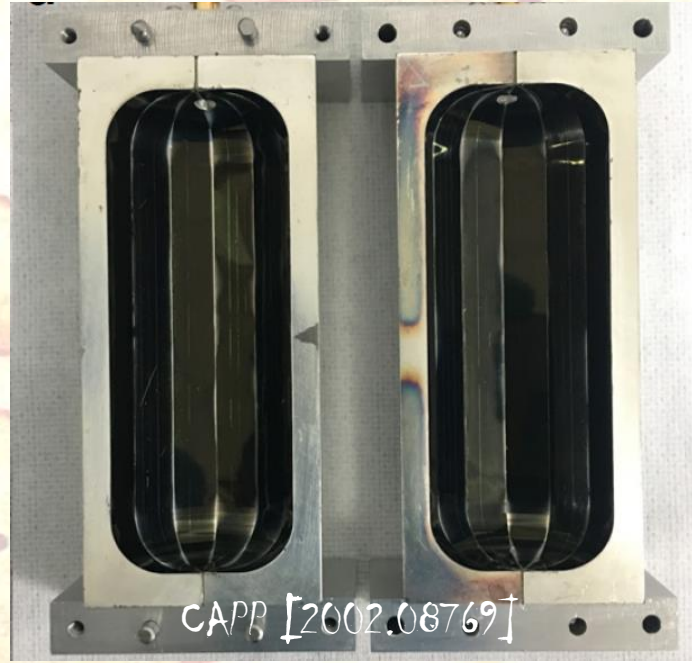
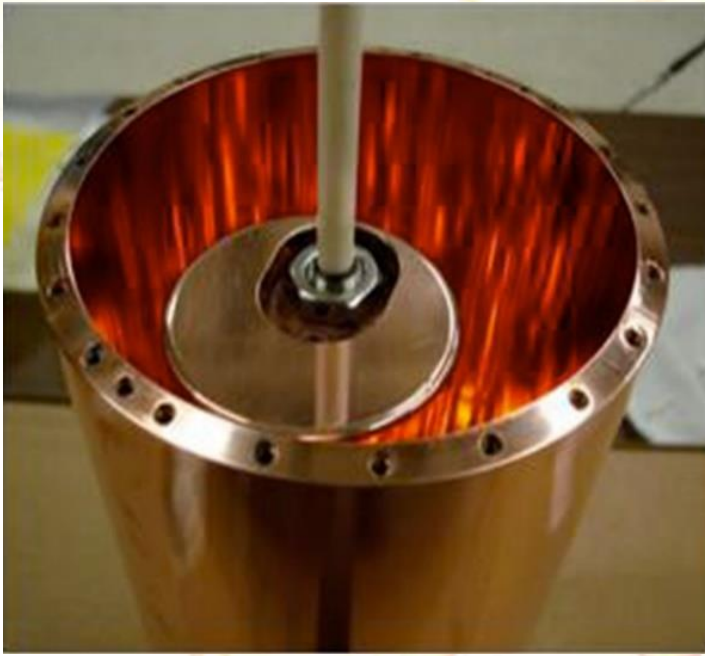
-> exploit wave mechanics & boost E-field

Tunable high Q- resonators
ADMX & Haystack

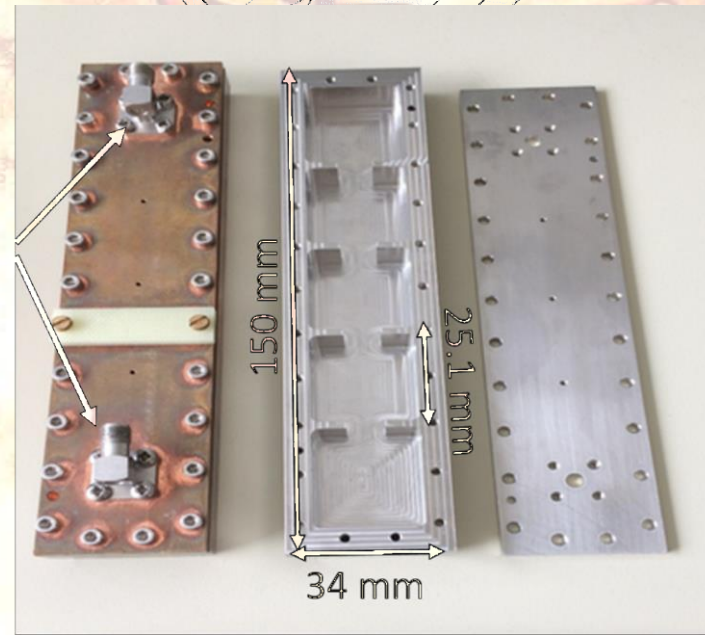
Superconducting cavities



Split cavities



CAPR [2002.08769]



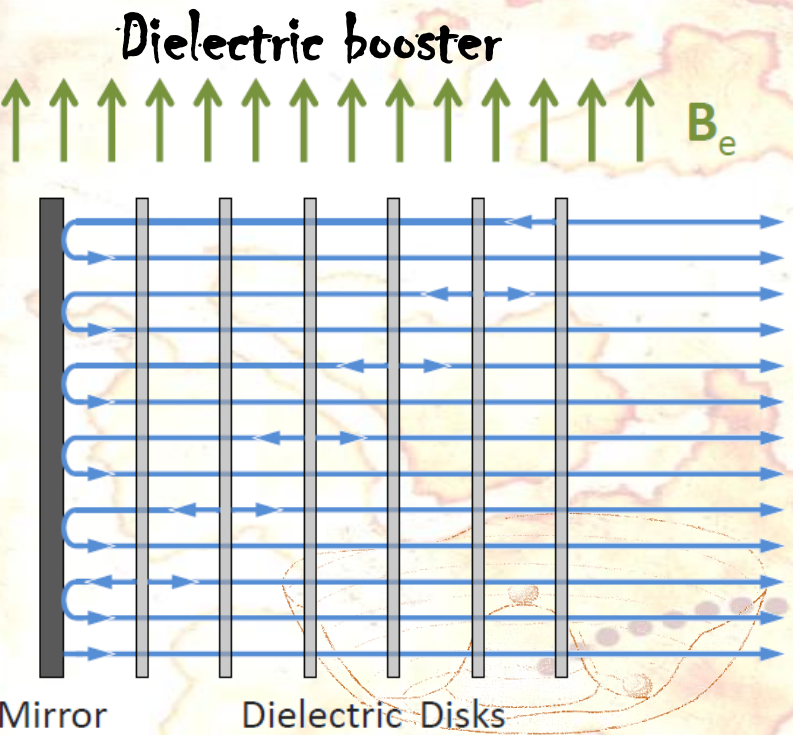
RADES [2104.13798]

PRL 124(2020)101303 PRL 118(2017)061302
Axions (and Axion Like Particles)

$$\text{Power} \propto B^2 \cdot V \cdot Q$$

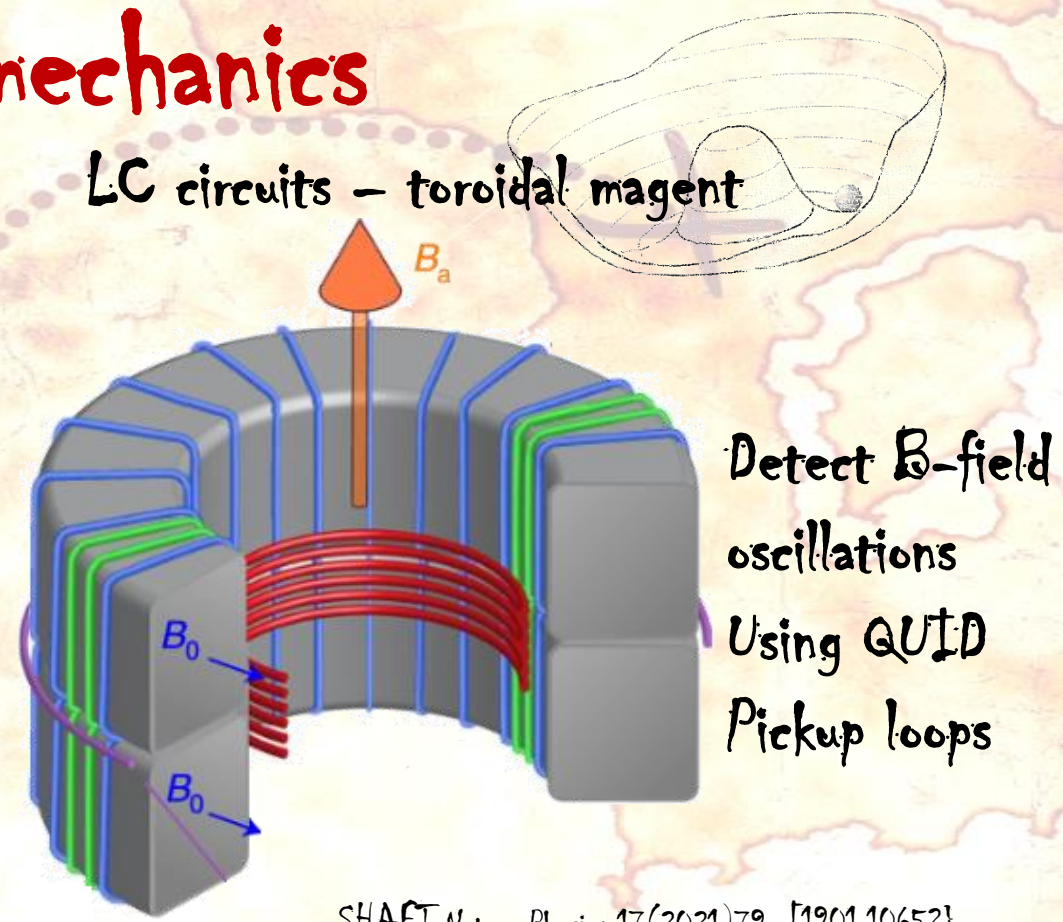
The bounty hunters' most important arms: Axion induced E-field oscillations:

-> exploit wave mechanics



MADMAX PRL118 (2017)091801

Axions (and Axion Like Particles)



SHAFT Nature Physics 17(2021)79 [1901.10652]

Laso DM radio - ADMX SLIC - ..

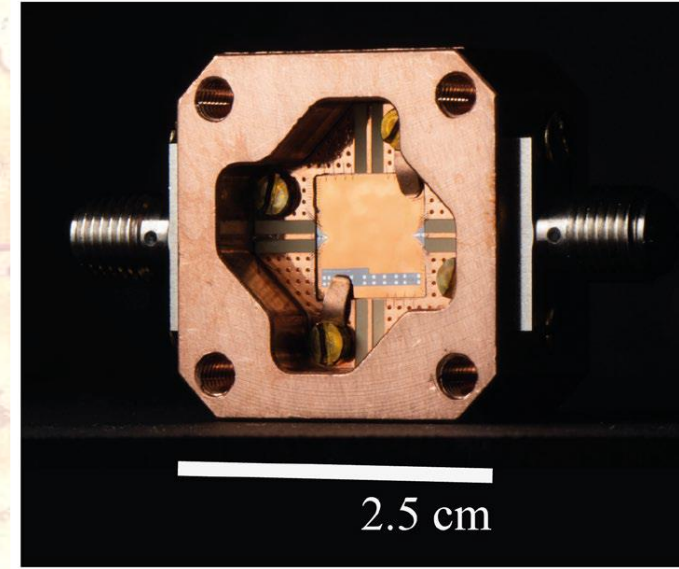
The bounty hunters' most important arms:

Low noise amplifiers

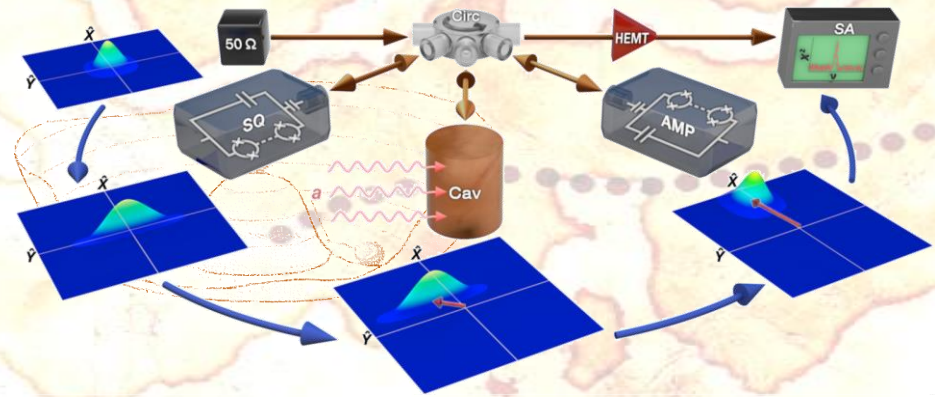
$$S.N.R. \propto P/T_{sys}$$

-> minimize noise

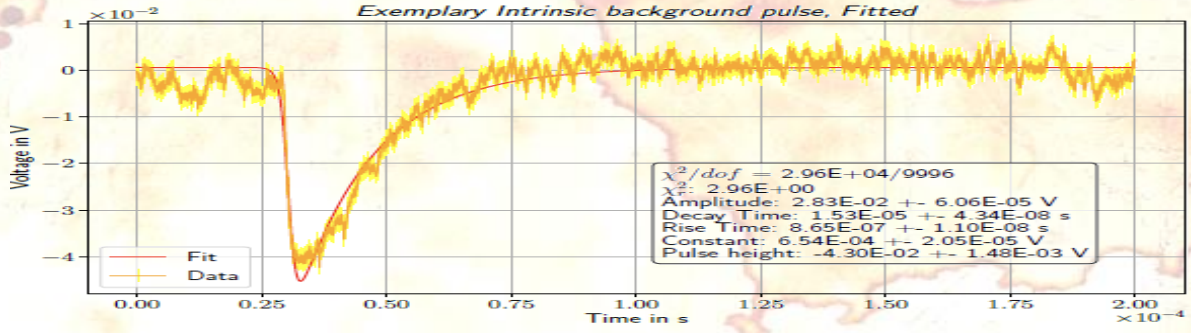
-> cryogenic temperatures



Quantum limited JPA
Haystac/ADMX

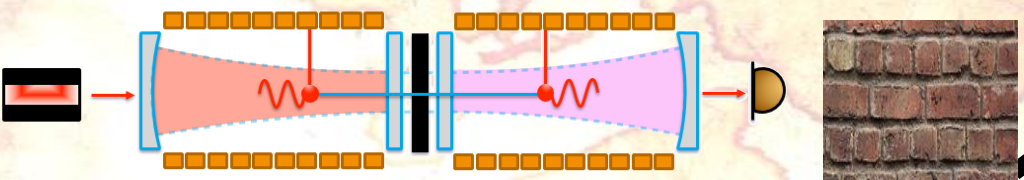


TWPA for MADMAX [2101.05815]



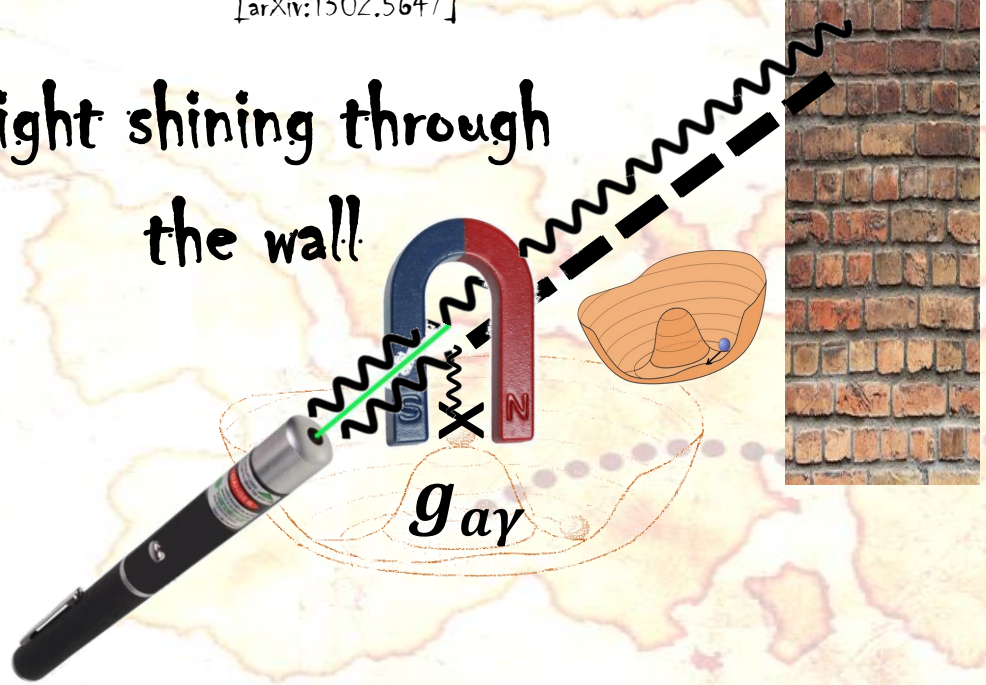
Single photon detectors
TES for ALPSII

The bounty hunters: Lab experiment

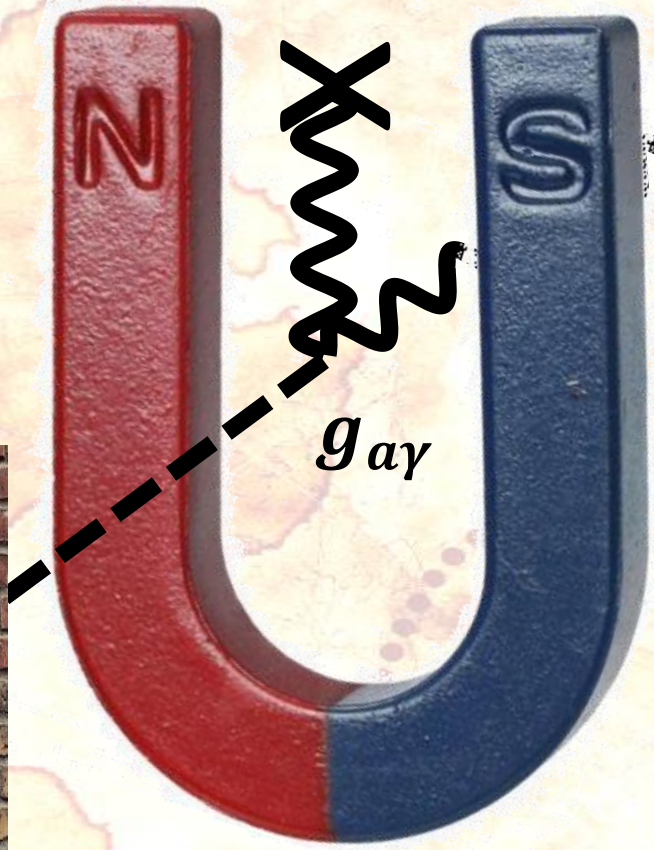


[arXiv:1302.5647]

Light shining through the wall.



Axions (and Axion Like Particles)



Laser as source



Light detection



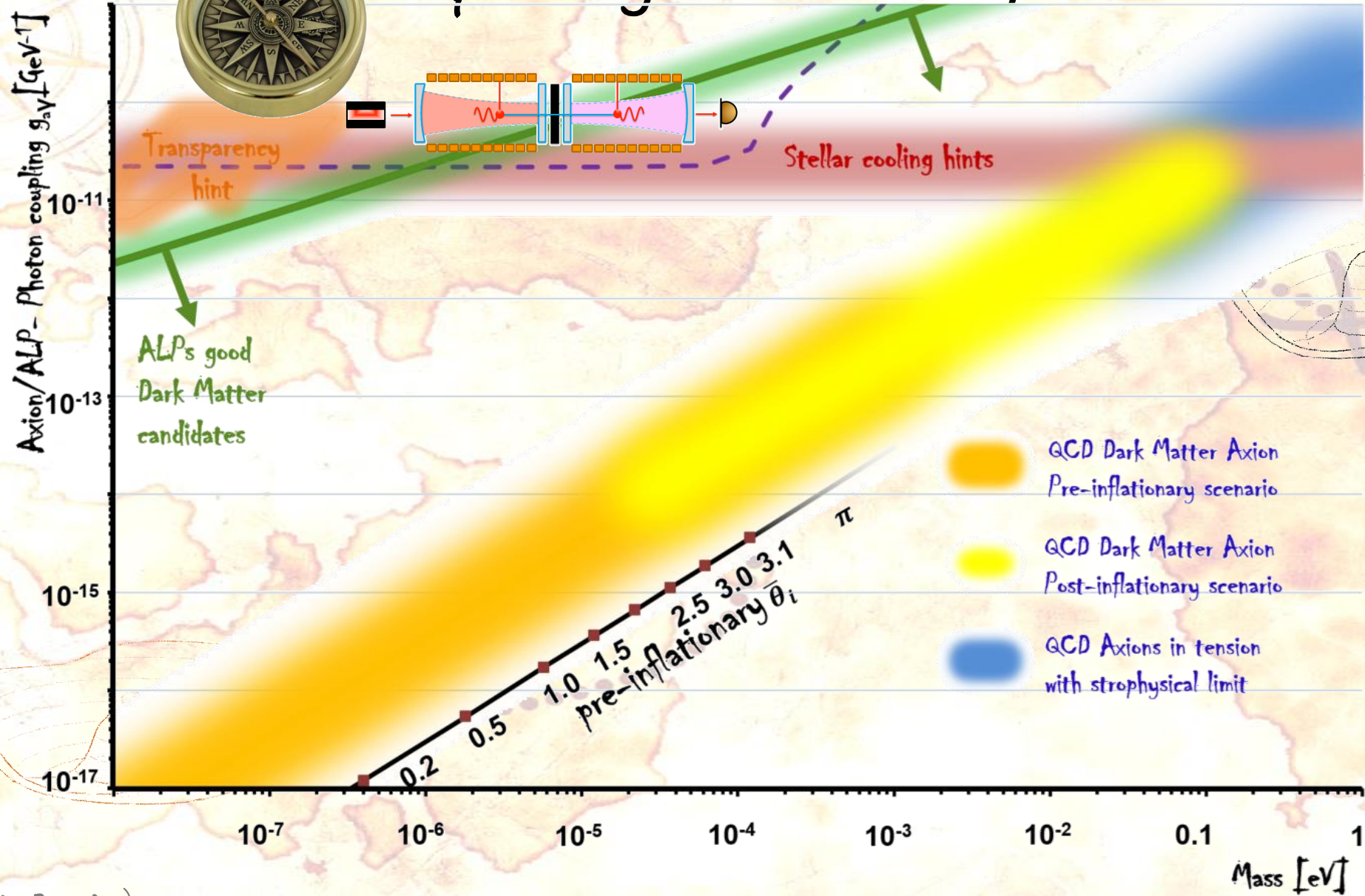
Image: DESY / Heiner Müller-Elsner



ALPSII at DESY

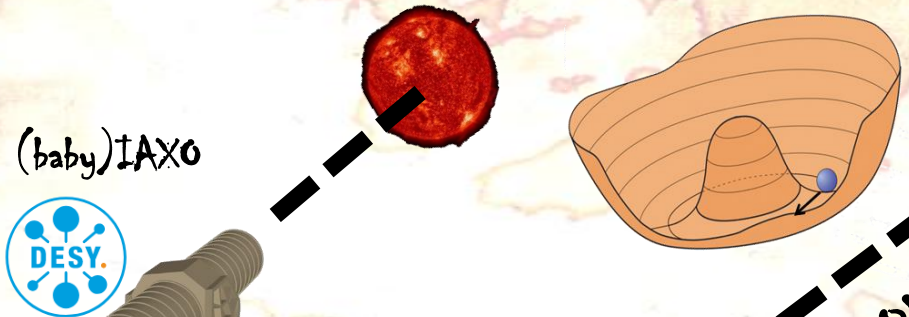
Exploring terra incognita

Axion/ALP - Photon coupling g_{γ} [GeV⁻¹]



The bounty hunters:

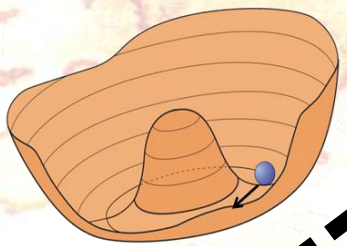
Helioscope $g_{\gamma\gamma}$ g_{ae}



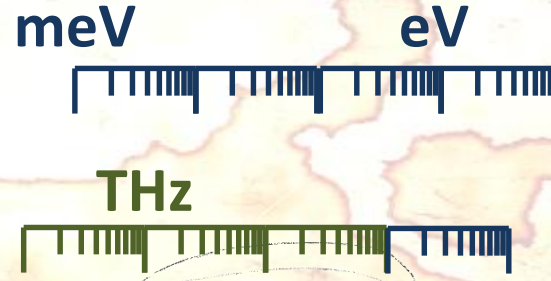
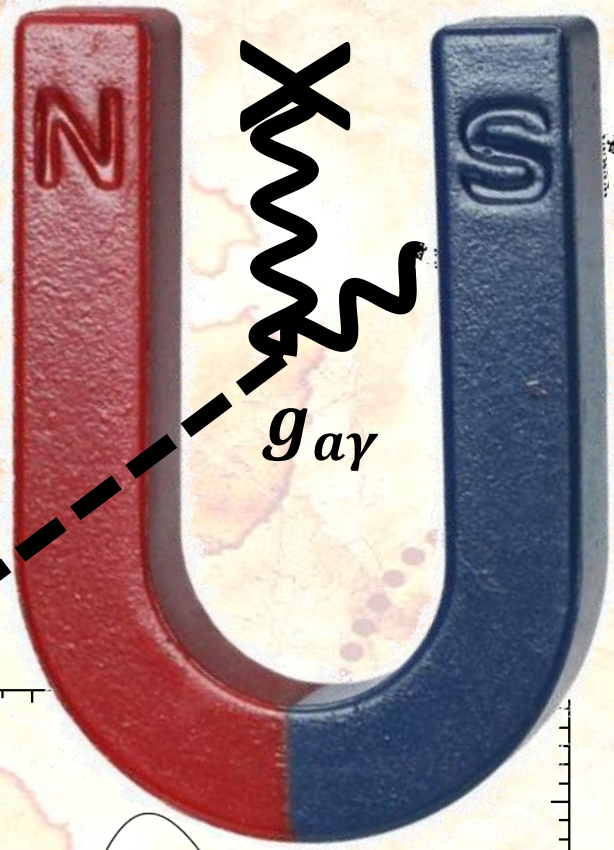
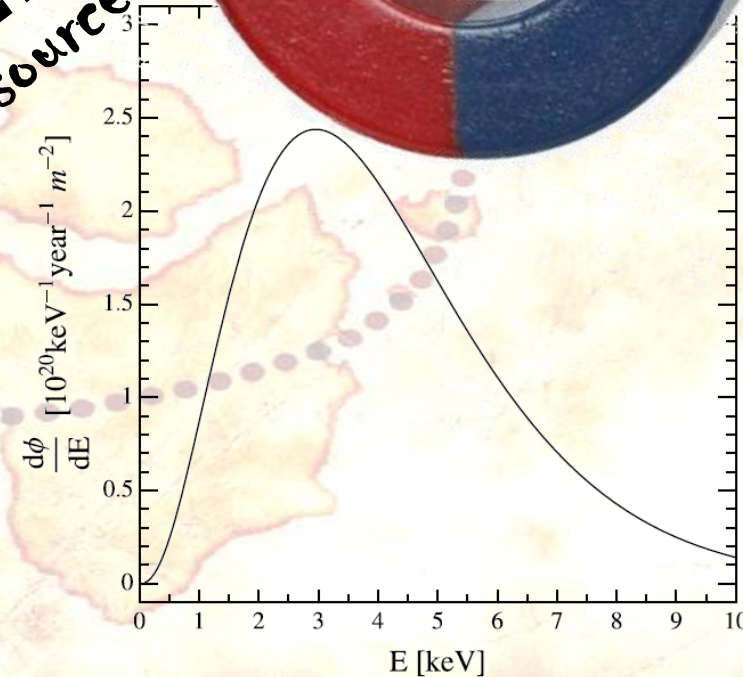
JCAP06(2019)047
[arXiv:1904.09155]



Axions (and Axion Like Particles)



Sun as source



$$h\nu_a = m_a c^2$$

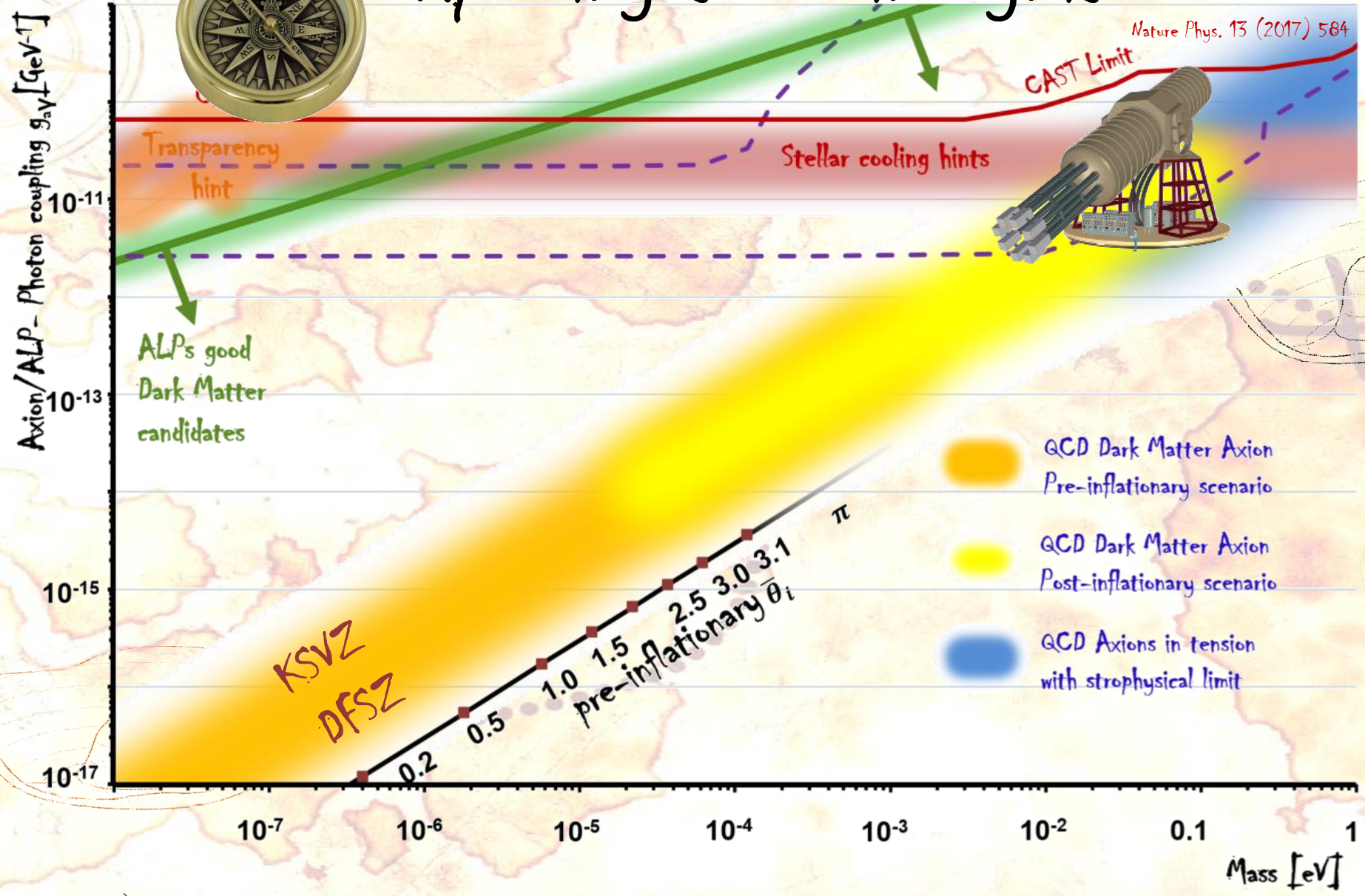
X-ray detection



CAST at CERN

Nature Phys. 13 (2017) 5

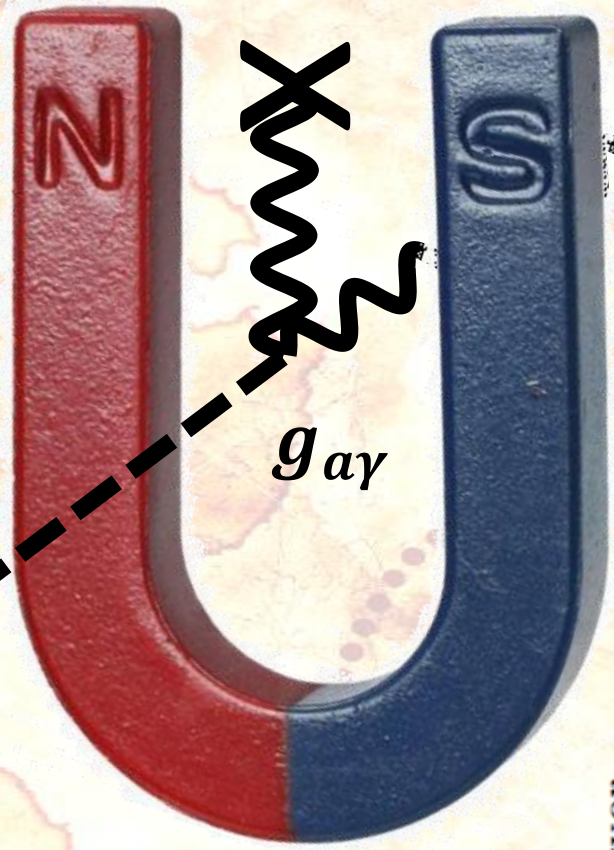
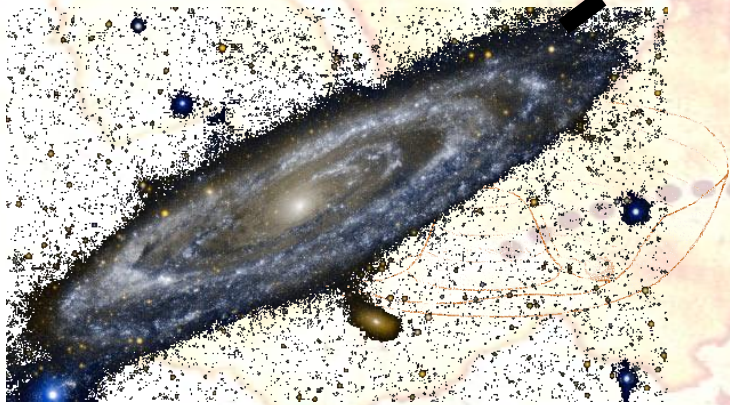
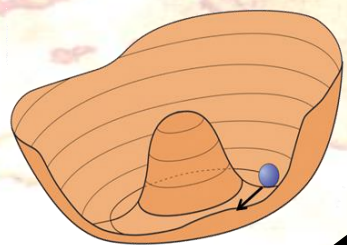
Exploring terra incognita



The bounty hunters:

Haloscope

$g_{a\gamma}$



$g_{a\gamma}$

γ

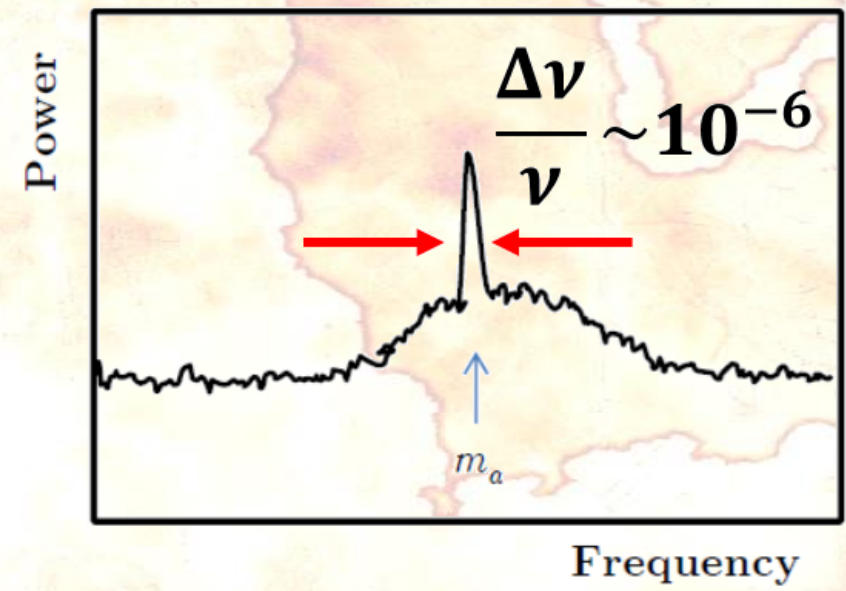


$$h\nu_a = m_a c^2$$

→ RF detection

Galactic DM as source

$$\langle v_{DM} \rangle = 10^{-3} c$$

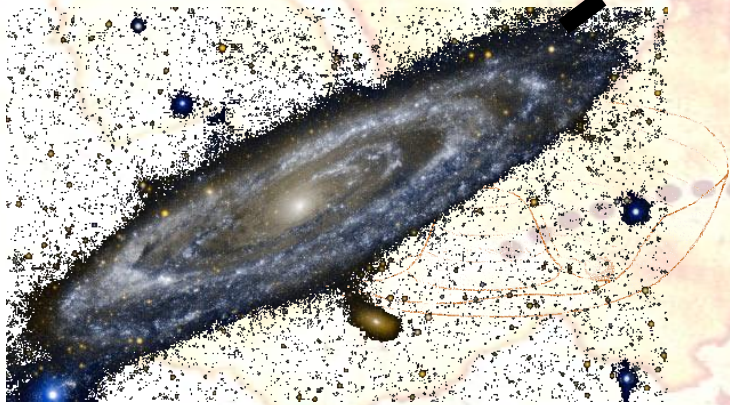


The bounty hunters:

Haloscope

$g_{a\gamma}$

- Strong magnet
- Boosting E-field
- Ultra low noise amplifier
- Cryogenic temperatures



Galactic DM as source

$$\langle v_{DM} \rangle = 10^{-3} c$$

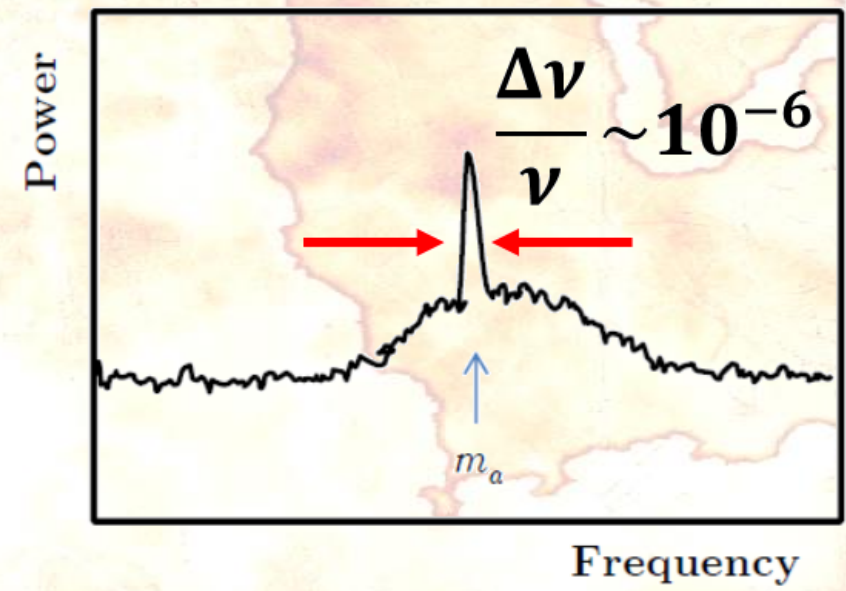


$g_{a\gamma}$



$$h\nu_a = m_a c^2$$

→ RF detection



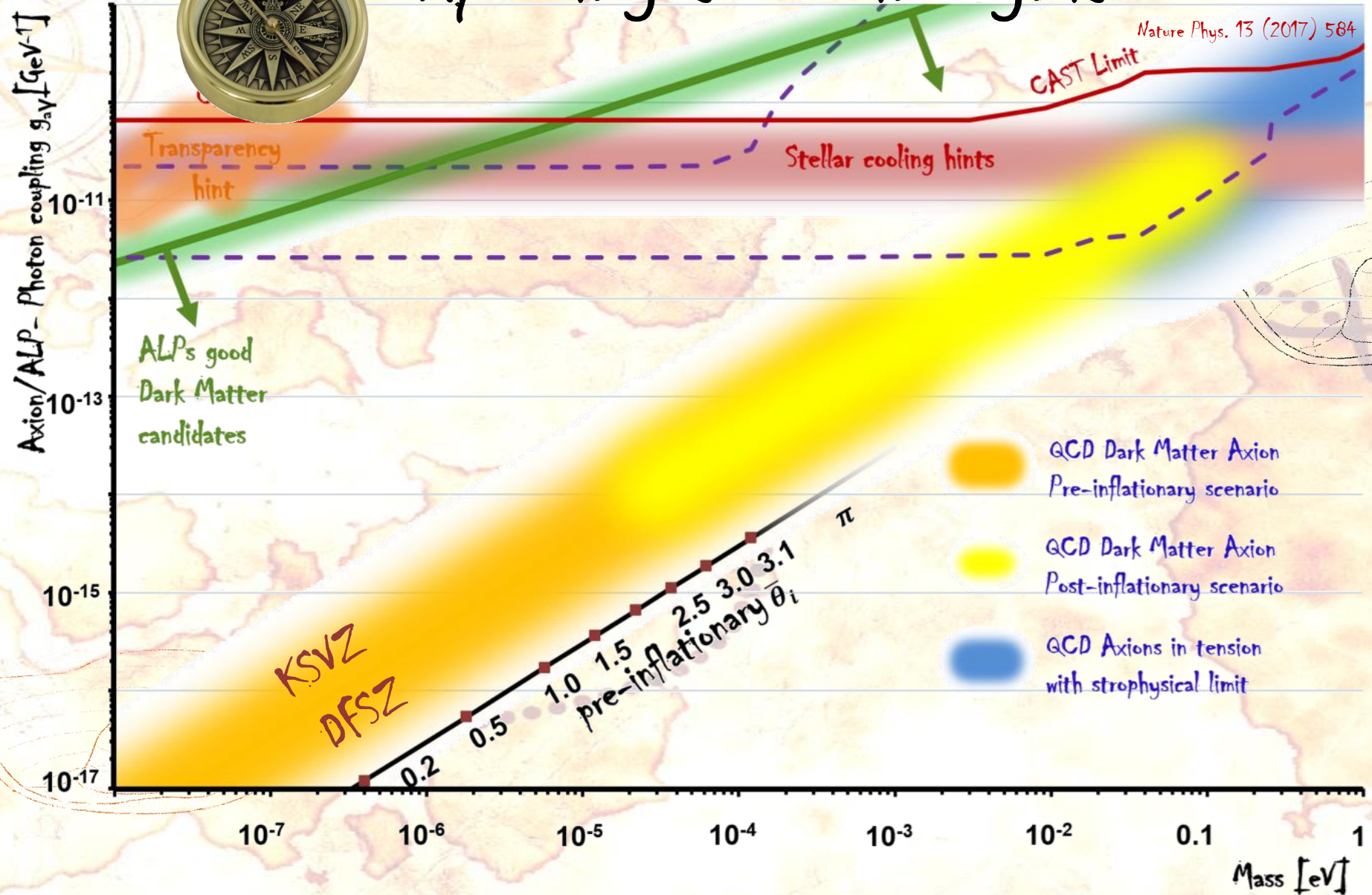
ADMX@University of Washington, USA

UW Seattle, LLNL, Fermilab, PNNL, NRAO, Uni Sheffield,
Uni Chicago, Uni Berkeley, Uni Florida, NIST, Uni Western Australia

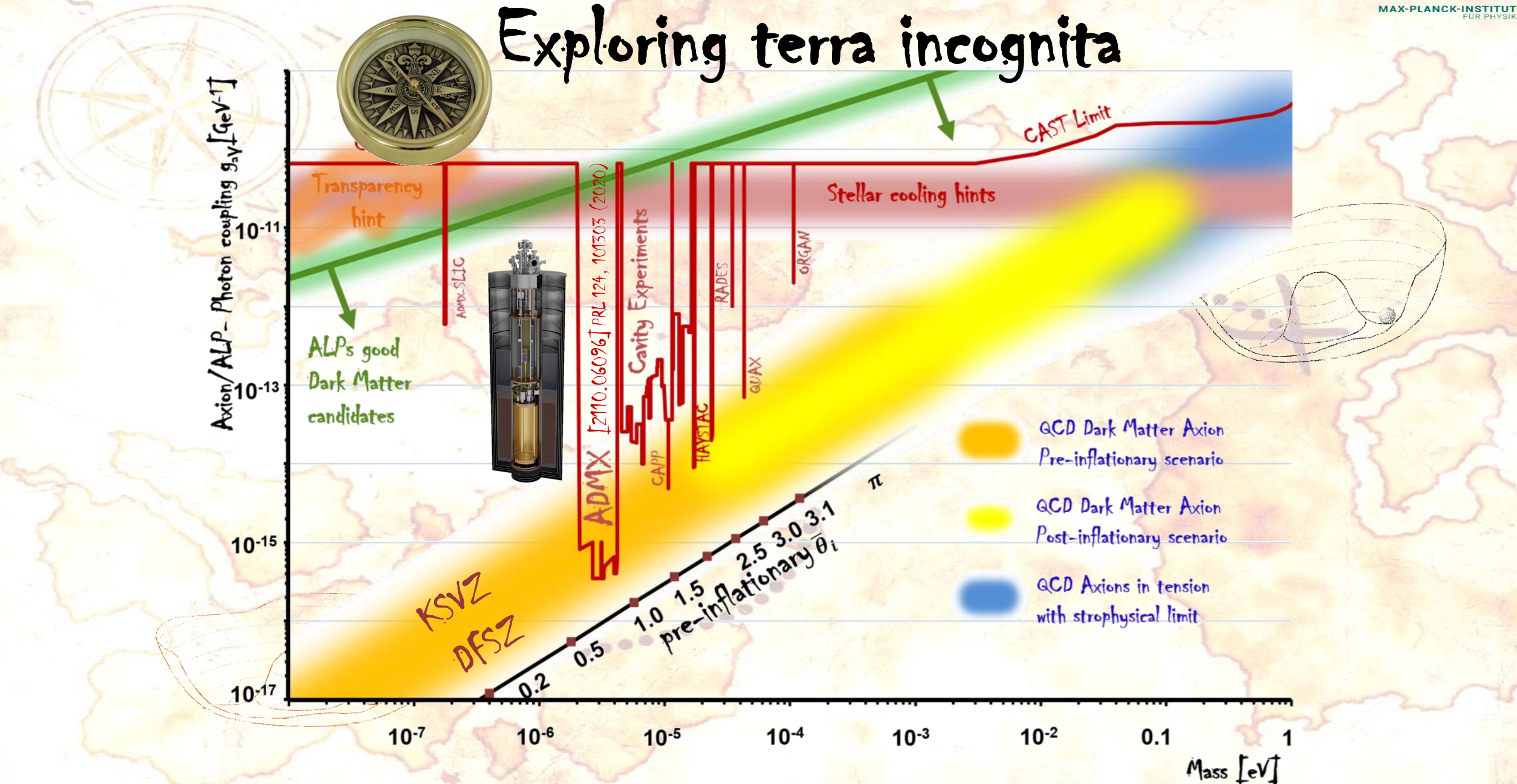


Axions (and Axion Like Particles) Cavity & multistrip SQUID amplifier @ sub K \rightarrow Sensitivity at quantum limit

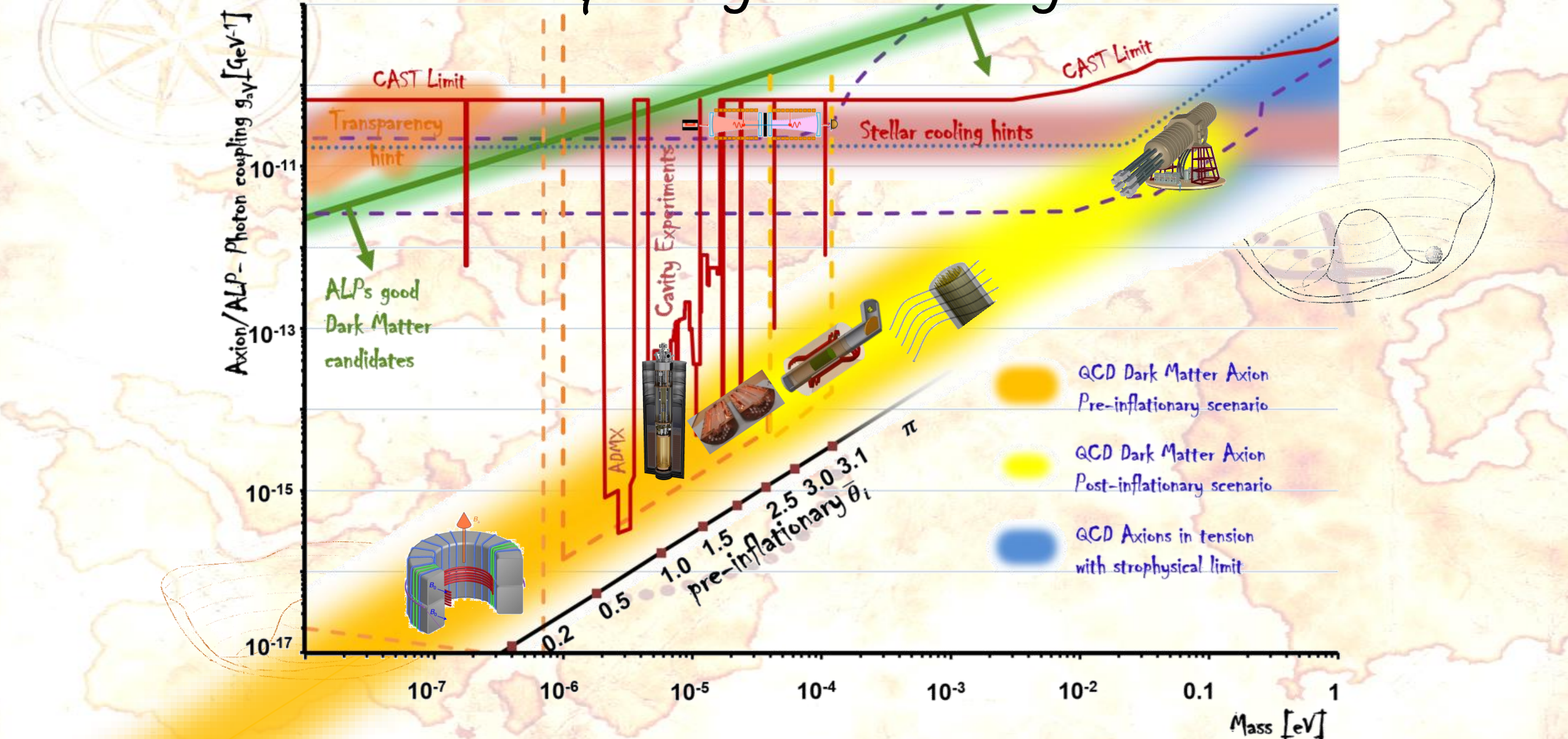
Exploring terra incognita



Exploring terra incognita

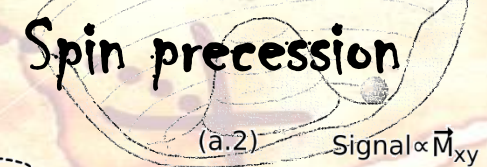
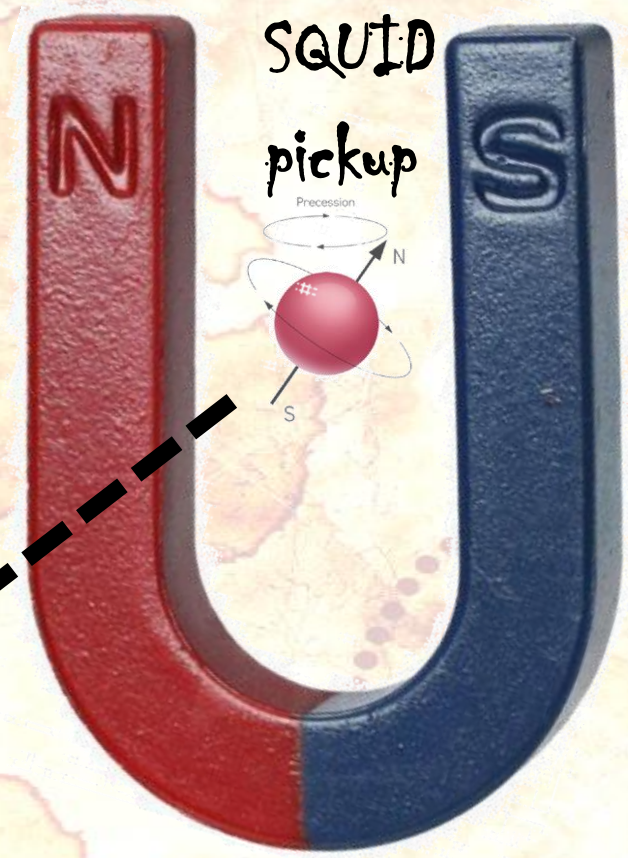
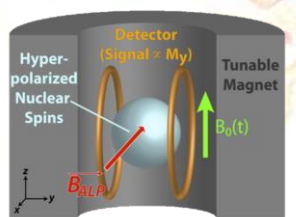


Exploring terra incognita

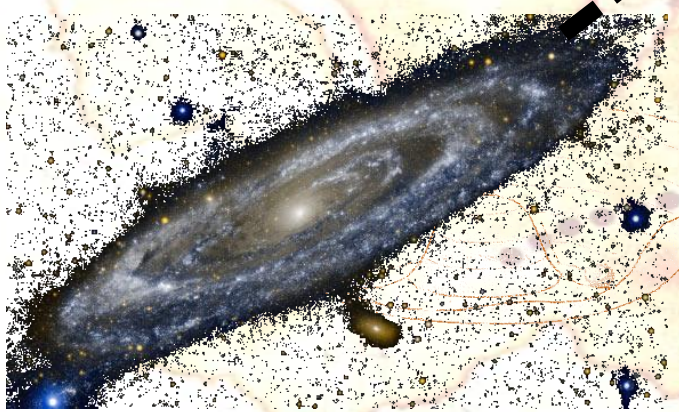


The bounty hunters: Haloscope

g_{aEDM} g_{aNN}
 Axion field oscillation
 → oscillating nEDM

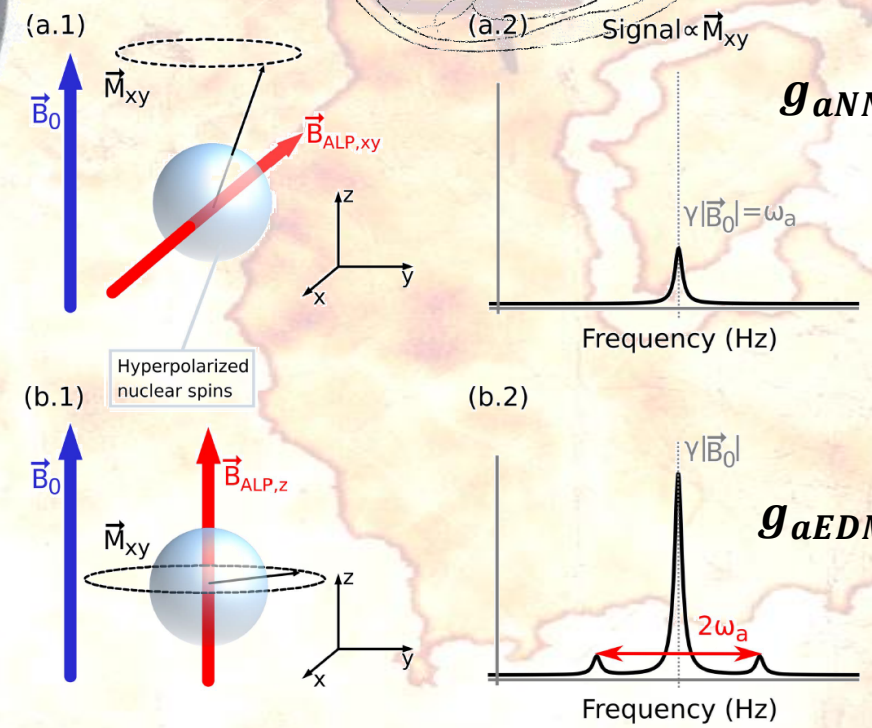


NMR techniques
 PhysRevLett.122.191302

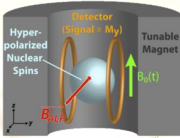
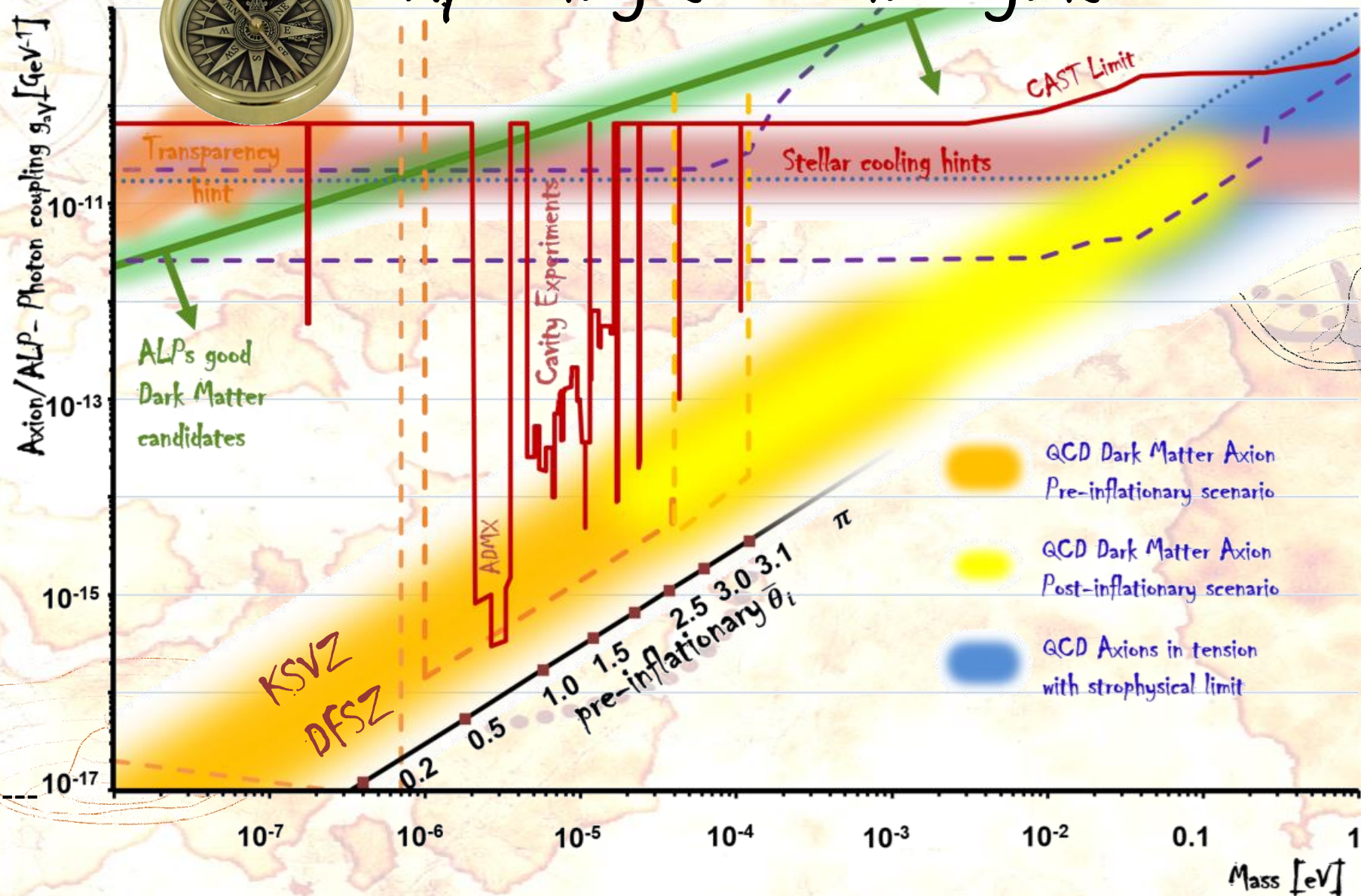


Galactic DM as source

$$\langle v_{DM} \rangle = 10^{-3} c$$

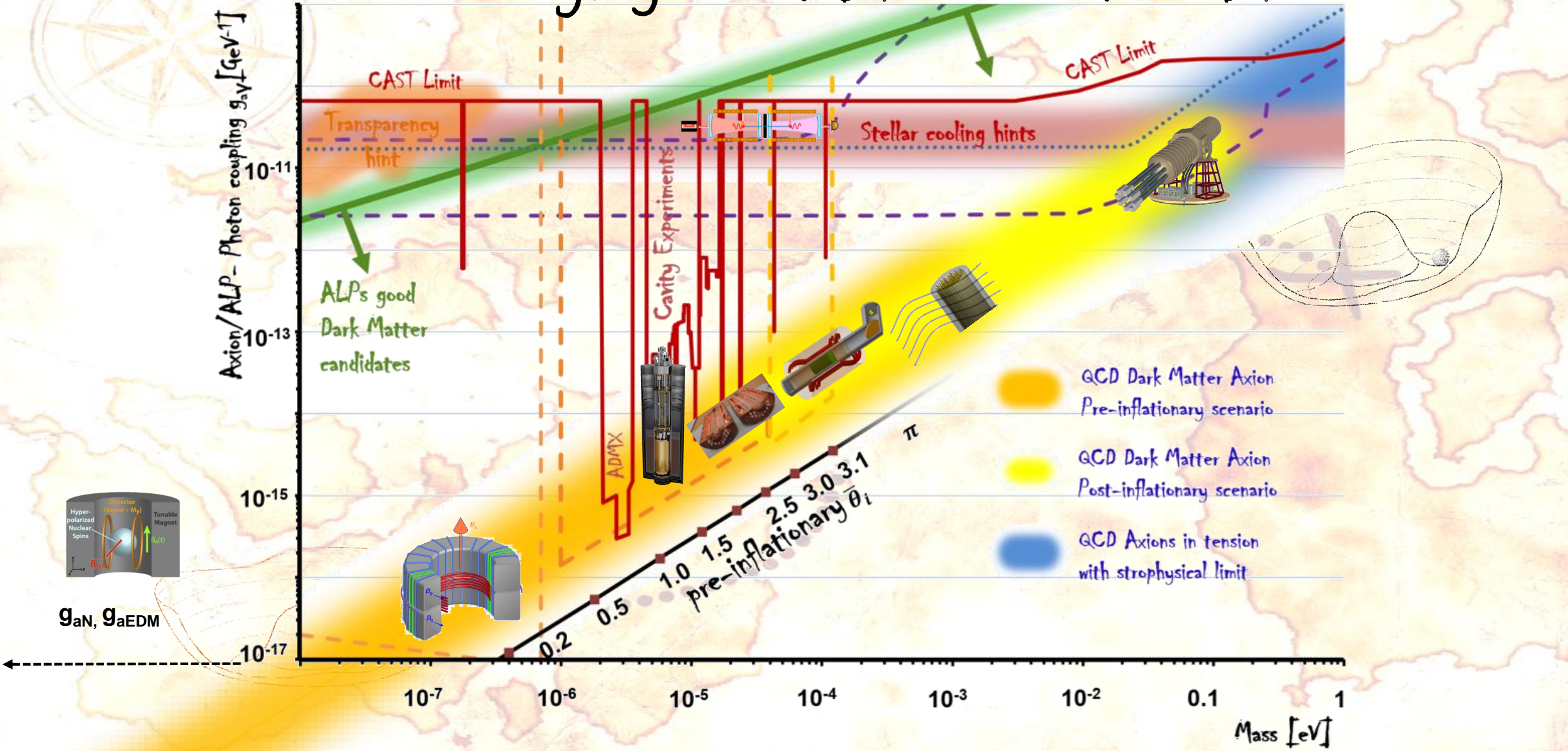


Exploring terra incognita



g_{aN}, g_{aEDM}

Surveying TERRA INCOGNITA

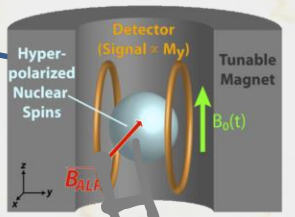


Status of axion mass survey

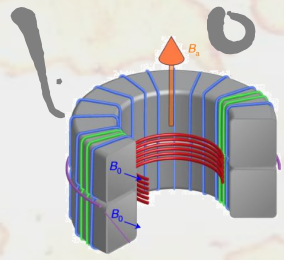
Exciting last 5 - 10 years:
plethora of approaches emerging

VERY COMPLEMENTARY!

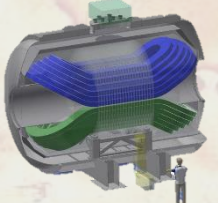
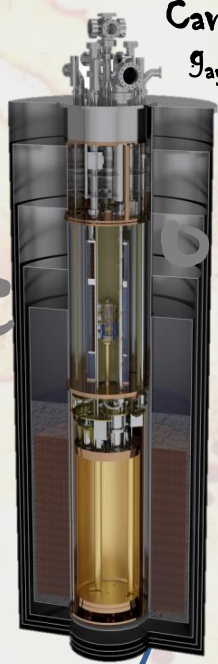
NMR / Spin-
precession
 g_{aN}, g_{aEDM}



LC circuit



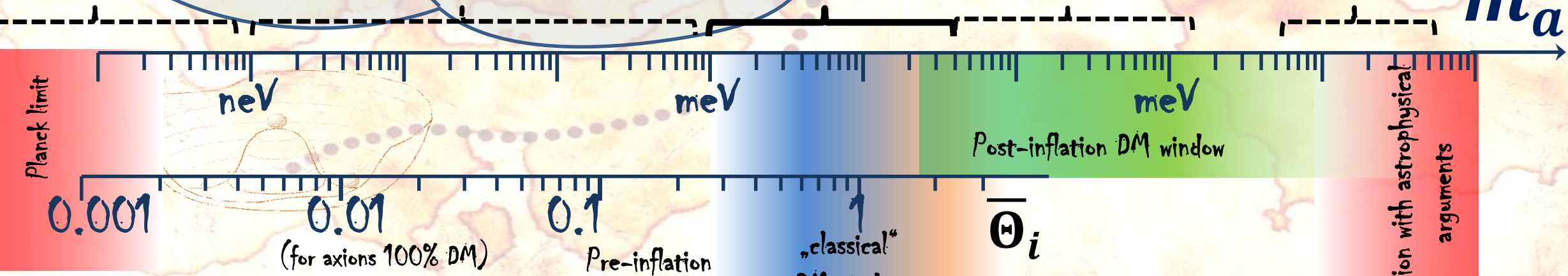
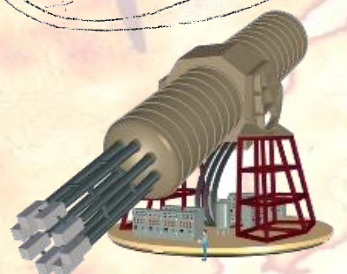
Cavity
 g_{ay}



Di-electric
haloscope



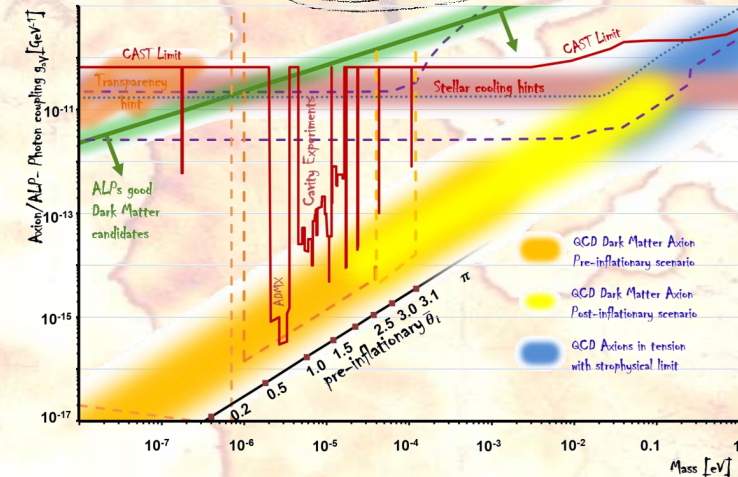
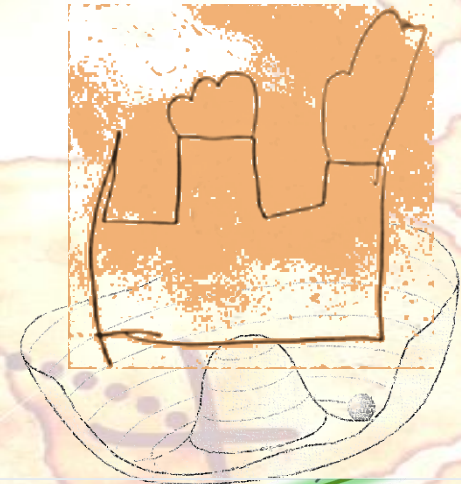
g_{ay}
Meta materials



Axions (and Axion Like Particles)

CONCLUSIONS:

- axion (& ALPs) very well motivated particle candidates
- Theory models give guidance: vast range to explore
- ADMX sensitive to QCD dark matter axion
- Last years: promising new approaches
- Hopefully not too distant future:



solve strong CP problem & find dark matter axion!

Thanks for your attention!



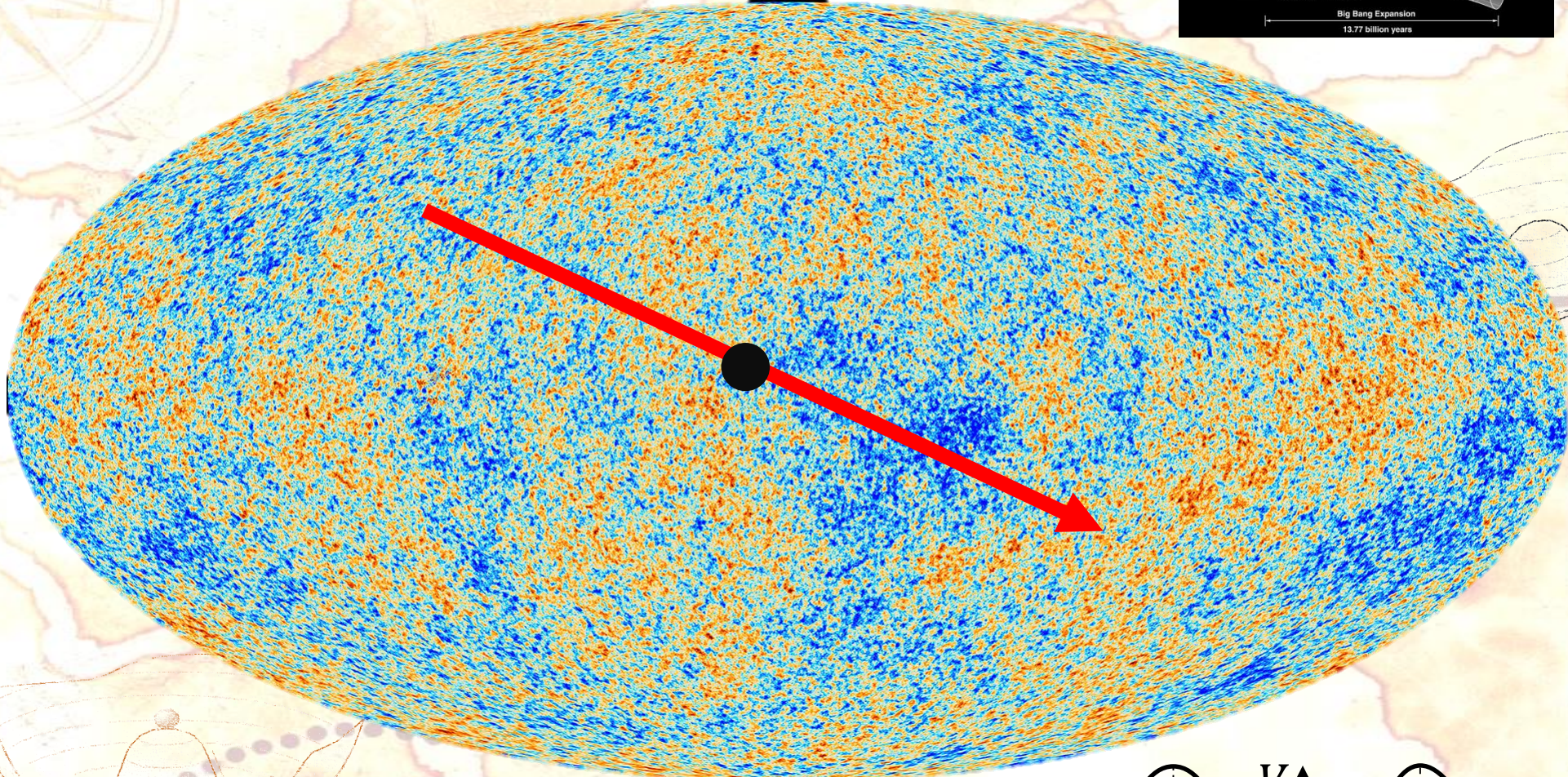
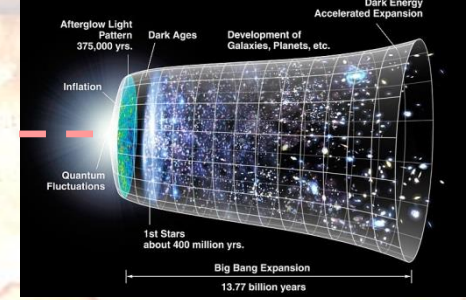
R. Peccei und H. Quinn,
Phys. Rev. Lett. **38**, 1440 (1977)
S. Weinberg, Phys. Rev. Lett. **40**, 223 (1978);
F. Wilczek, Phys. Rev. Lett. **40**, 279 (1978)

The Birth of Axions

Frank Wilczek
Institute for Advanced Study
Princeton, NJ 08540

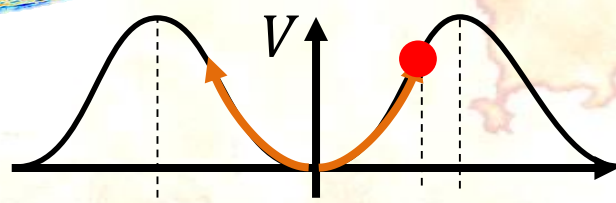
usual, very light particle. I called this particle the *axion*, after the laundry detergent, because that was a nice catchy name that sounded like a particle and because this particular particle solved a problem involving *axial* currents.

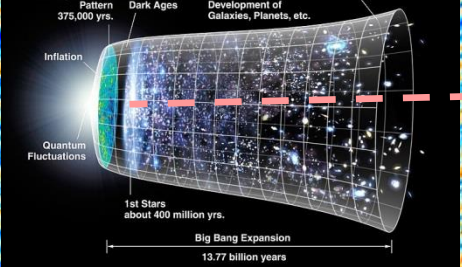
Pre-inflationary scenario



One value of $\bar{\theta}_i$ in entire visible universe

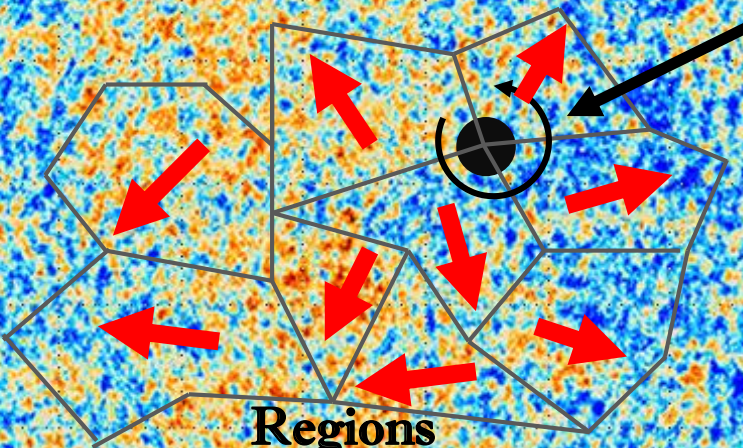
$$0 < |\bar{\theta}_i| < \pi$$





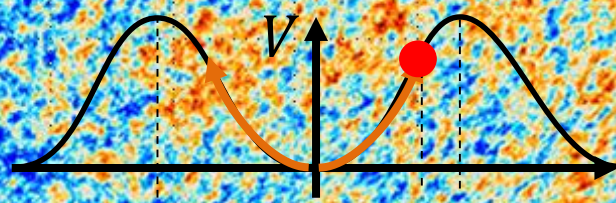
Post-inflationary Scenario

Complications by decay of topological defect



Regions much smaller in reality


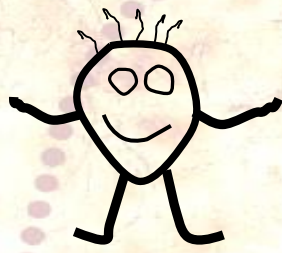
Average of all possible $\bar{\theta}_i$
→ Prediction for overall density



The size of DM Axions



λ de Broglie
 $\langle v_{DM} \rangle \sim 10^{-3} c$


Predictions:

$|\Theta_i|$ arbitrary $0 - \pi$



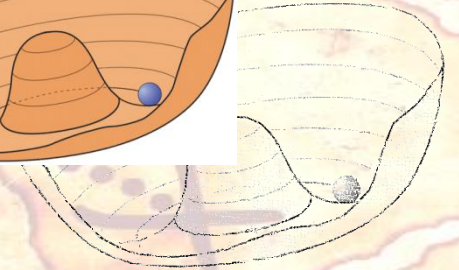
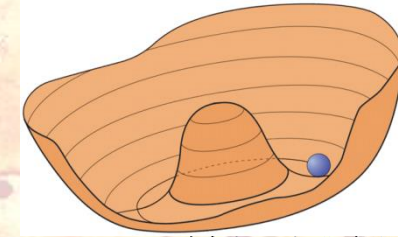
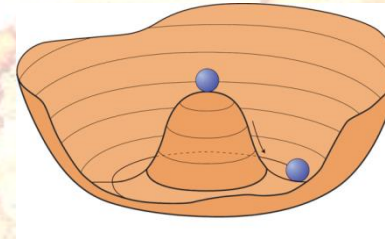
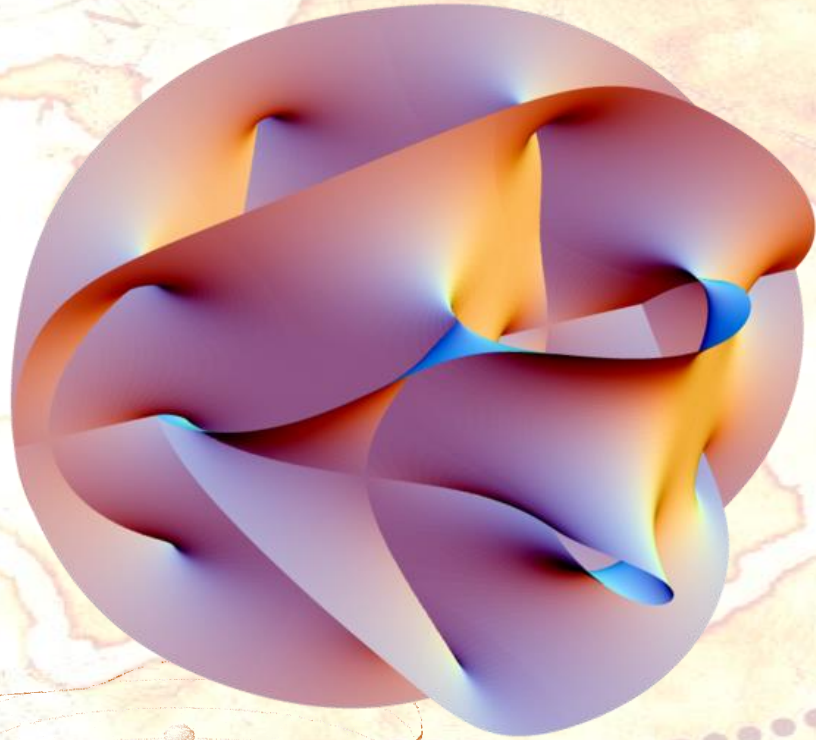
Pre-
Post-
inflationary scenario

$|\Theta_i| \emptyset$ of all
possible values




DM axions fit into experiment!

ALPs emerging from string compactification: the Axiverse



No direct relation btw.
 m_{ALP} and f_{ALP}

Some astrophysical inconsistencies:

- Transparency hint
- Cooling anomalies

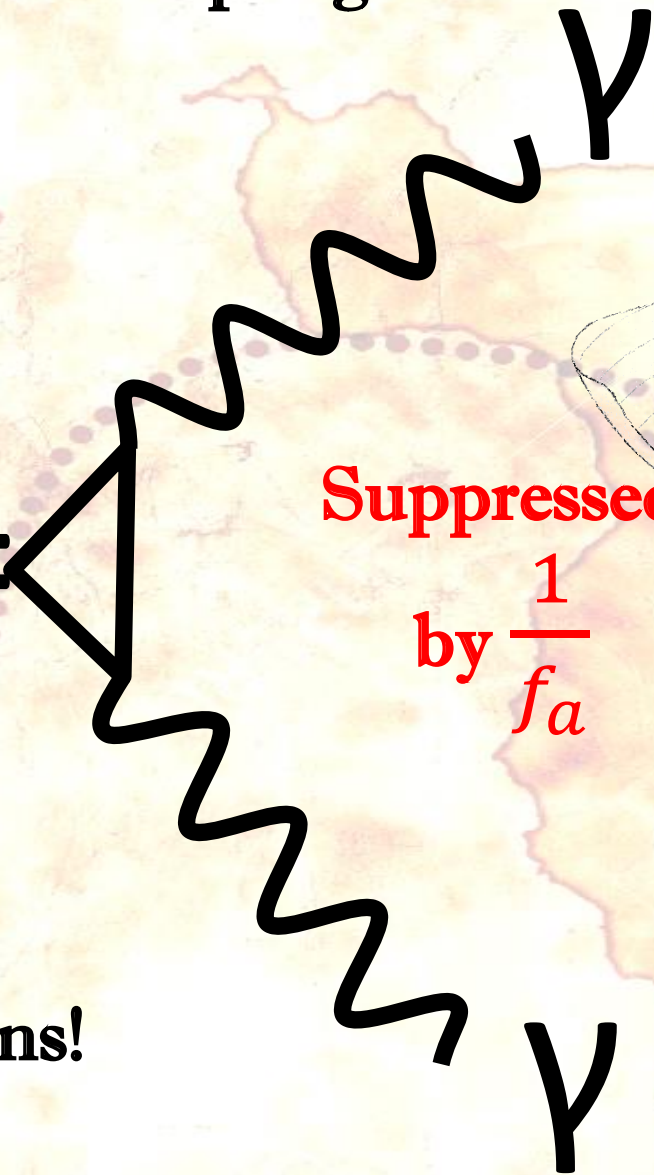
Could be explained by ALPs

Axion (ALP) - Photon Coupling:

The Axion (ALP) carries same quantum numbers as η^0 and π^0



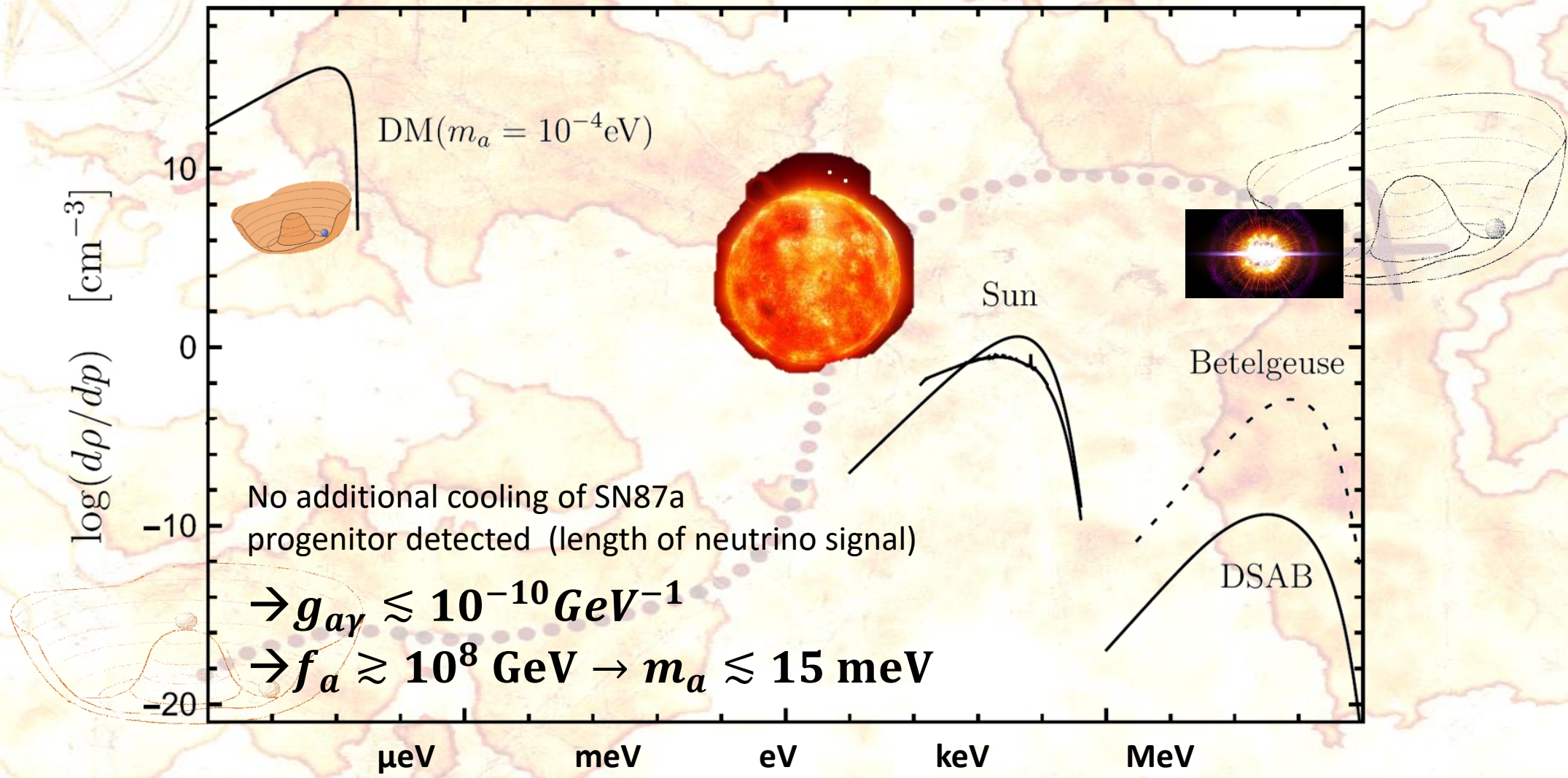
Suppressed
by $\frac{1}{f_a}$



→ Quantum mechanical mixing with π^0 & $\eta^0 \Rightarrow 2$ photons!

Axion - Sources:

I.G. Irastorza, J. Redondo / Progress in Particle and Nuclear Physics 102 (2018) 89-159

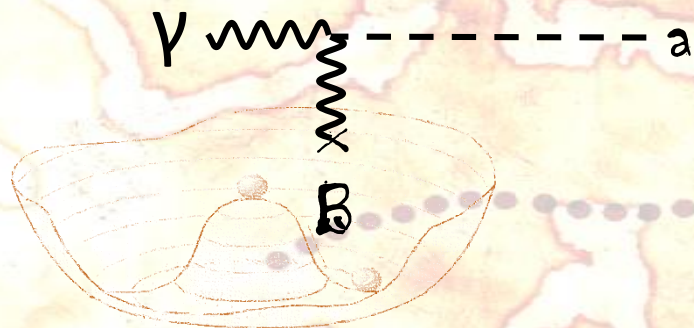
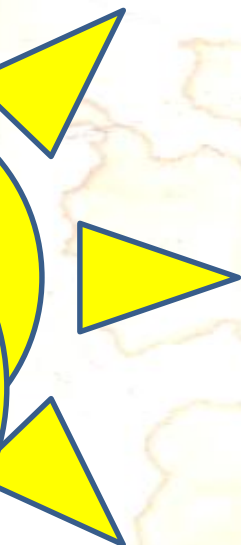
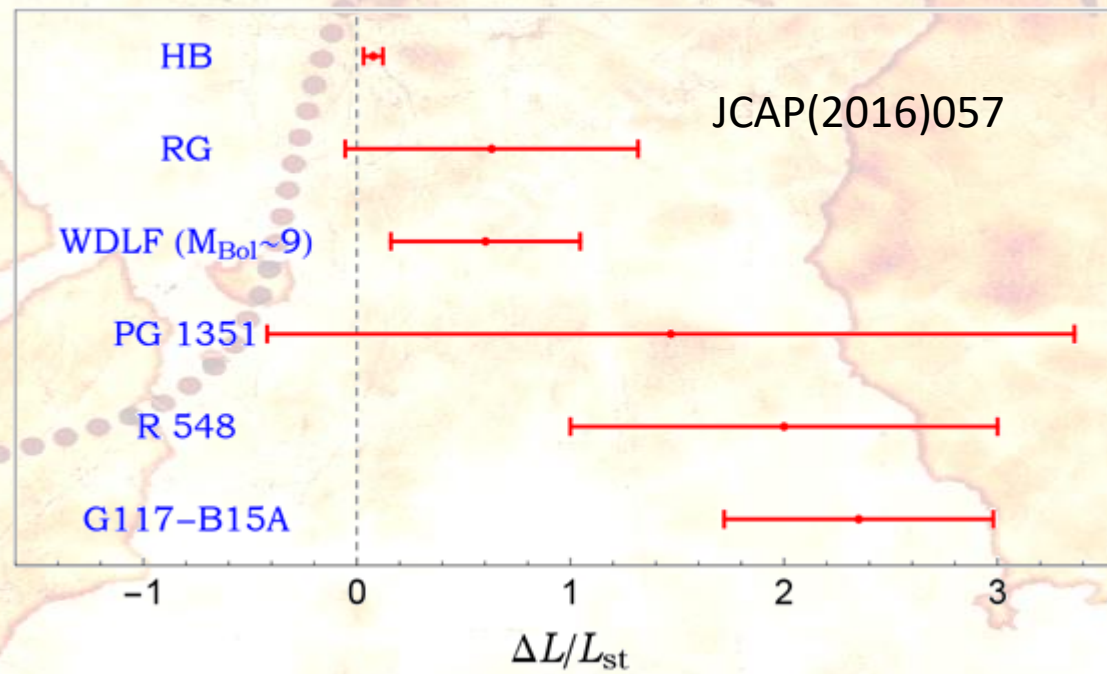


Evidence fo ALPs?

Transparency of intergalactic medium:



Anomalous cooling rate of white dwarfs and HB stars (?)



Cavities in B-Field:

Adjusting resonance frequency: “Tuning Rod”



ADMX

HAYSTAC

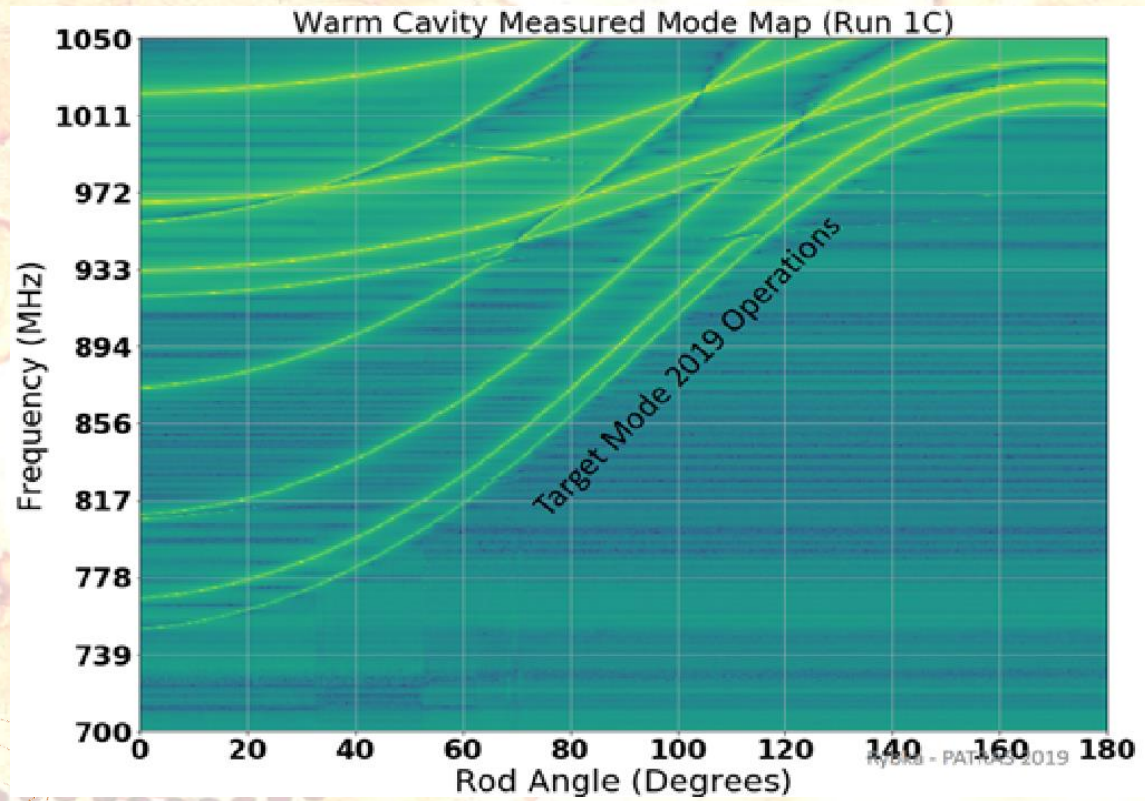
University of Washington, USA Yale University, USA

$$P_{sig} \propto B^2 V Q_{cav}$$

$$P_{sig}(B=6.8 \text{ T}, V=136 \text{ l}, Q=10^5) \sim 2 \cdot 10^{-22} \text{ W}$$

Cavities in B-Field:

Adjusting resonance frequency: “Tuning Rod”



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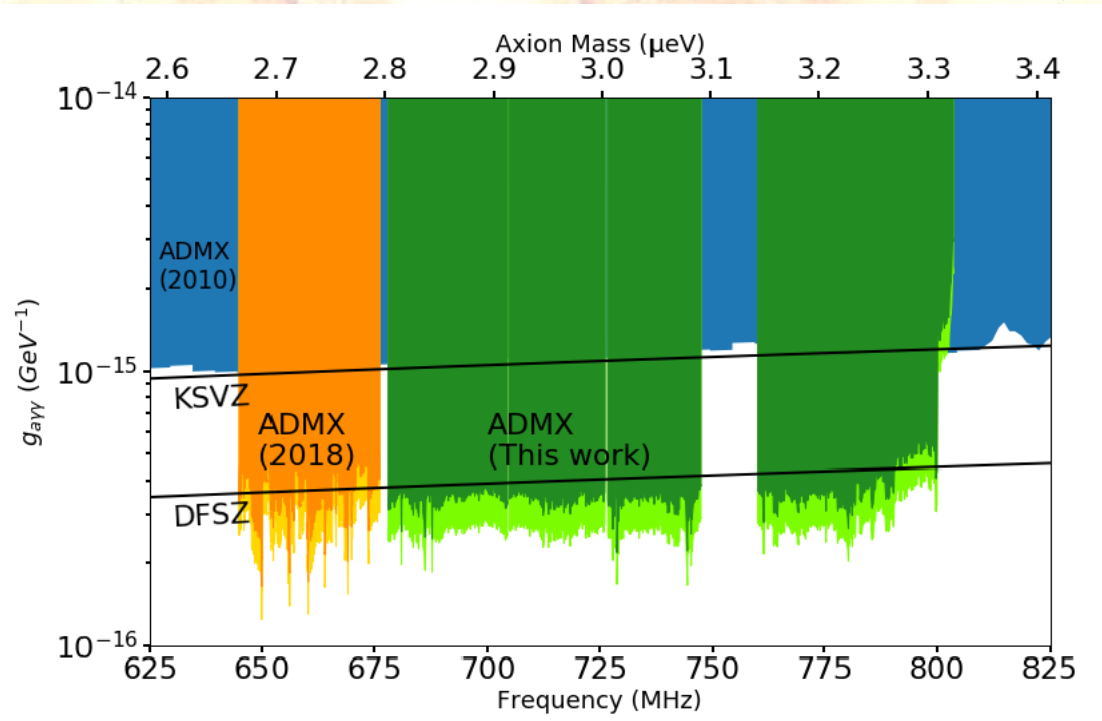
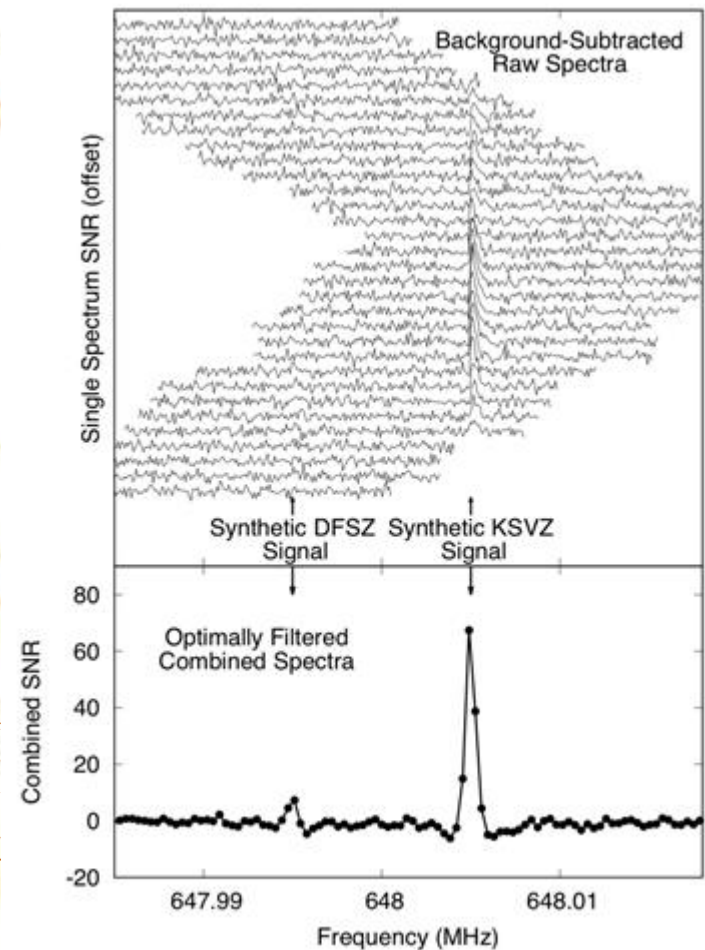
$$P_{sig}(B=6.8 T, V=136 l, Q=10^5) \sim 2 \cdot 10^{-22} W$$

ADMX@University of Washington, USA

Measurements ongoing!

Sensitive to DM axion masses $\sim 2\text{-}4\mu\text{eV}$

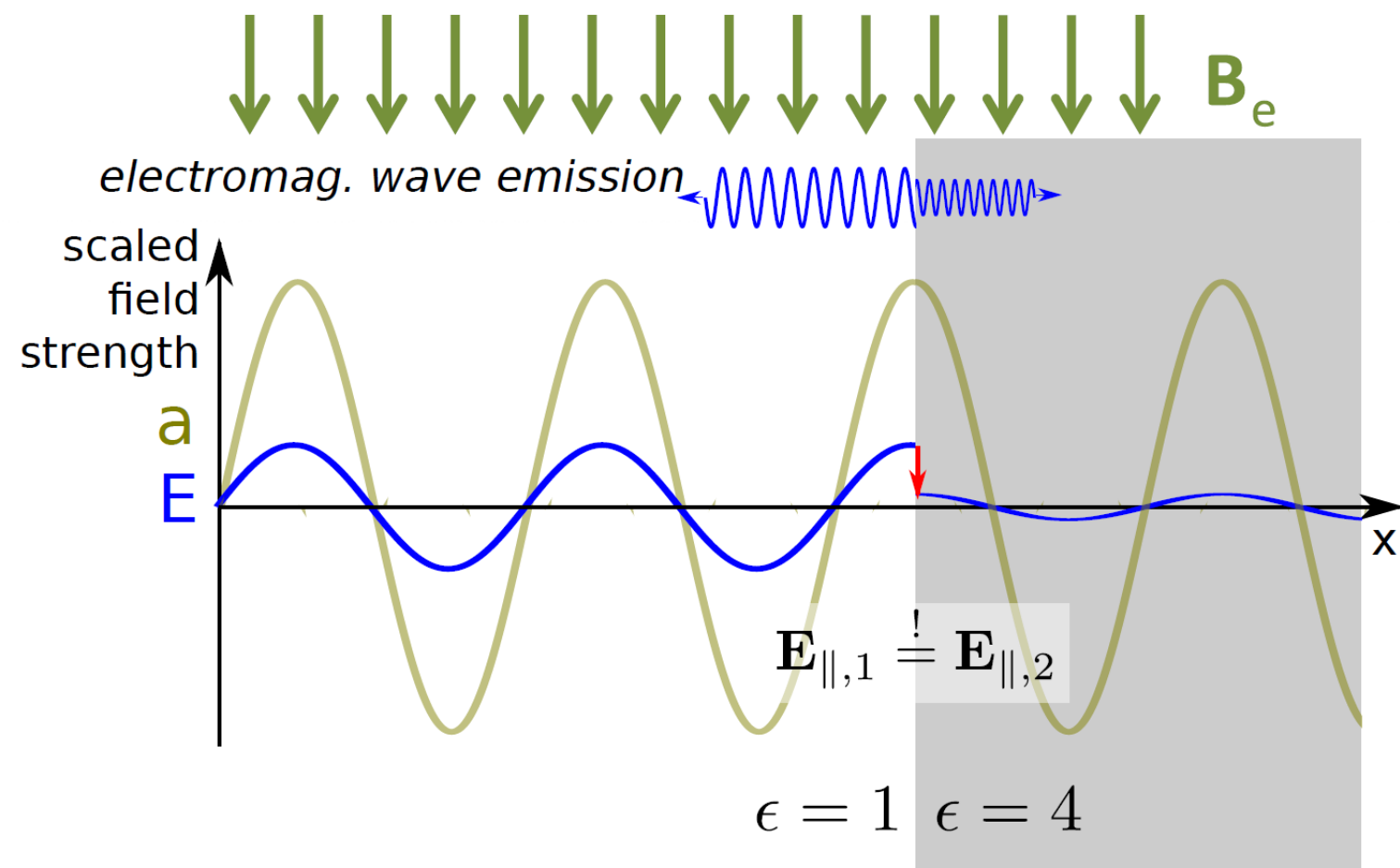
Potential up to $\sim 40\mu\text{eV}$



ADMX: Phys. Rev. Lett. 124, 101303 (2020)

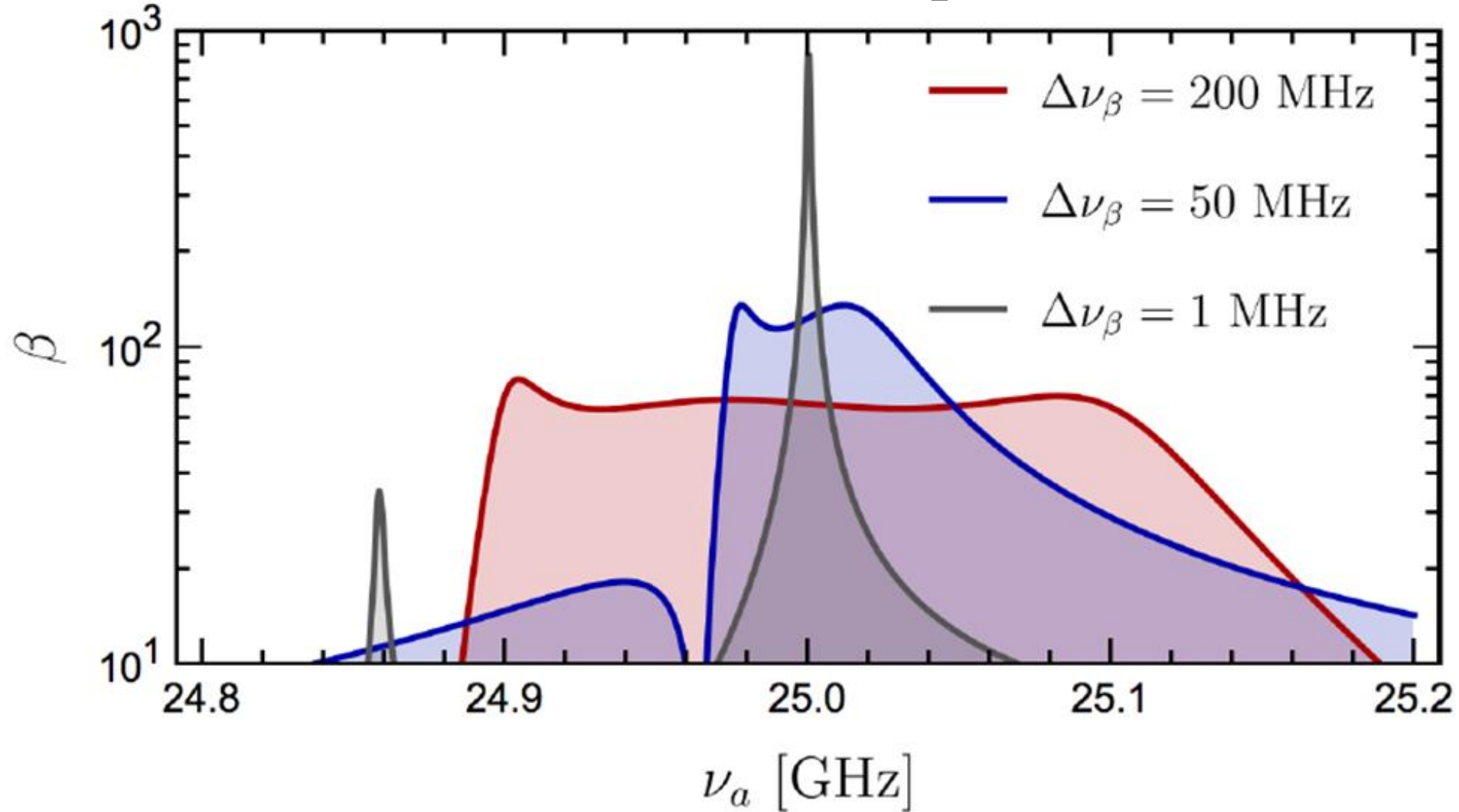
HAYSTAC: Phys. Rev. D 97, 092001 (2018)

Dielectric Haloscope



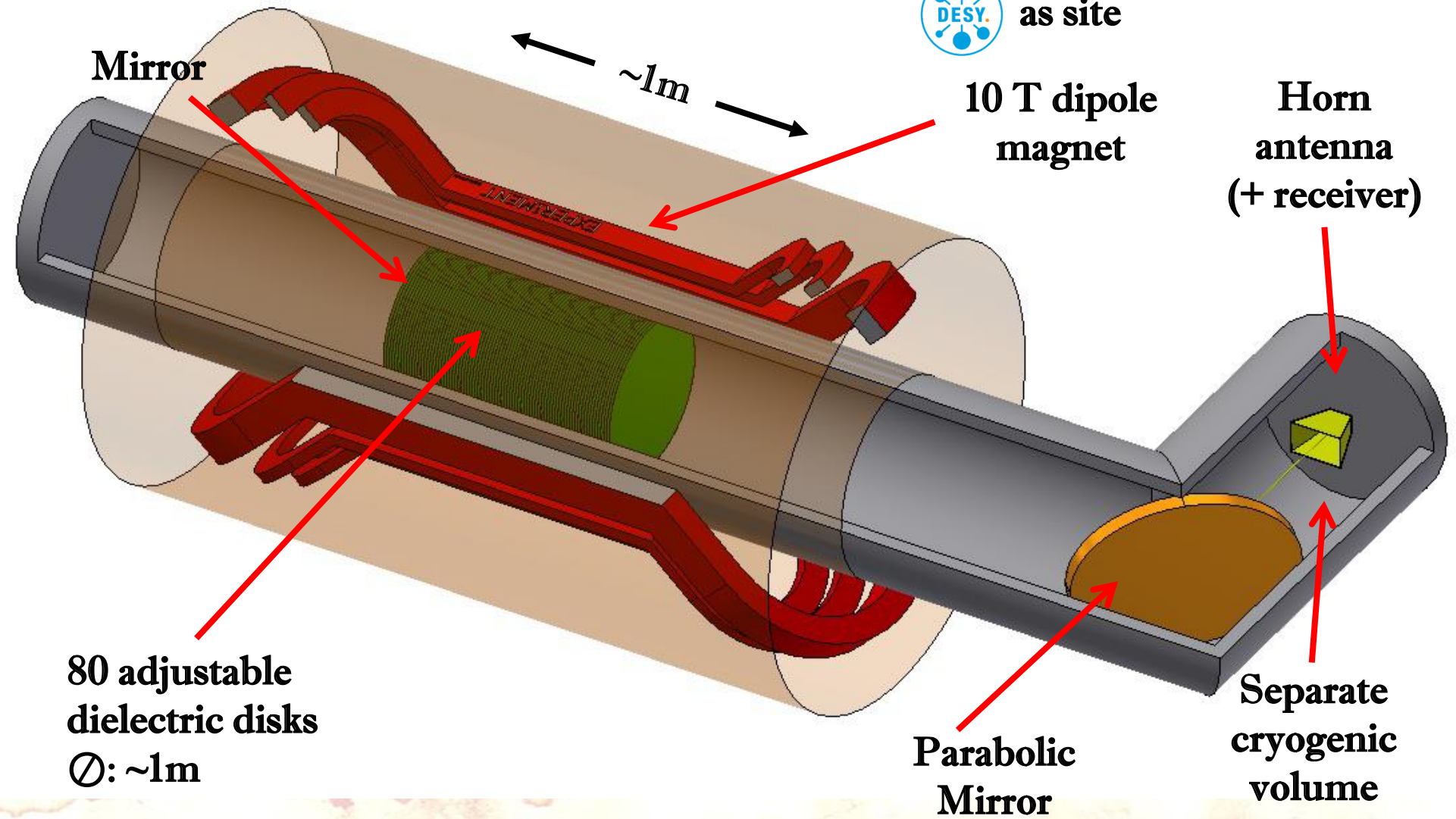
$$\left(\frac{P}{A}\right)_{\text{mirror}} \sim 2 \cdot 10^{-27} \frac{W}{m^2} \left(\frac{B_{\parallel}}{10 T}\right)^2 (g_{a\gamma\gamma} m_a)^2$$

Dielectric Haloscope





Magnetized disk and Mirror Axion eXperiment



80 adjustable dielectric disks
Ø: ~1m

Parabolic Mirror

Separate cryogenic volume

