



Universität Hamburg

DER FORSCHUNG | DER LEHRE | DER BILDUNG



# MADMAX

## Towards a Dielectric Axion Haloscope

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Universität Hamburg

On behalf of the MADMAX Collaboration

12<sup>th</sup> December 2022

2<sup>nd</sup> DMLab Meeting

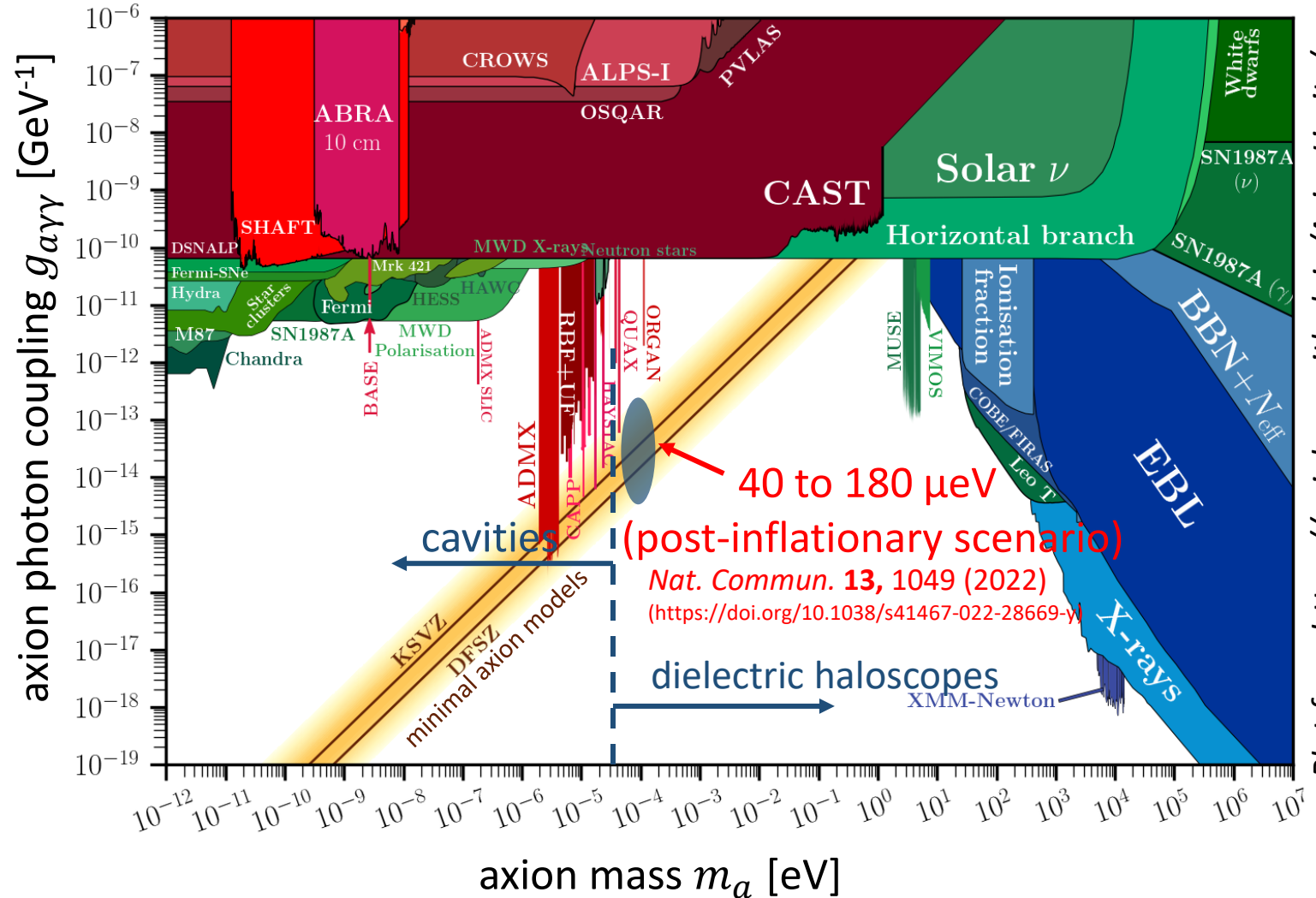
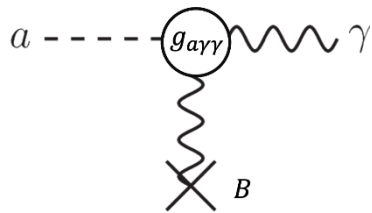
DESY Hamburg

# Axion Parameter Space



## The Axion:

- Pseudo Nambu-Goldstone boson
- Small mass and small couplings
- Connected to solution of the strong CP problem
- Primakoff/Sikivie effect: Photon-Axion conversion in strong EM fields
- **Axion** can explain (part of) **Cold Dark Matter**

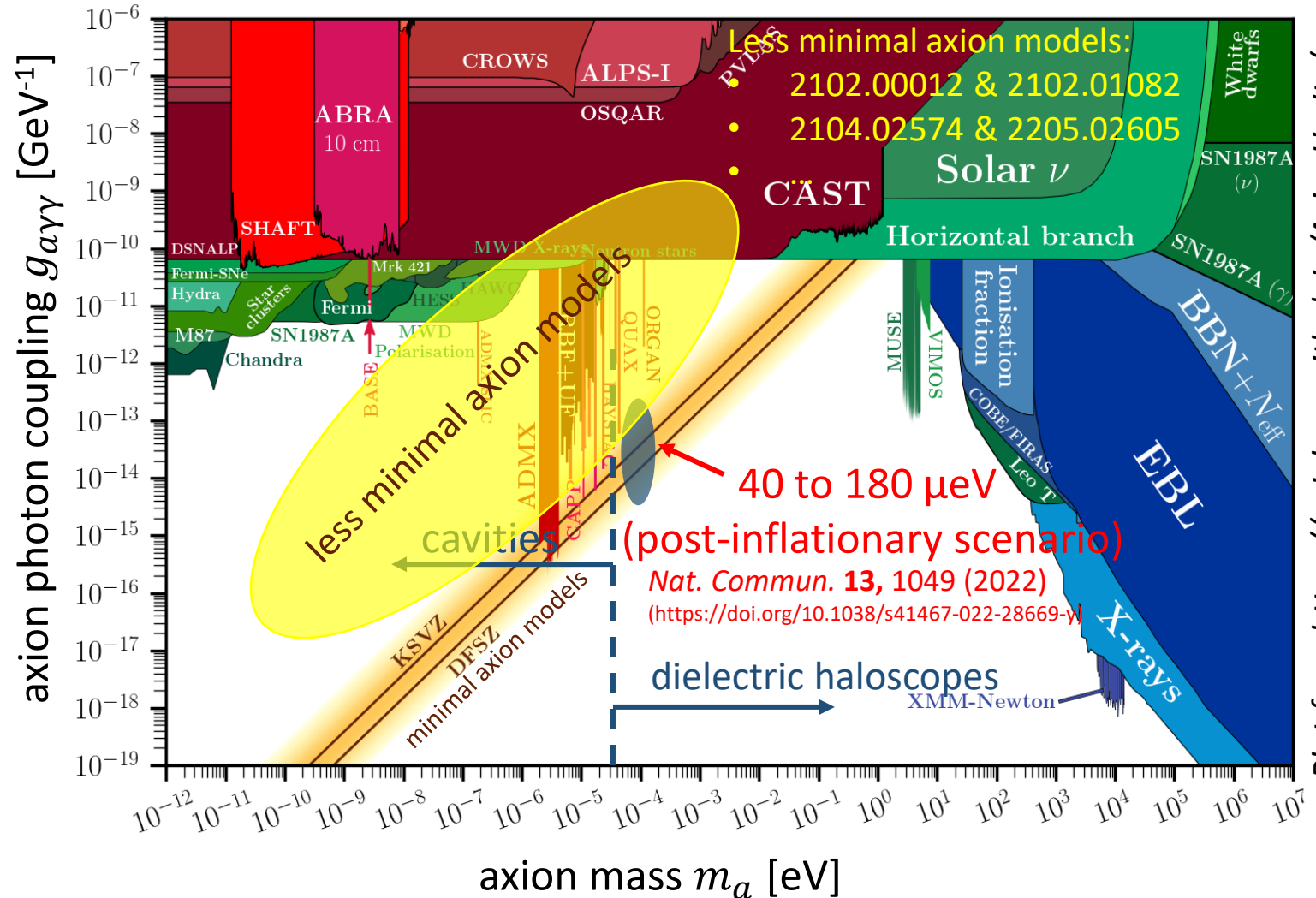
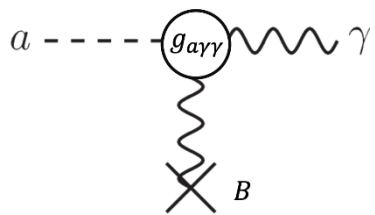


# Axion Parameter Space



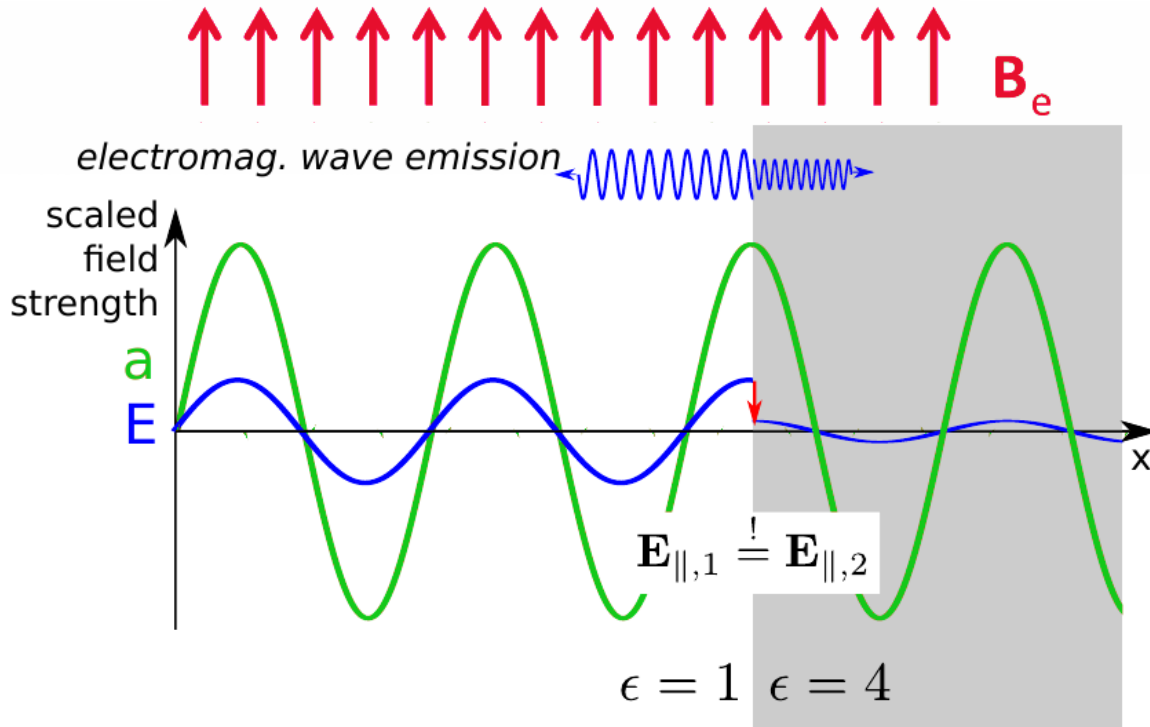
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Plot from <https://cajohare.github.io/AxionLimits/>

# Dielectric Haloscope



In an external magnetic field  $B_e$  the axion field  $a(t)$  sources an oscillating electric field  $E_a$

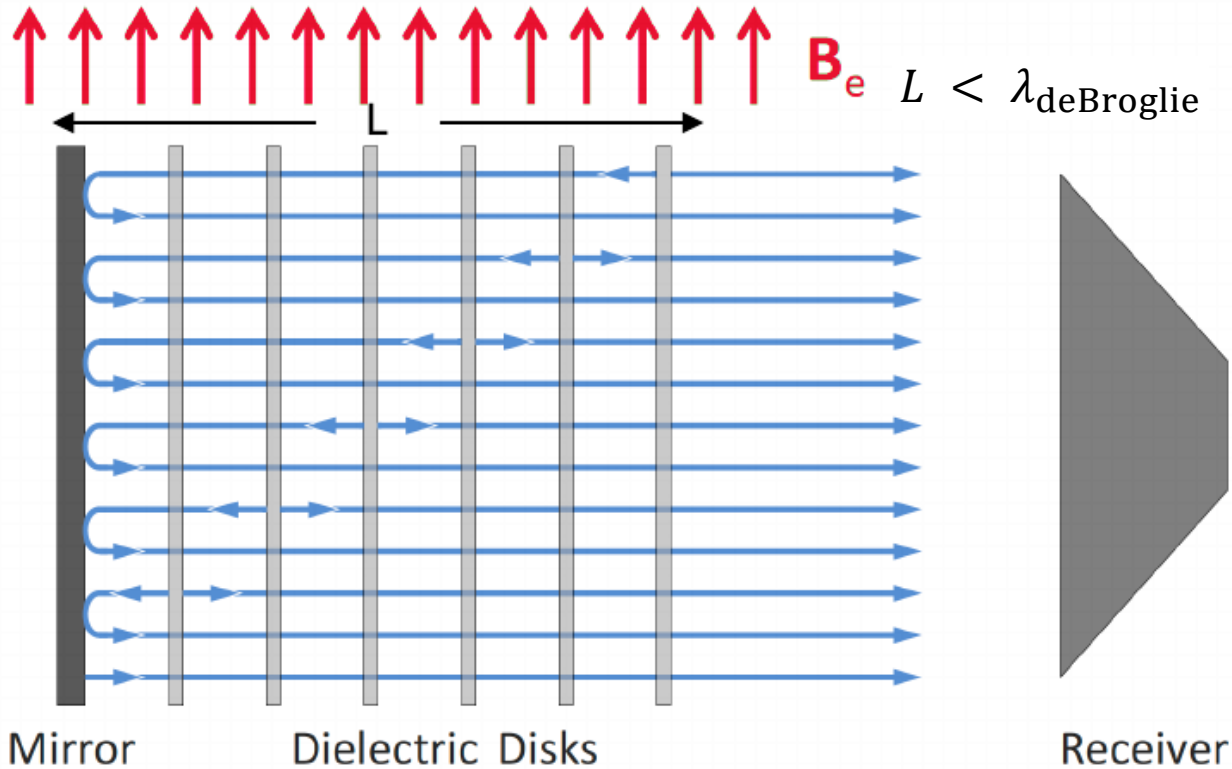
$$E_a \cdot \epsilon \sim 10^{-12} \text{ V/m for } B_e = 10 \text{ T}$$

$E_a$  is different in materials with different  $\epsilon$

At the surface,  $E_{||}$  must be continuous  
 $\rightarrow$  Emission of electromagnetic waves

Power emitted from a single surface:  $P/A = 2.2 \cdot 10^{-27} \frac{\text{W}}{\text{m}^2} C_{a\gamma} \left( \frac{B}{10 \text{ T}} \right)^2 \mathcal{O}(C_{a\gamma}) = 1$

# Dielectric Haloscope



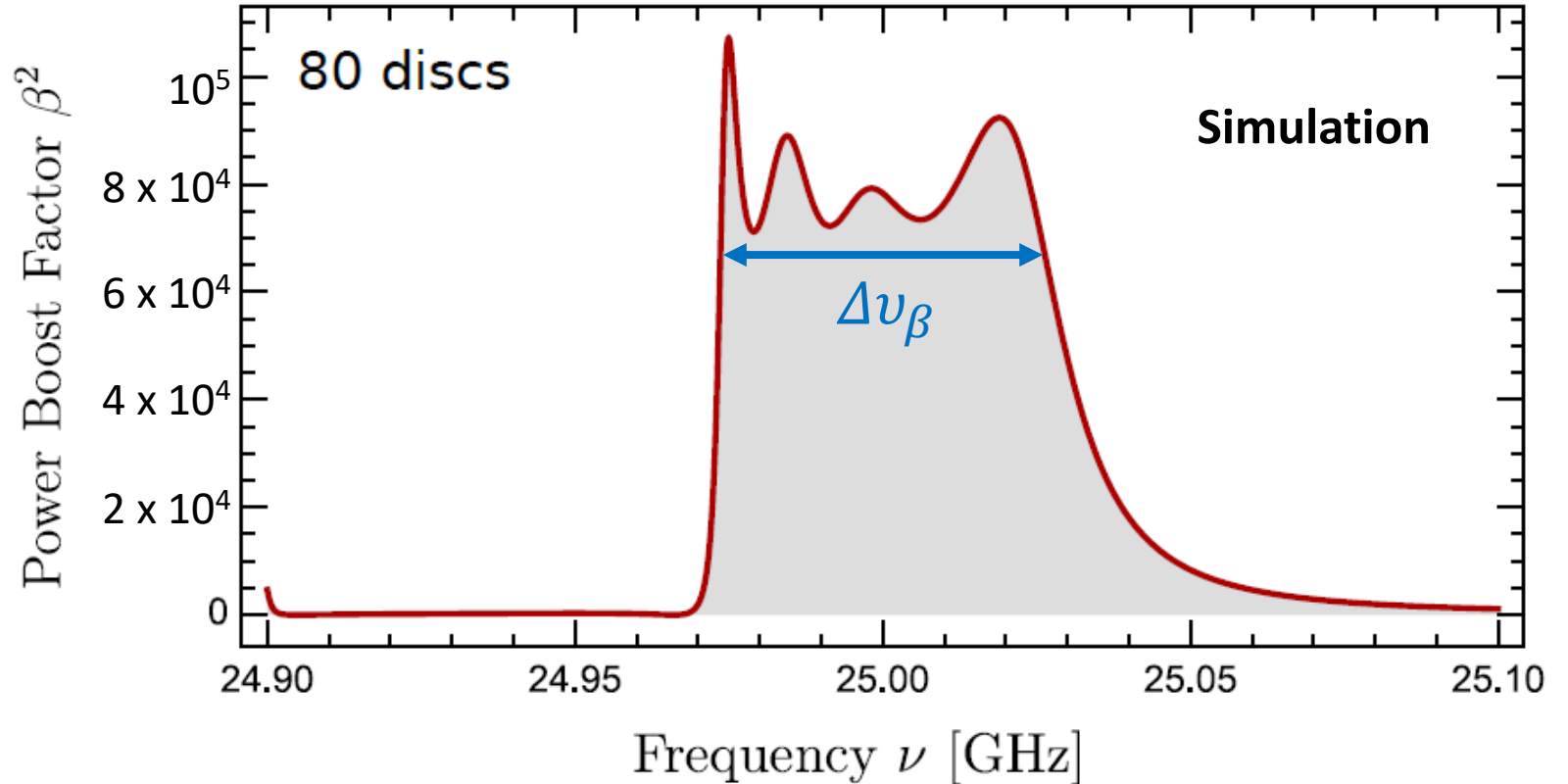
- Boost emitted power through:
- coherent emission from multiple interfaces
  - constructive interference effects

Power boost factor:

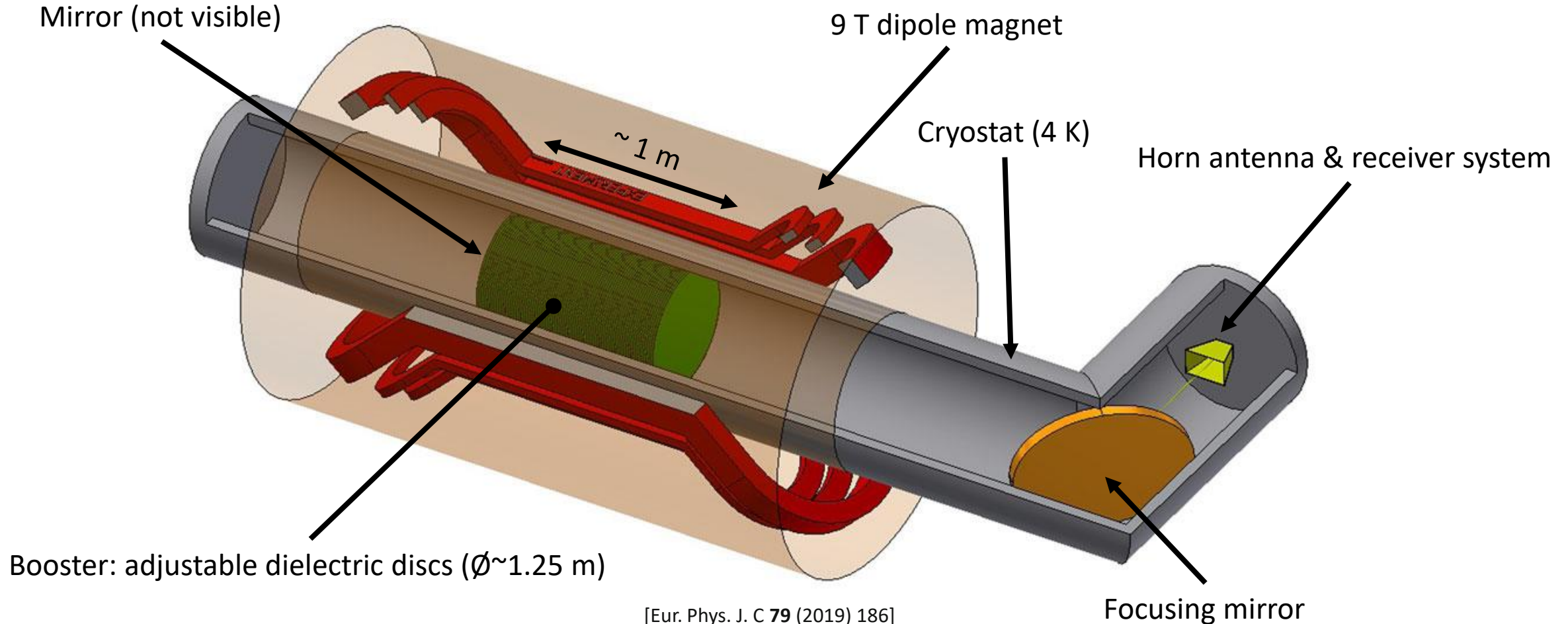
$$\beta^2 = P_{\text{total}} / P_{\text{mirror}}$$

Power emitted from all interfaces:  $P/A = 2.2 \cdot 10^{-27} \frac{\text{W}}{\text{m}^2} C_{a\gamma} \left( \frac{B}{10 \text{ T}} \right)^2 |\beta^2|$

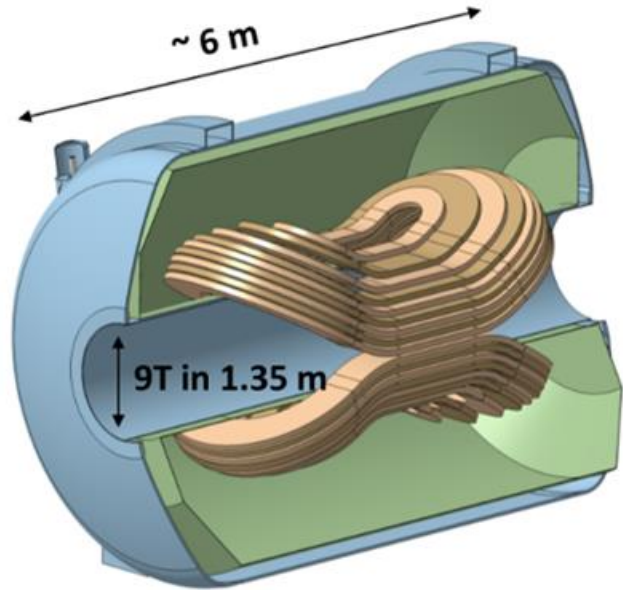
- In perfect world (1D simulation):  
 $|\beta^2| > 10^4$  achievable  
with 80 discs and  $\epsilon = 24$
- Non-uniform disk spacing  
of  $\sim \lambda/2$  can achieve  
broadband response
- Tuning of sensitive  
frequency range by  
adjusting disc spacing
- Area law:  $\beta^2 \Delta\nu_\beta \sim \text{const.}$



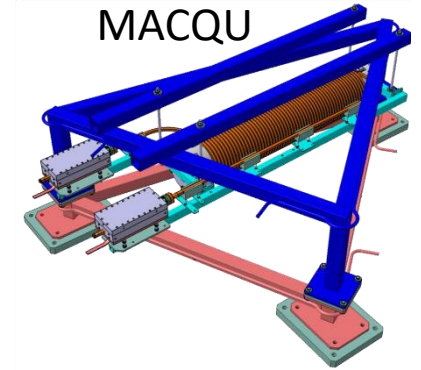
## MAGnetized DISC and MIRROR Axion eXperiment



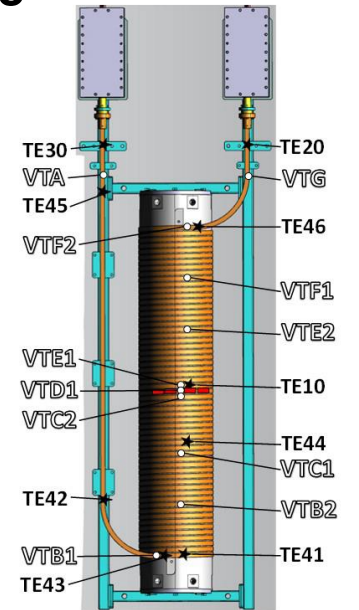
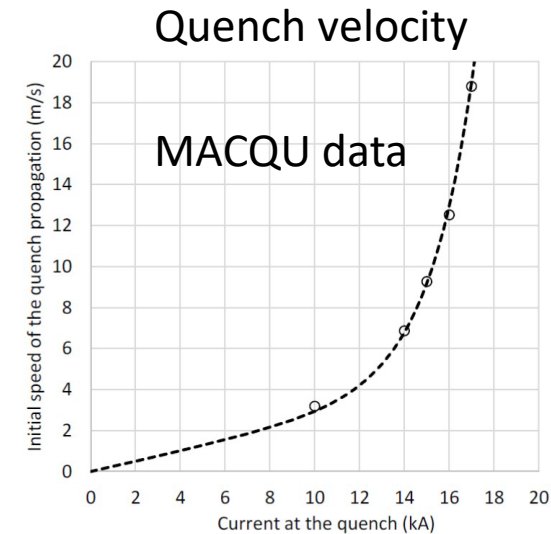
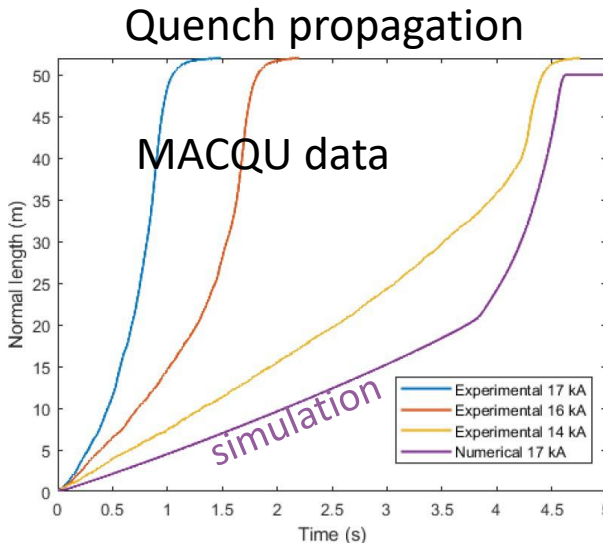
# MADMAX Magnet Update



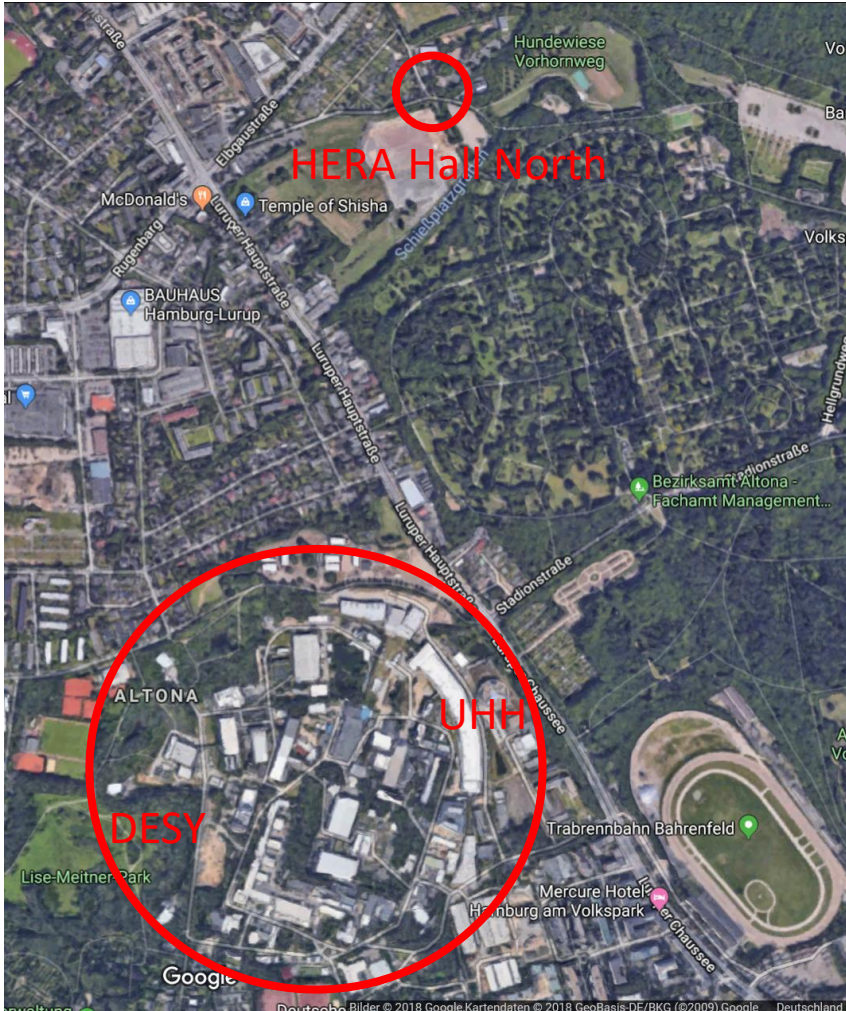
- Dipole Magnet most critical item for full-size MADMAX
- Design for 9 T large bore conceptually very well advanced
- **Novel conductor: cable in copper conduit**  
→ **production is feasible**
- Quench propagation velocity was measured in dedicated setup: **MA**dmax **C**oil for **Q**uench **U**nderstanding  
→ Main project risk mitigated: **Quench propagation according to requirements for safe operation**



Development in innovation partnership



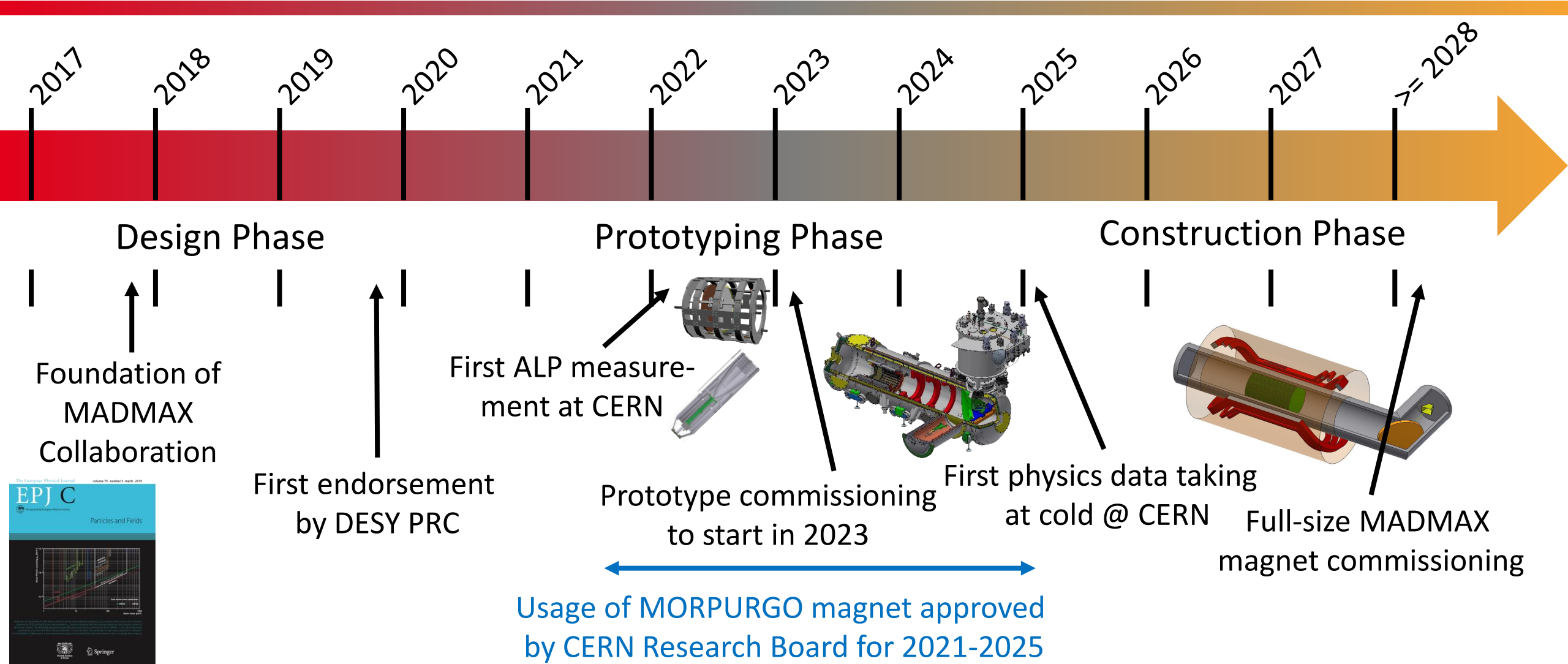




- MADMAX to be operated at HERA Hall North
- Make use of DESY infrastructure  
→ Cryoplatform to be operational in 2025
- Benefit: re-use H1 yoke as magnetic shielding



# Time Scale



# The MADMAX Prototype



Prototype cryostat (4 K)

Prototype booster

Focusing mirror

~ 3.2 m

~ 4.2 m

RF absorbers

Horn antenna & receiver

Receiver cryostat

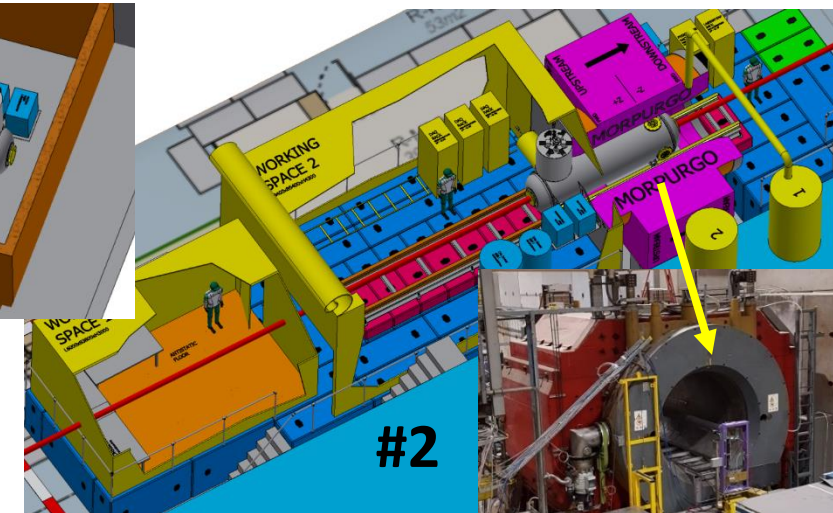
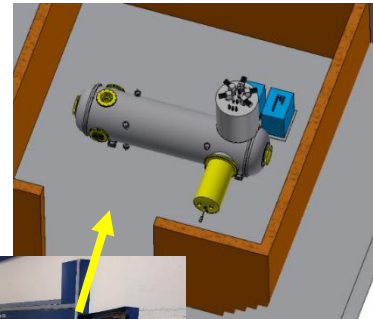
- Scaled-down version of MADMAX:
  - Reduced number of discs
  - 1/16 disc area
  - 1/5 magnetic field
- **Main goal #1:** Demonstrating and prototyping key technologies
- **Main goal #2:** Competitive ALP search with a dielectric haloscope

CLUSTER OF EXCELLENCE  
QUANTUM UNIVERSE



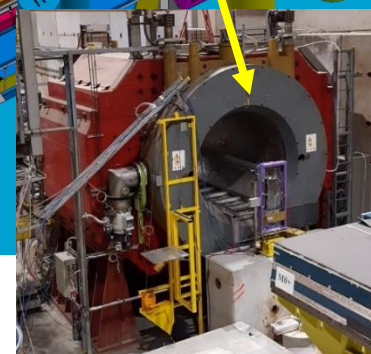
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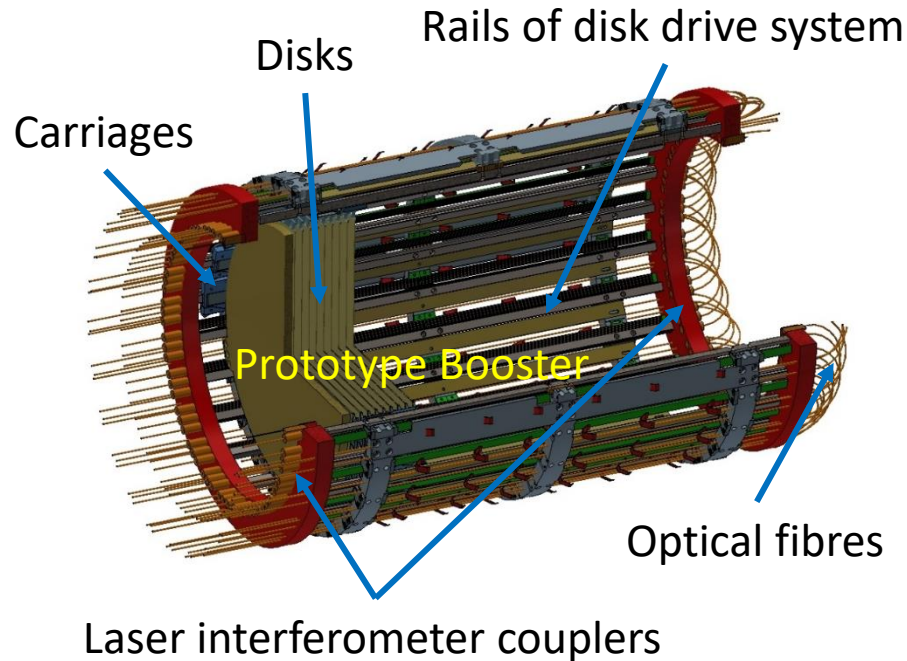
SHELL @Hamburg



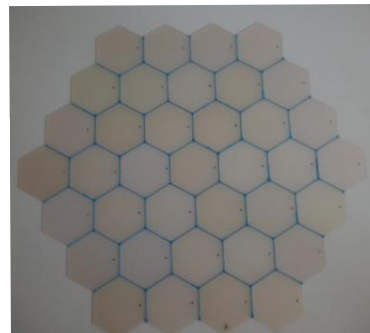
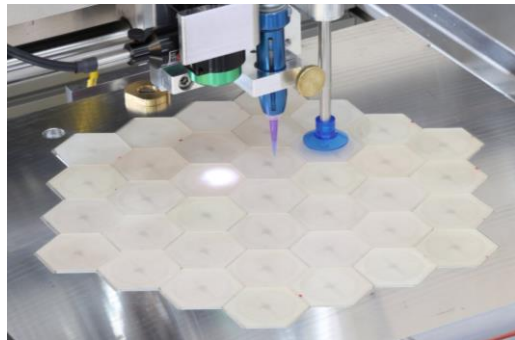
#2

MORPURGO @ CERN



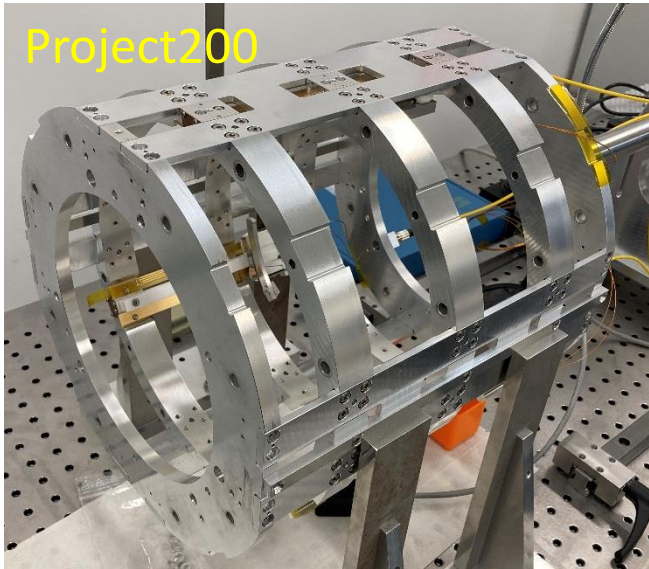


- Booster is the heart of MADMAX
- Need to manipulate many large area discs with precision  $< 10 \mu\text{m}$
- Operating conditions:
  - Cryogenic temperatures: 4 K
  - High magnetic field: up to  $\sim 10 \text{ T}$
  - Vacuum or cold gHe exchange gas
- Long travel range
- Disk weight: 600 g for  $\varnothing 300 \text{ mm}$
- **Piezo-driven actuator system with feedback from laser interferometer with absolute precision**

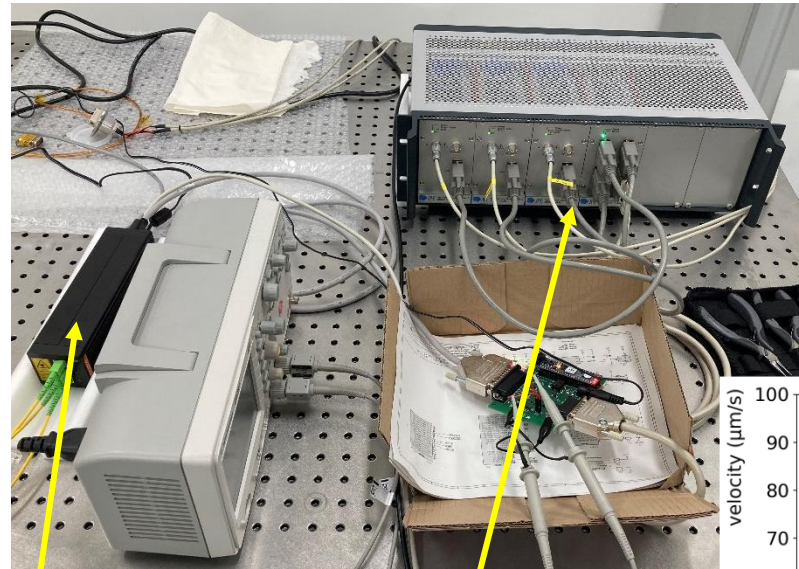


- Candidate disk materials:
  - **$\text{LaAlO}_3$**  ( $\epsilon \approx 24$ ,  $\tan\delta \approx \text{a few } 10^{-5}$ )
  - Sapphire ( $\epsilon \approx 9$ ,  $\tan\delta \approx 10^{-5}$ )
- $\text{LaAlO}_3$  available as 3" wafers at maximum
- **Tiling necessary  $\rightarrow$  Semi-automatic gluing machine**

# Testing the Disk Drive



Project200



Laser interferometer

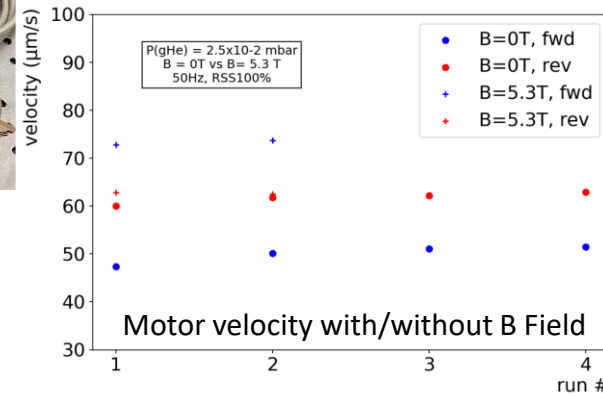
Piezo controllers

- Project200 built as **mechanical demonstrator**
- Three JPE piezo actuators on self-built carriages
- **Piezo controller system for driving a disc with three motors**
- Attocube laser interferometer for displacement measurement
- Single 200 mm sapphire disk in titanium disc ring can be mounted



## Single motor tests

- Motor developed by company JPE
- Motor successfully tested at RT and 4 K
- Begin of 2022 test in ALPS II magnet  
→ **Motor works in 5.3 T field and at 5 K**



Single motor test rig

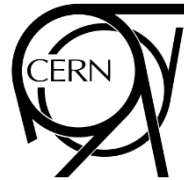


ALPS II magnet test stand

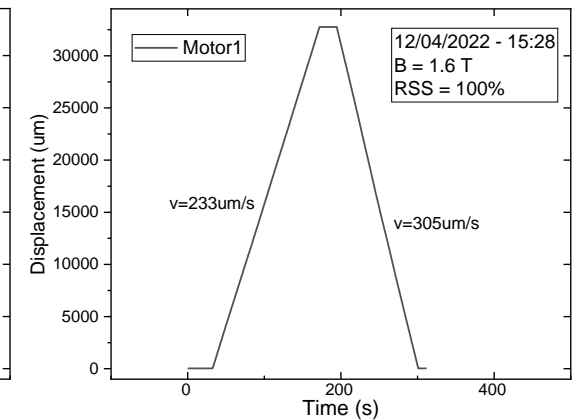
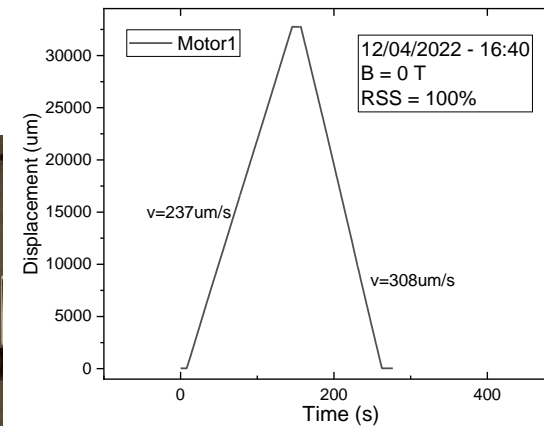


# Testing the Disk Drive

- Project200 successfully tested at CERN Cryolab and in CERN's Morpurgo magnet
- All three piezo motors work at cryogenic temperatures and in 1.6 T field (at RT)
- Attocube laser interferometer works at cryogenic temperatures
- Project200 backbone structure keeps optics alignment during cool-down
- A disc can be moved with three motors using the laser interferometer feedback

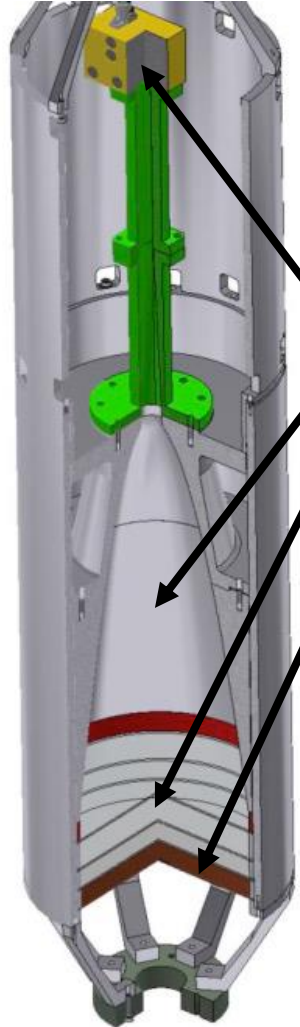


No difference in disk velocity with/without B field



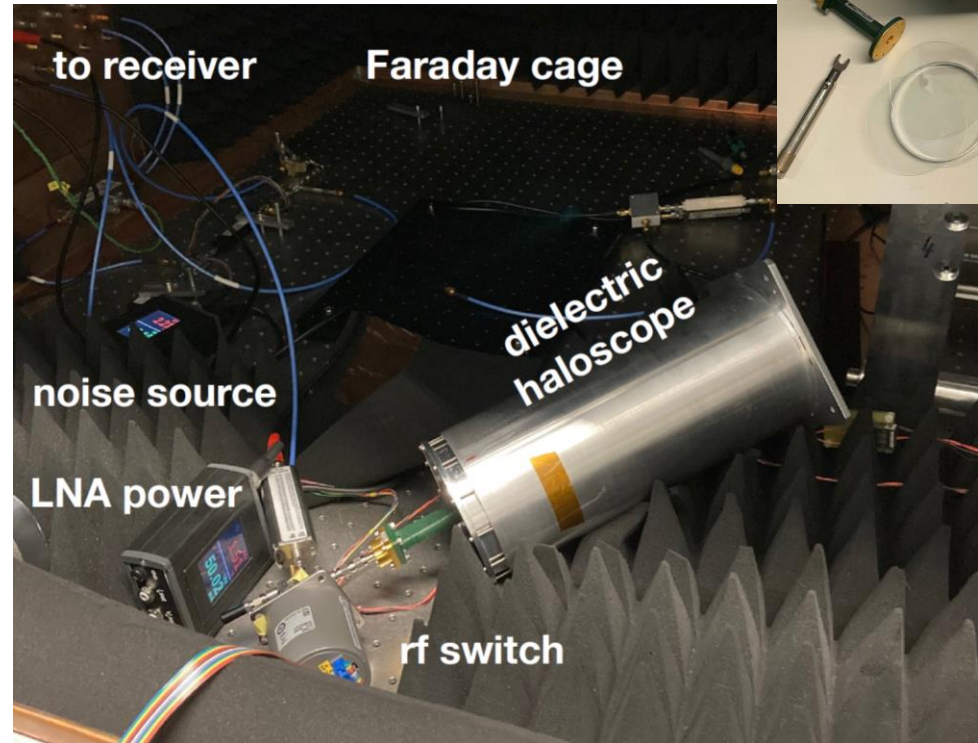
Project200 inside Morpurgo

# Closed Booster System

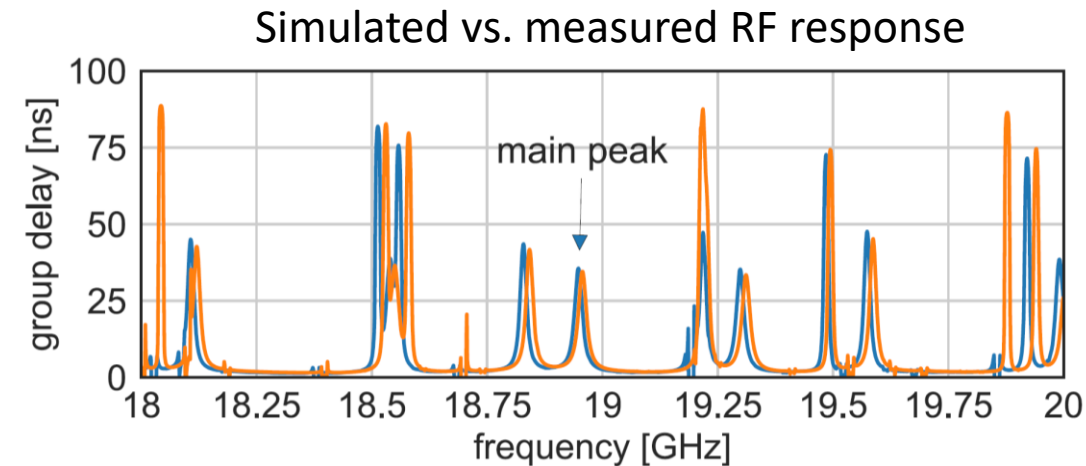
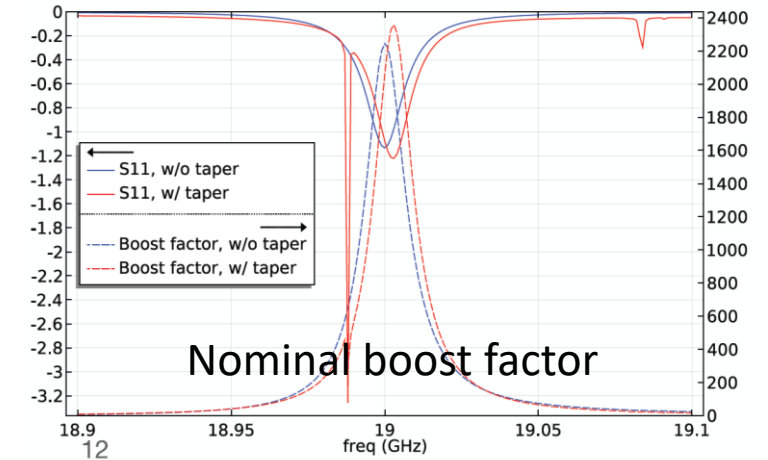
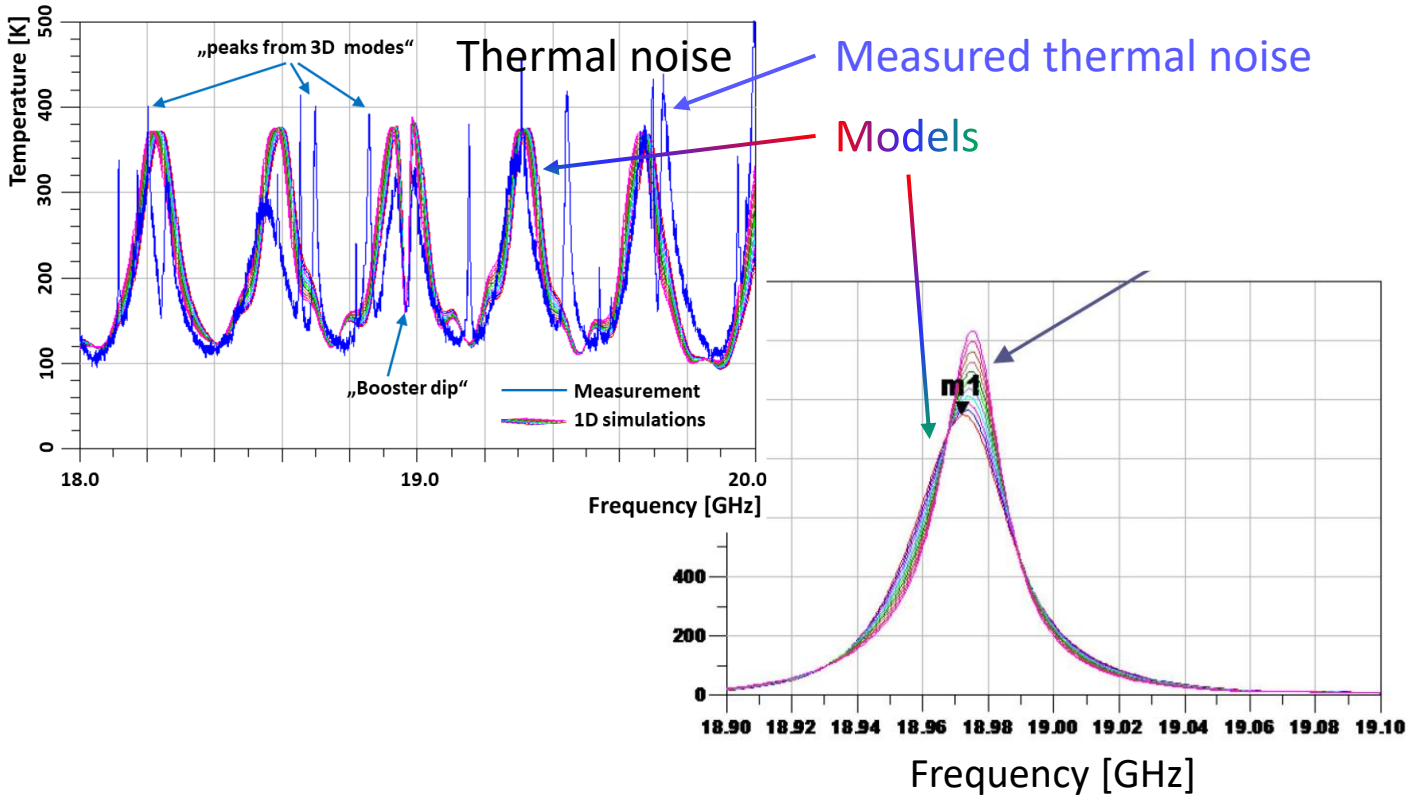


- Simple closed system to understand behaviour
- Can be operated at cryogenic temperatures
- Hidden Photon search and ALP search with Closed Booster 100

- Receiver
- Parabolic taper
- 3x Ø100 mm sapphire disks (fixed distances)
- Copper mirror

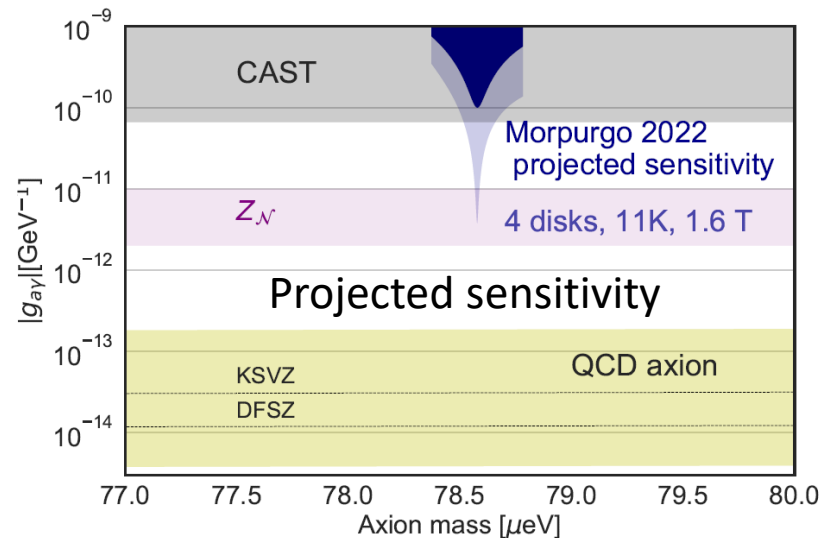
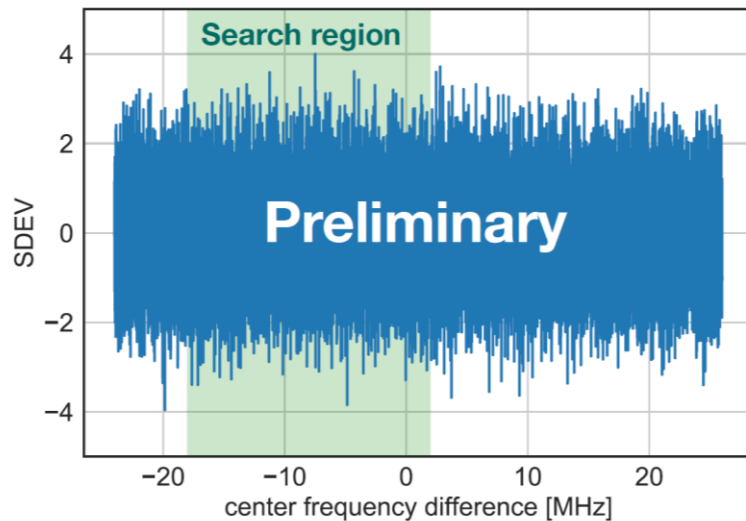


- Measurement of thermal noise & RF response (reflectivity and/or group delay)
- Match model to reproduce measured quantities
- Extract boost factor from model

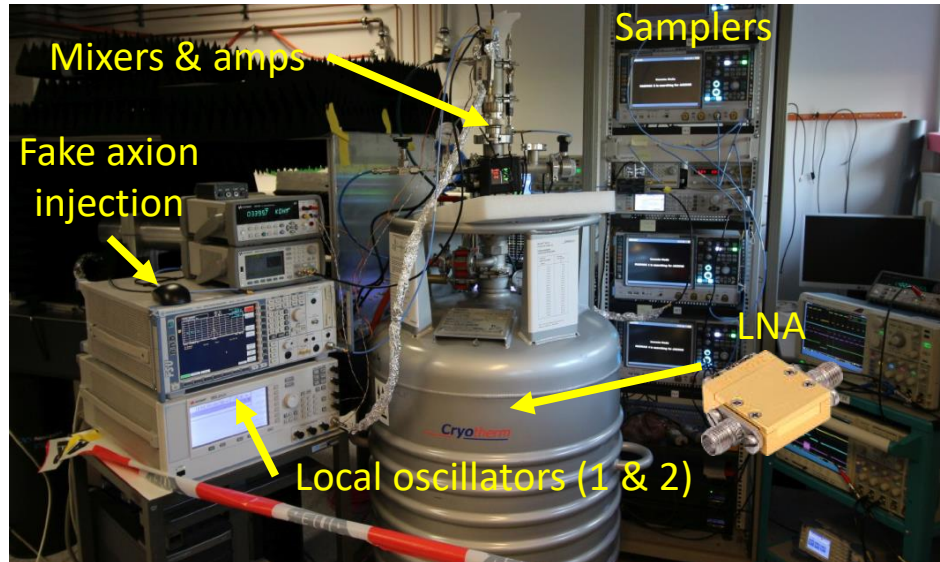
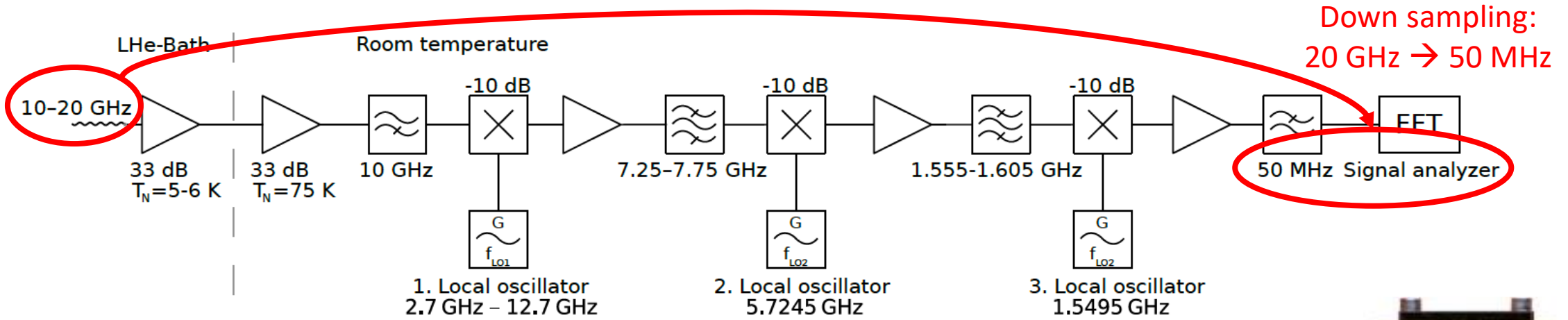




- Opportunity to perform ALP search in CERN's Morpurgo magnet (1.6 T) was used in Mar/Apr 2022
- In total 10 h at 1.6 T with  $\sim 200$  K noise temperature
- Sensitivity not dominated by RFI in CERN North Hall
- Possibilities for an upgrade allowing to cool the setup to  $< 10$  K in Morpurgo currently in preparation



# Receiver Chain



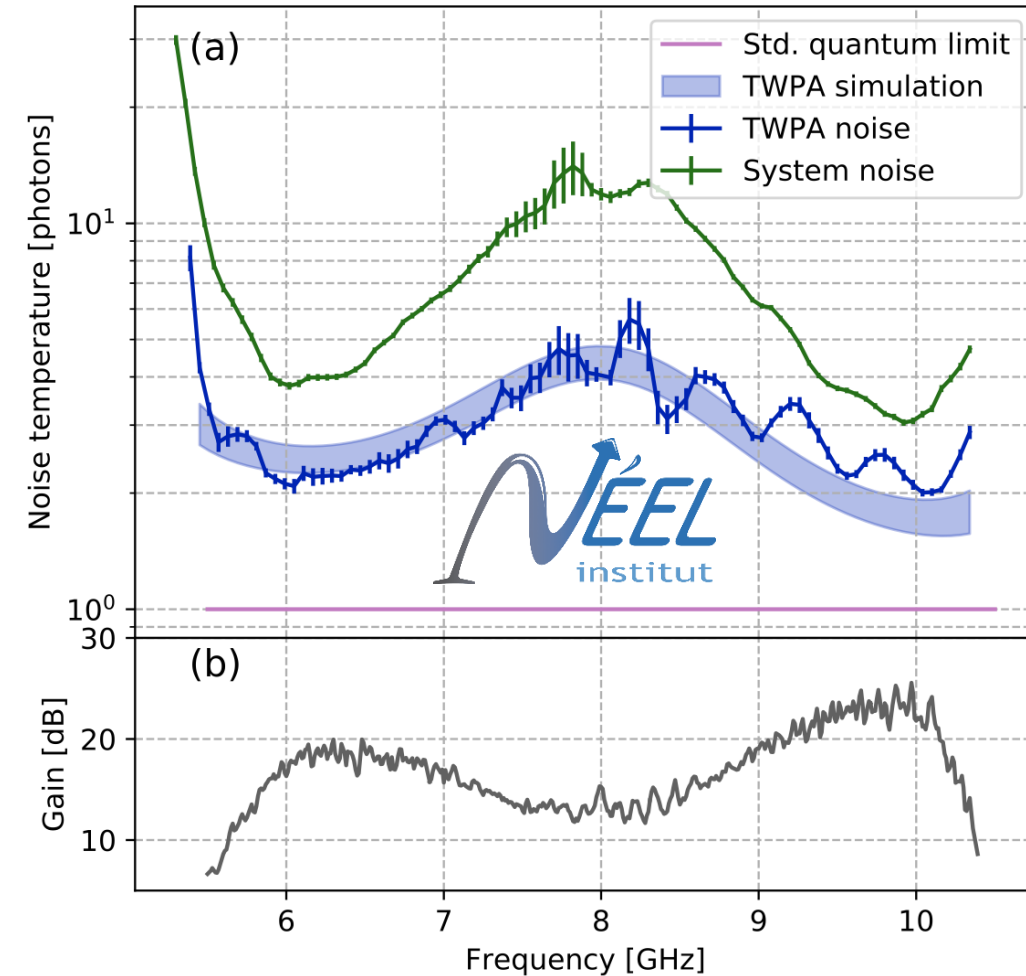
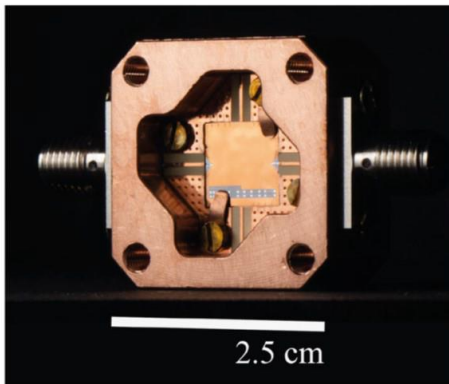
- Receiver chain with low-noise amplifier and three mixing stages
- Amplifiers for high frequencies developed: TWPAs for  $< 30$  GHz



Low-noise cryogenic amplifier (noise temperature 5 to 6 K)

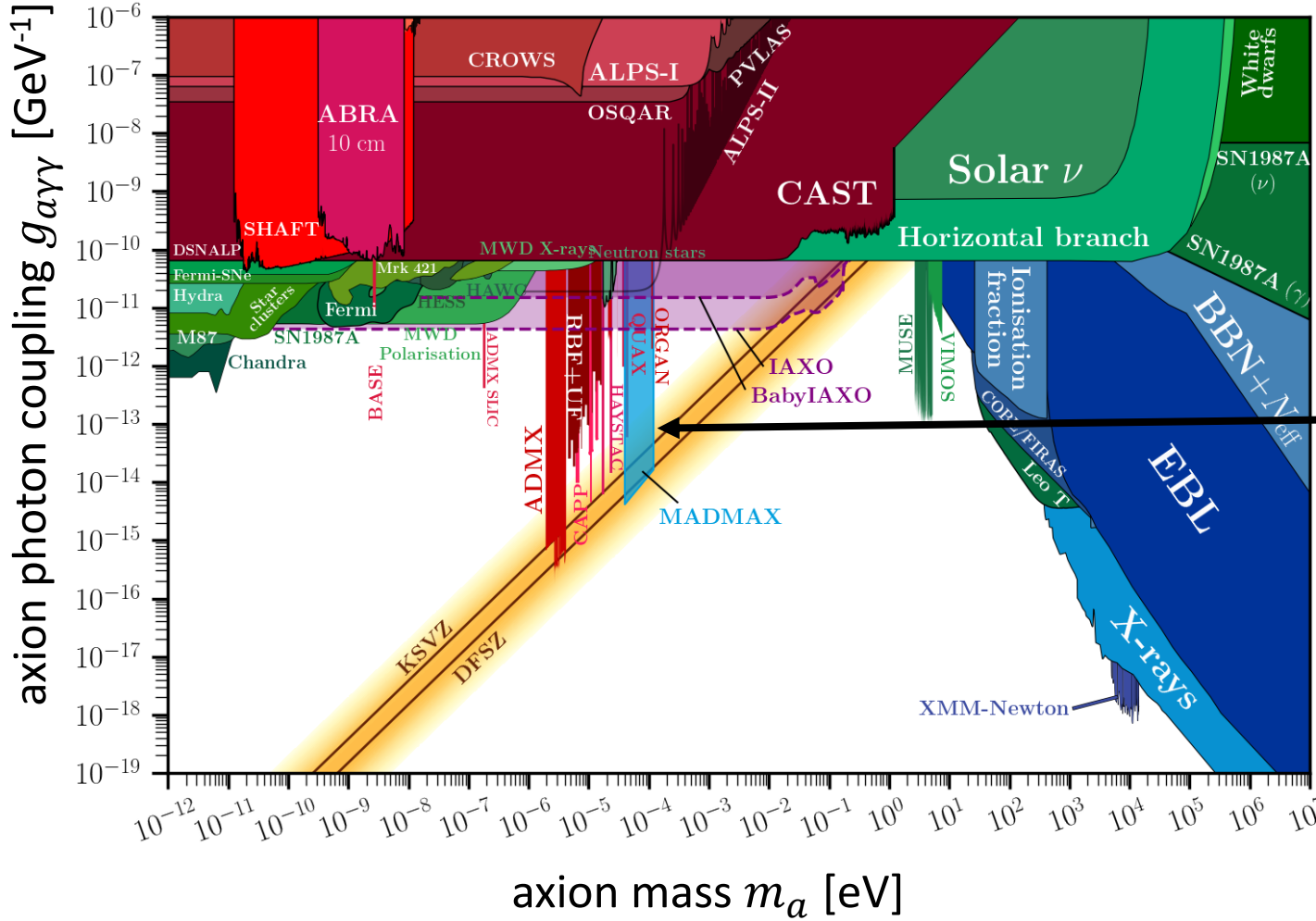
Test setup at MPP with 4 samplers and fake axion injection: Detection of  $1.2 \times 10^{-22}$  W signal within few days

- Traveling wave parametric amplifier (TWPA)
- **First 10 GHz TWPA** produced (PRX 10, 021021)
- Added noise: 1 K above quantum limit (20 dB gain @ 10 GHz)
- Future development to 30 GHz



[Reversed Kerr TWPA arXiv:2101.05815]

# MADMAX Sensitivity



MADMAX:

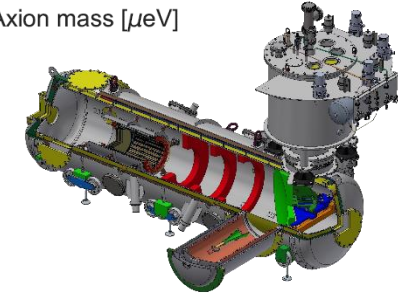
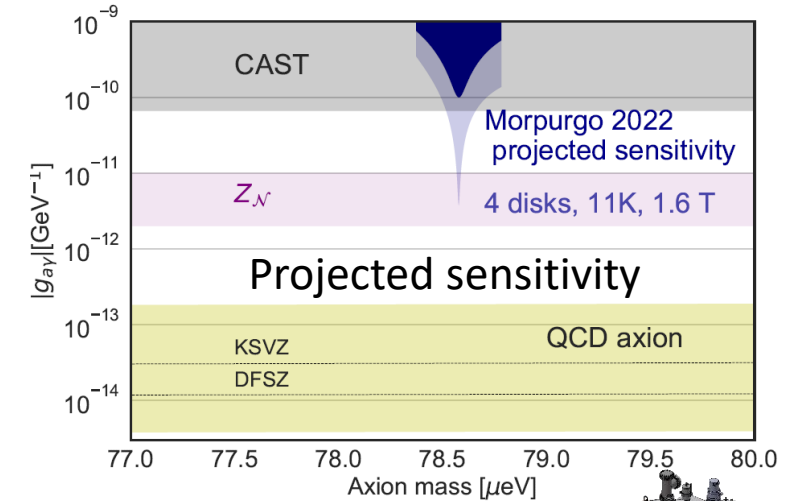
$$N_{\text{disc}} = 80$$

$$A_{\text{disc}} = 1.2 \text{ m}^2$$

$$B_{\parallel} = 9 \text{ T}$$

$$T_{\text{sys}} = 8 \text{ K}$$

- **MA**gnetized **D**isk and **M**irror **A**xion **eX**periment: dielectric haloscope to detect post-inflationary DM axions
- Continue work on analysis & calibration of Closed Booster 100
- Commissioning of full-size MADMAX at DESY in HERA Hall North starting (earliest) in 2028 with the huge magnet
- Commissioning of MADMAX Prototype to start in 2023  
 → First operation at CERN during YETS 2024/25



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# MADMAX Collaboration Meeting September 2022 @ Hamburg

