

The search for axion dark matter with a dielectric haloscope:

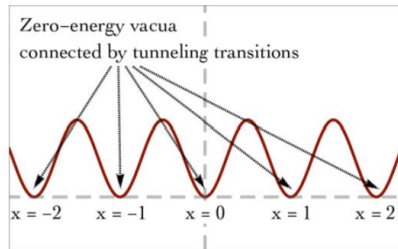


Béla Majorovits



MAX-PLANCK-INSTITUT FÜR PHYSIK

on behalf of the MADMAX collaboration



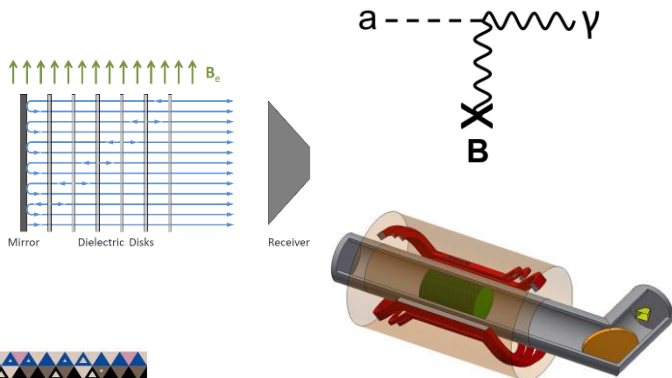
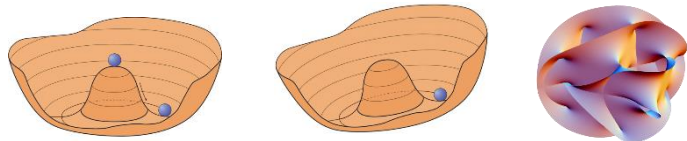
- The strong CP problem and the axion

- Axions and ALPs as dark matter

- How to detect axions

- Dielectric haloscope

- 



CP violation in the QCD vacuum

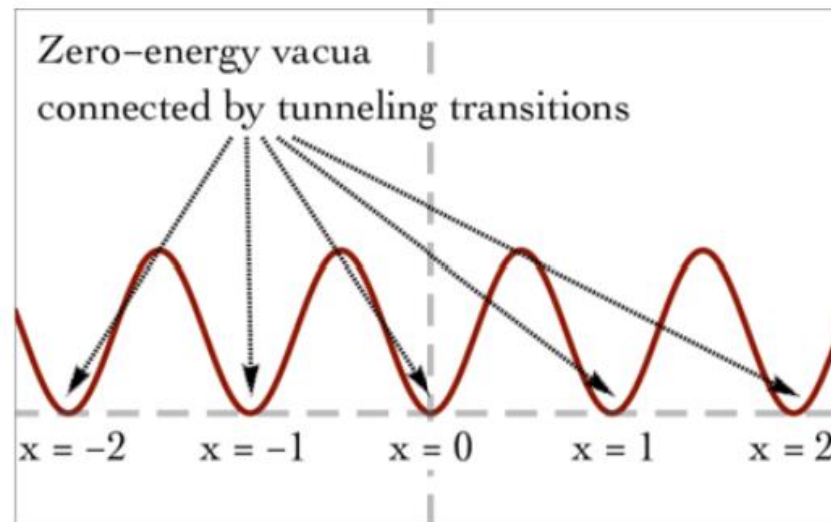
QCD SU(3) gauge group non-Abelian

gauge transformations not commutative!

→ Gauge in-equivalent zero energy states $|n\rangle$

→ Separated by potential barrier: tunneling → instantons

→ Single $|n\rangle$ **not stable** vacuum state



CP violation in the QCD vacuum

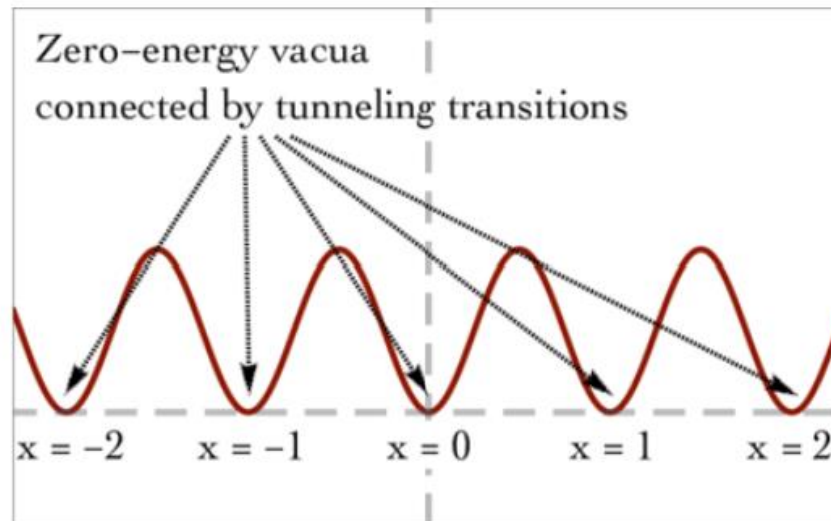
→ Physical ground state of QCD vacuum defined by

gauge invariant superposition of zero energy vacuum states:

θ – vacuum: $|\theta\rangle = \sum_n e^{in\theta} |n\rangle$

→ CP violating term in QCD Lagrangian:

$$\ominus \frac{\alpha_s}{8\pi} G_{\mu\nu a} \tilde{G}_a^{\mu\nu}$$



CP violation in QCD

CP violating phase through CKM matrix

→ Physically observable CP violation expected:

$$\bar{\Theta} = \Theta - \arg \det M_q$$

Random phase
from Θ -vacuum

phases from Yukawa coupling:
CKM matrix

CP violating term in QCD Lagrangian:

$$\bar{\Theta} \frac{\alpha_s}{8\pi} G_{\mu\nu a} \tilde{G}_a^{\mu\nu} \in \mathcal{L}_{\text{QCD}}$$

CP violation in QCD - the strong CP problem

CP violating term in QCD

→ induces neutron EDM:

$$\bar{\Theta} \frac{\alpha_s}{8\pi} G_{\mu\nu a} \tilde{G}_a^{\mu\nu} \in \mathcal{L}_{\text{QCD}}$$

$$d_n \sim \bar{\Theta} \cdot 10^{-16} \text{ e cm}$$

Limit on EDM of neutron:

$$d_n < 2 \cdot 10^{-26} \quad \text{Abel et al., Phys. Rev. Lett. 124, 081803 (2020)}$$

$$\bar{\Theta} < 10^{-10}$$

**Two seemingly independent terms
cancel each other to 1 in 10^{10}**

→ Strong CP problem

The strong CP problem ?

arXiv > hep-th > arXiv:2001.07152

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High Energy Physics - Theory

[Submitted on 20 Jan 2020 (v1), last revised 28 Sep 2021 (this version, v5)]

Absence of CP violation in the strong interactions

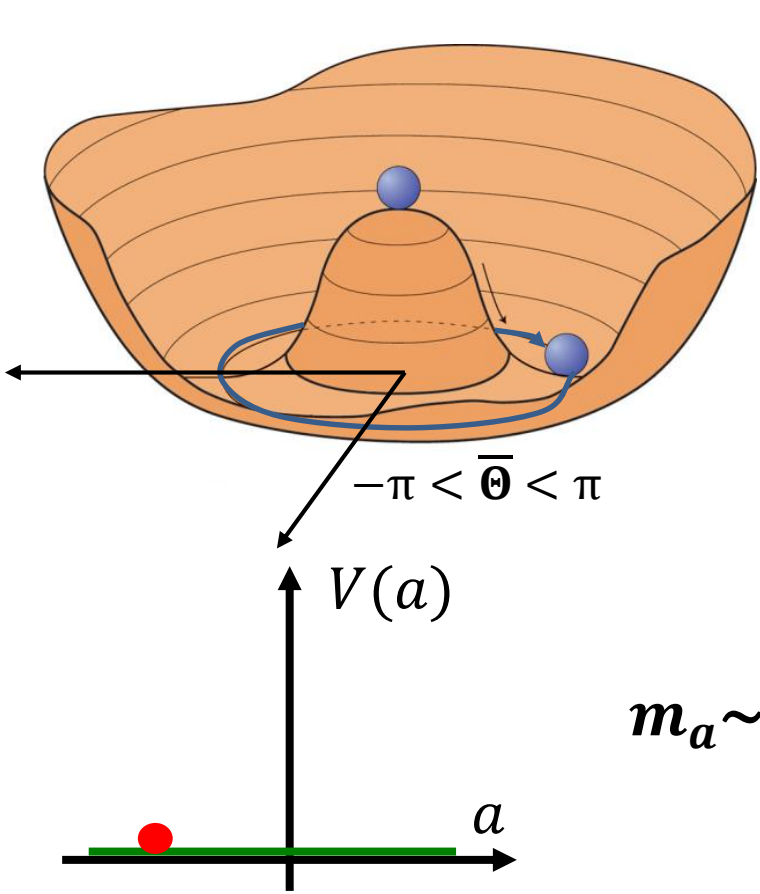
Wen-Yuan Ai, Juan S. Cruz, Bjorn Garbrecht, Carlos Tamarit

We derive correlation functions for massive fermions with a complex mass in the presence of a general vacuum angle. For this purpose, we first build the Green's functions in the one-instanton background and then sum over the configurations of background instantons. The quantization of topological sectors follows for saddle points of finite Euclidean action in an infinite spacetime volume and the fluctuations about these. For the resulting correlation functions, we therefore take the infinite-volume limit before summing over topological sectors. In contrast to the opposite order of limits, the chiral phases from the mass terms and from the instanton effects then are aligned so that, in absence of additional phases, these do not give rise to observables violating charge-parity symmetry. This result is confirmed when constraining the correlations at coincident points by using the index theorem instead of instanton calculus.

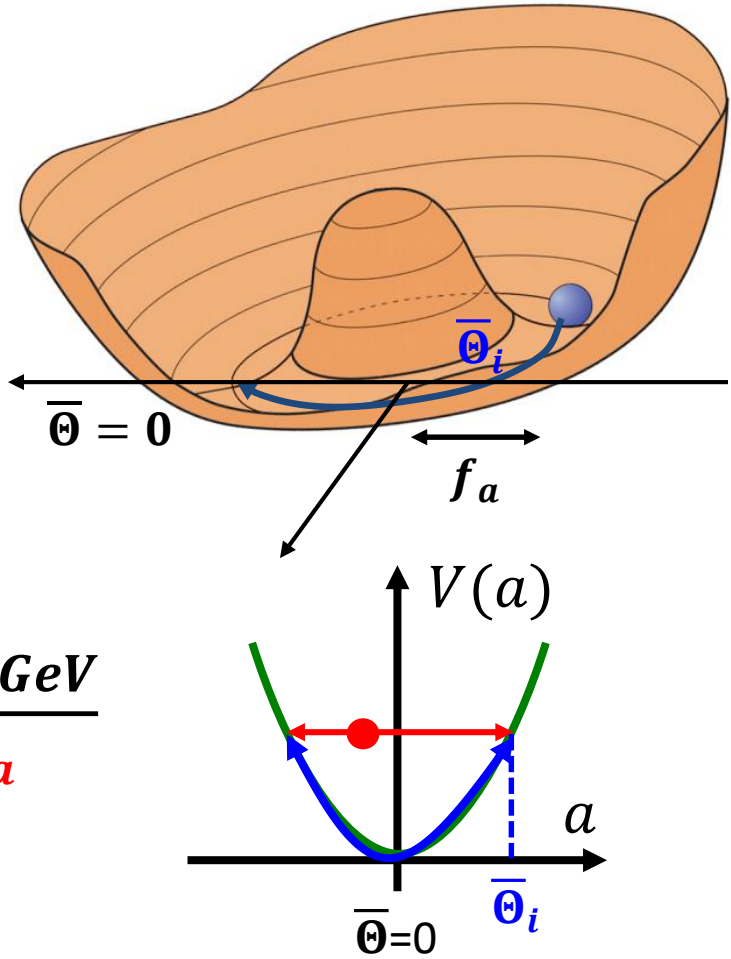
Comments: 52 pages, matches published version (except for the title that has been changed in journal) with supplementary material included
Subjects: **High Energy Physics - Theory (hep-th)**; High Energy Physics - Phenomenology (hep-ph)
Report number: TUM-HEP-1249/20, CP3-20-02
Cite as: [arXiv:2001.07152](https://arxiv.org/abs/2001.07152) [**hep-th**]
(or [arXiv:2001.07152v5](https://arxiv.org/abs/2001.07152v5) [**hep-th**] for this version)
<https://doi.org/10.48550/arXiv.2001.07152> 
Journal reference: Phys.Lett.B 822 (2021) 136616
Related DOI: <https://doi.org/10.1016/j.physletb.2021.136616> 

Axion as solution to strong CP problem

Make $\bar{\theta}$ dynamical \rightarrow U(1) with spontaneous Peccei Quinn symmetry breaking



QCD:
explicit
symmetry
breaking



$$m_a \sim 5.7 \mu eV \frac{10^{12} GeV}{f_a}$$

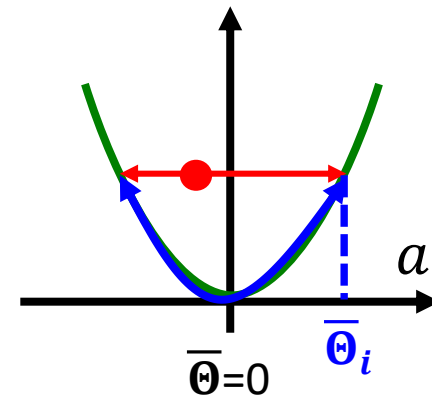
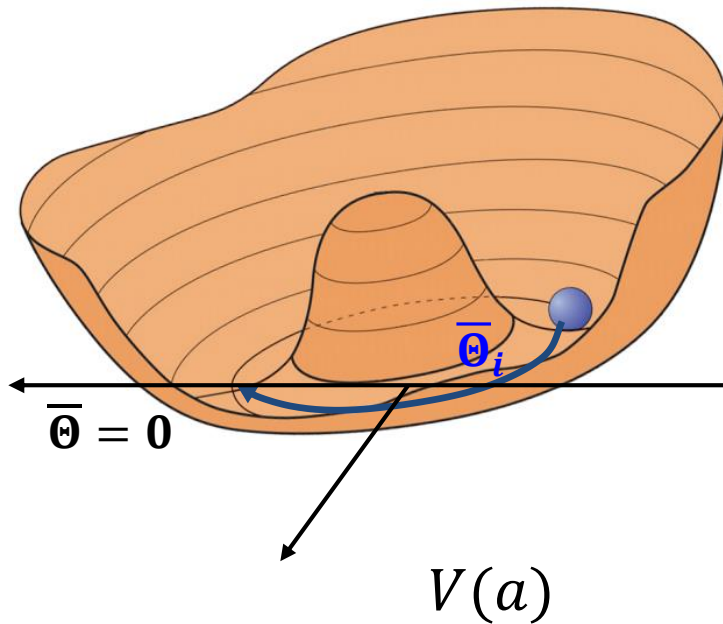


Axion as solution to strong CP problem

If axion exists:

→ **Contribution to Dark Matter:**

as relic oscillations of $\bar{\Theta}$ around minimum



Oscillations amplitude (particle density) damped by expansion of universe $H(t)$

Damping depends on ratio oscillation frequency (m_a) to $H(t)$



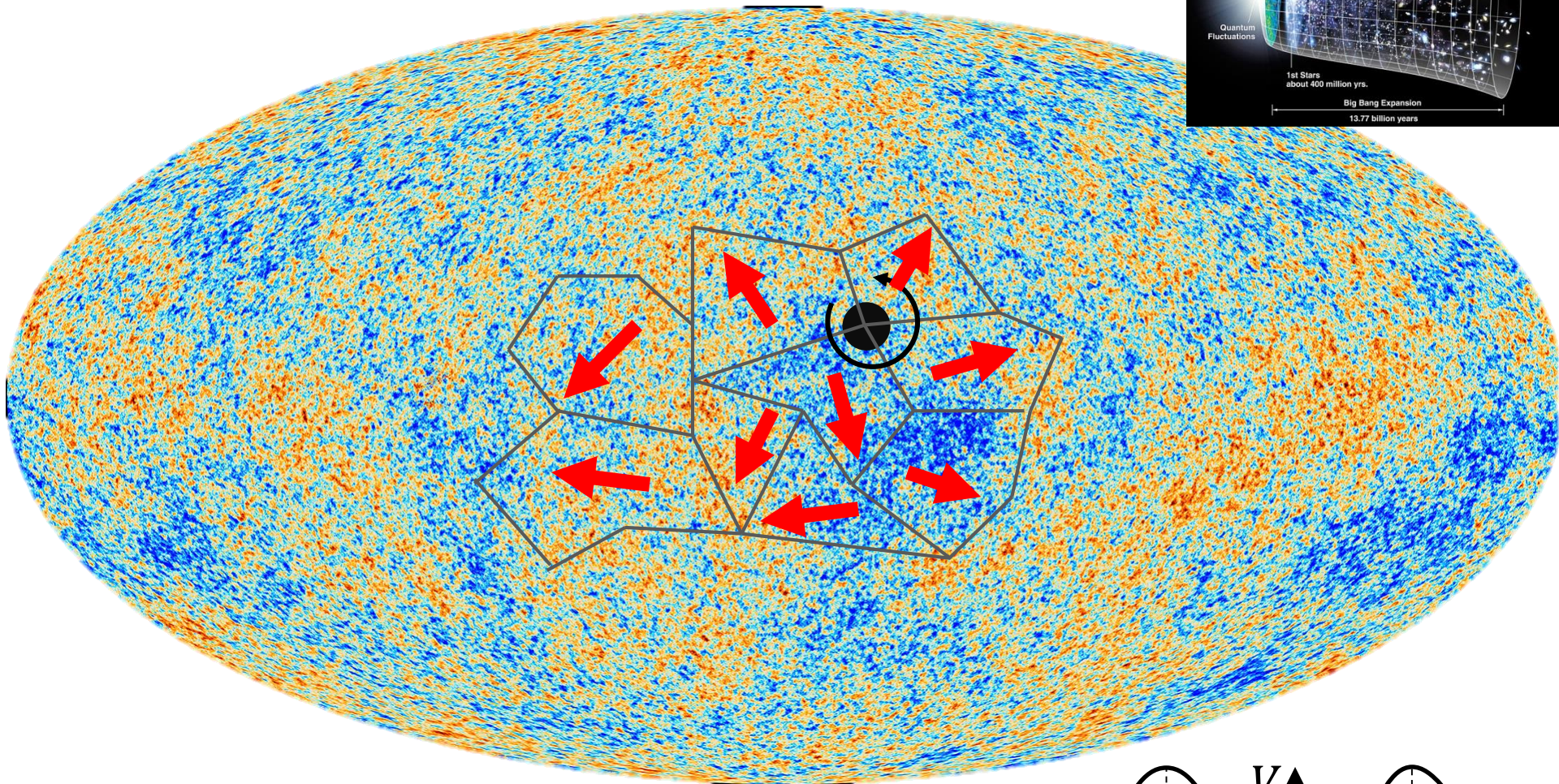
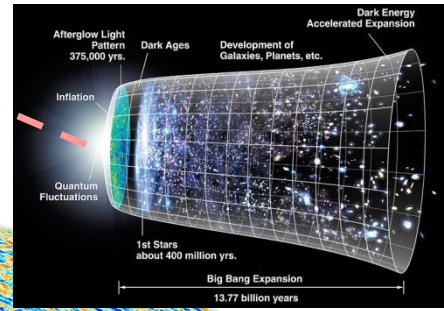
R. Peccei und H. Quinn,
Phys. Rev. Lett. **38**, 1440 (1977)
S. Weinberg, Phys. Rev. Lett. **40**, 223 (1978);
F. Wilczek, Phys. Rev. Lett. **40**, 279 (1978)

The Birth of Axions

Frank Wilczek
Institute for Advanced Study
Princeton, NJ 08540

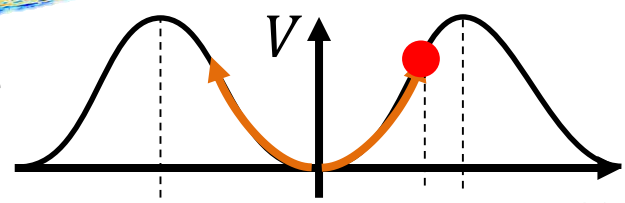
usual, very light particle. I called this particle the *axion*, after the laundry detergent, because that was a nice catchy name that sounded like a particle and because this particular particle solved a problem involving *axial* currents.

Pre-inflationary scenario

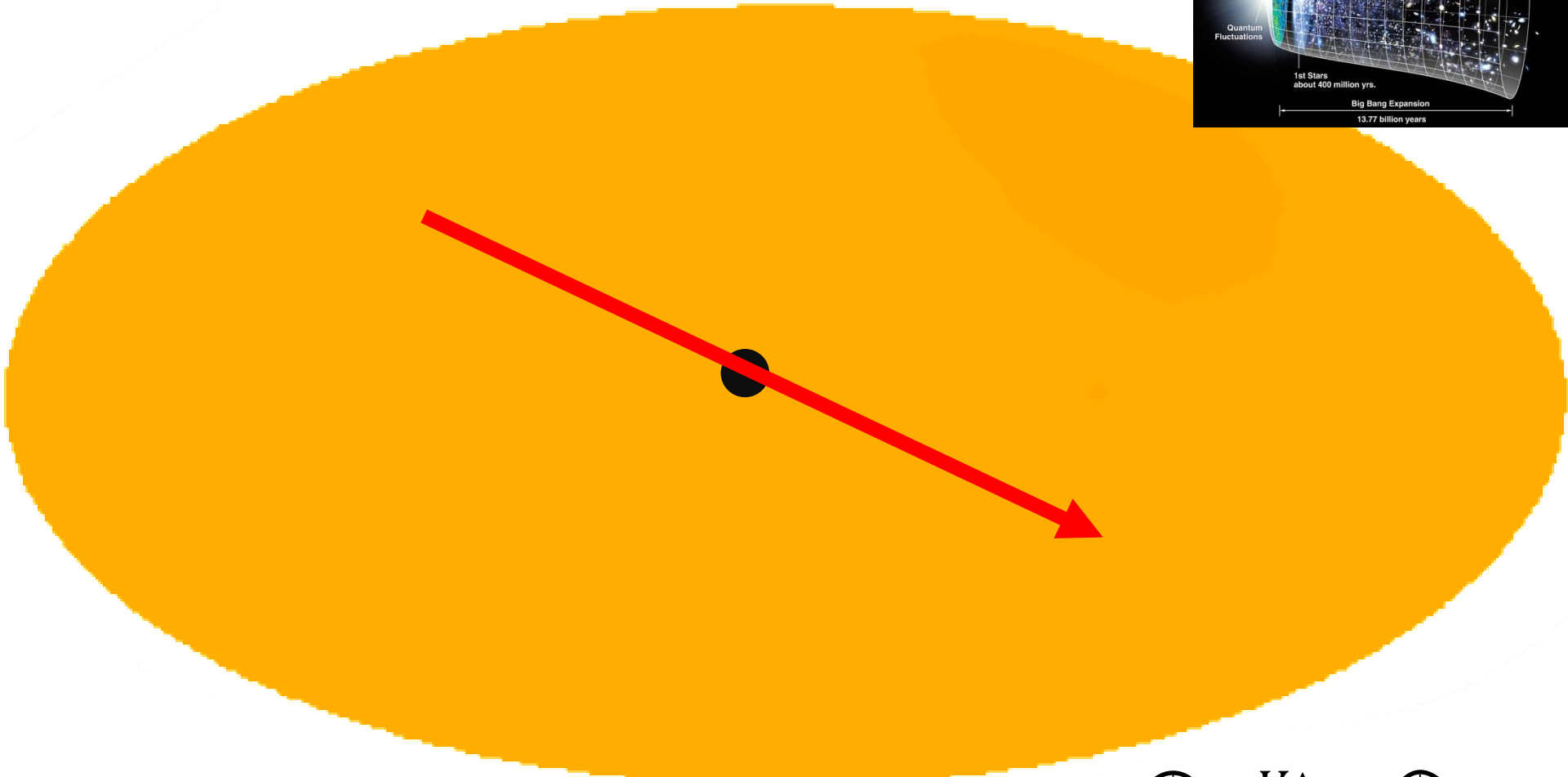
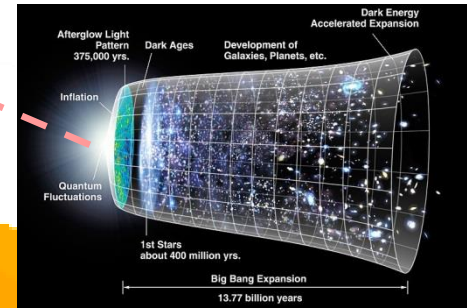


One value of $\bar{\theta}_i$ in entire visible universe

$$0 < |\bar{\theta}_i| < \pi$$

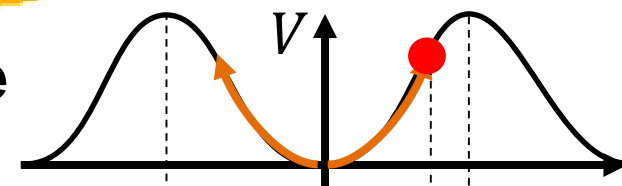


Pre-inflationary scenario



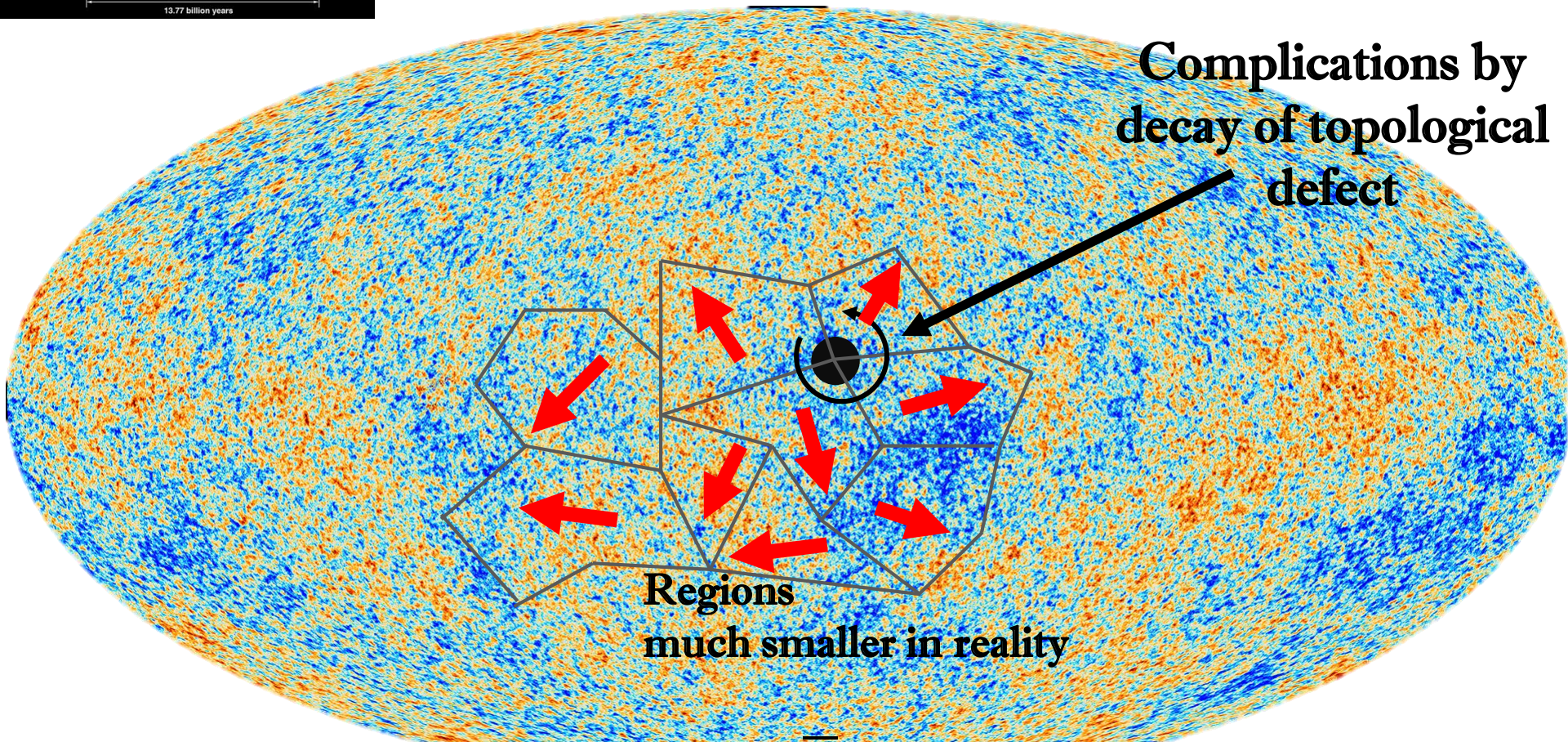
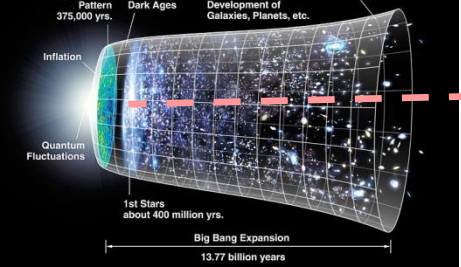
One value of $\bar{\theta}_i$ in entire visible universe

$$0 < |\bar{\theta}_i| < \pi$$

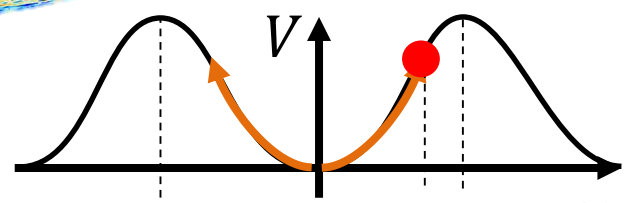


The search for axion dark matter with a dielectric haloscope: 

Post-inflationary Scenario

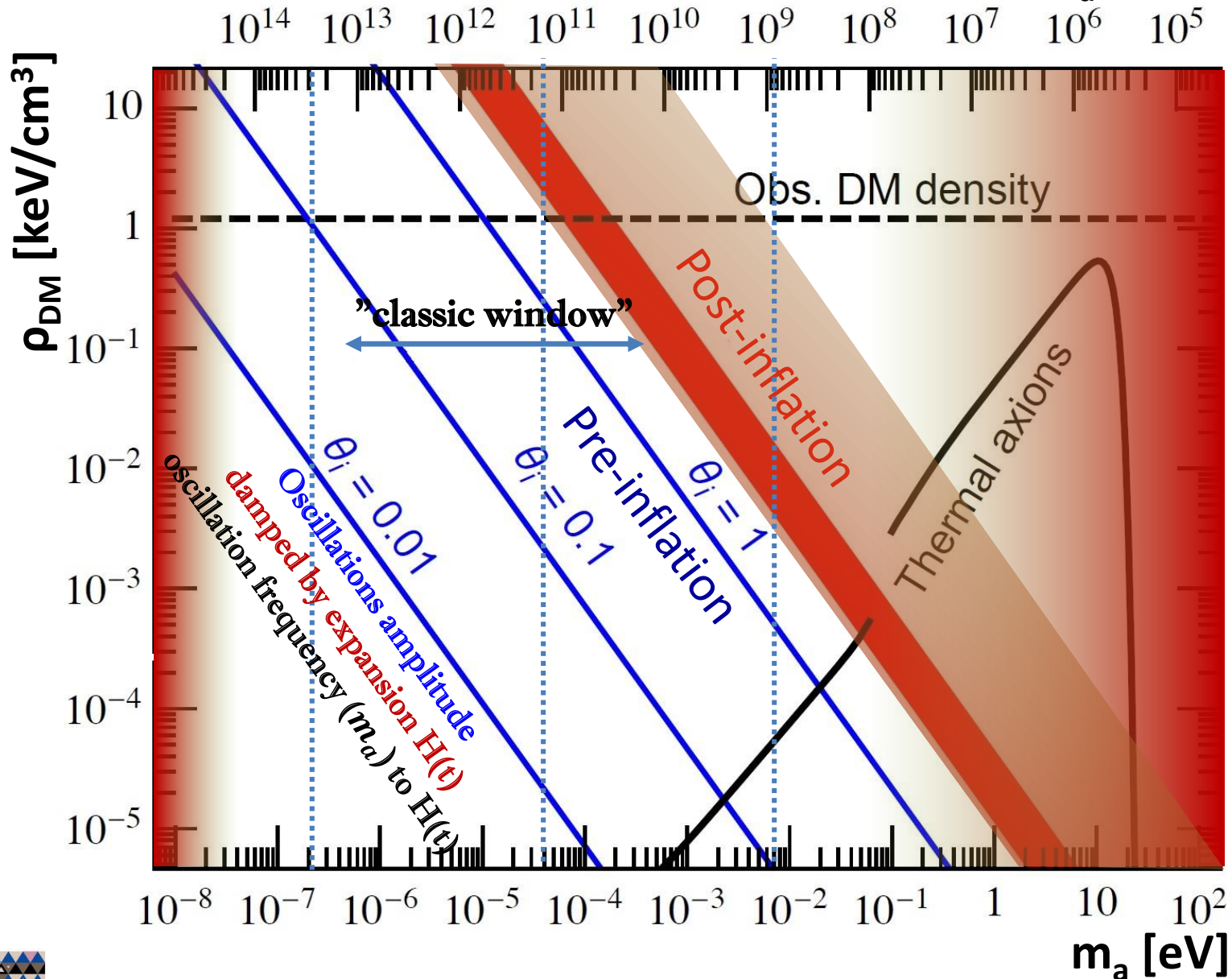


Average of all possible $\overline{\theta}_i$
→ Prediction for overall density

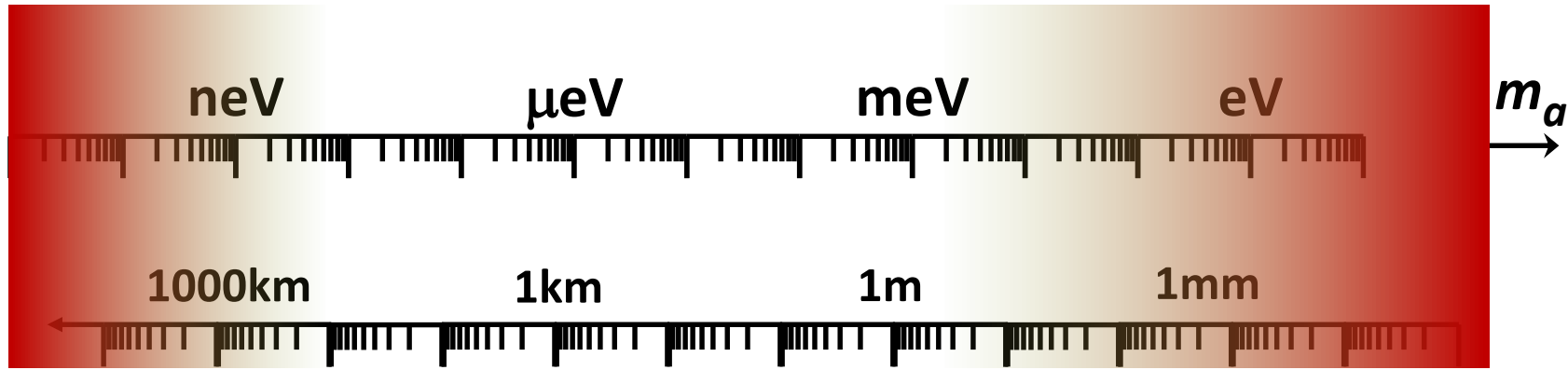


Axions as dark matter

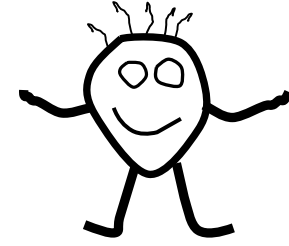
f_a [GeV]



The size of DM Axions



$\lambda_{\text{de Broglie}}$
 $\langle v_{DM} \rangle \sim 10^{-3} c$



Predictions:

$|\Theta_i|$ arbitrary $0 - \pi$



Pre-
Post-
inflationary scenario

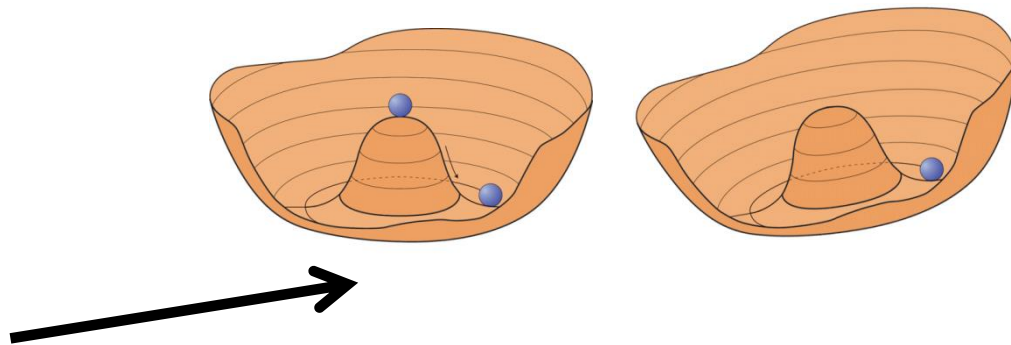
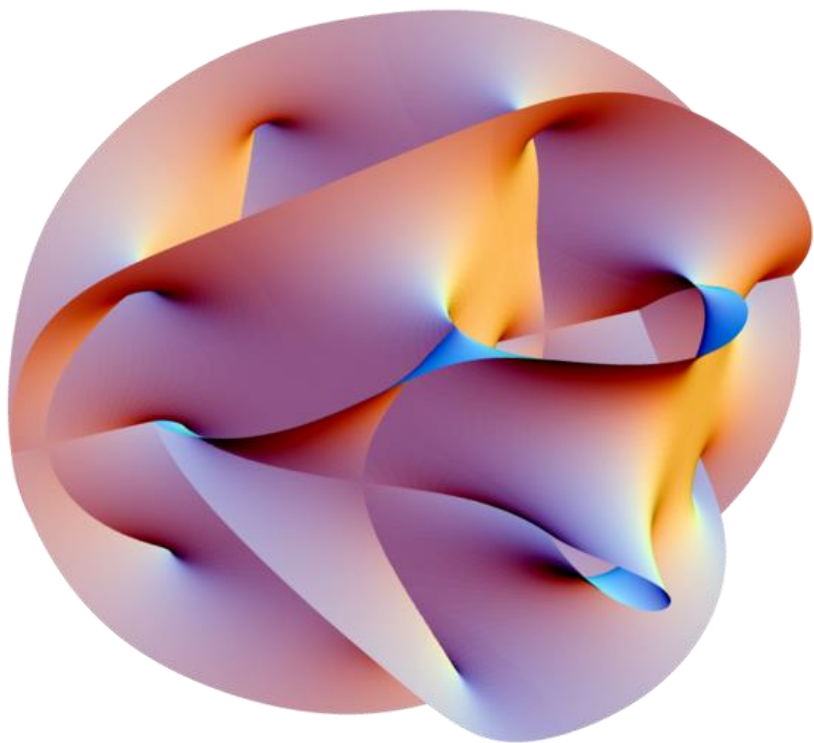
$|\Theta_i| \emptyset$ of all
possible values



DM axions fit into experiment!



ALPs emerging from string compactification: the Axiverse



No direct relation btw.
 m_{ALP} and f_{ALP}

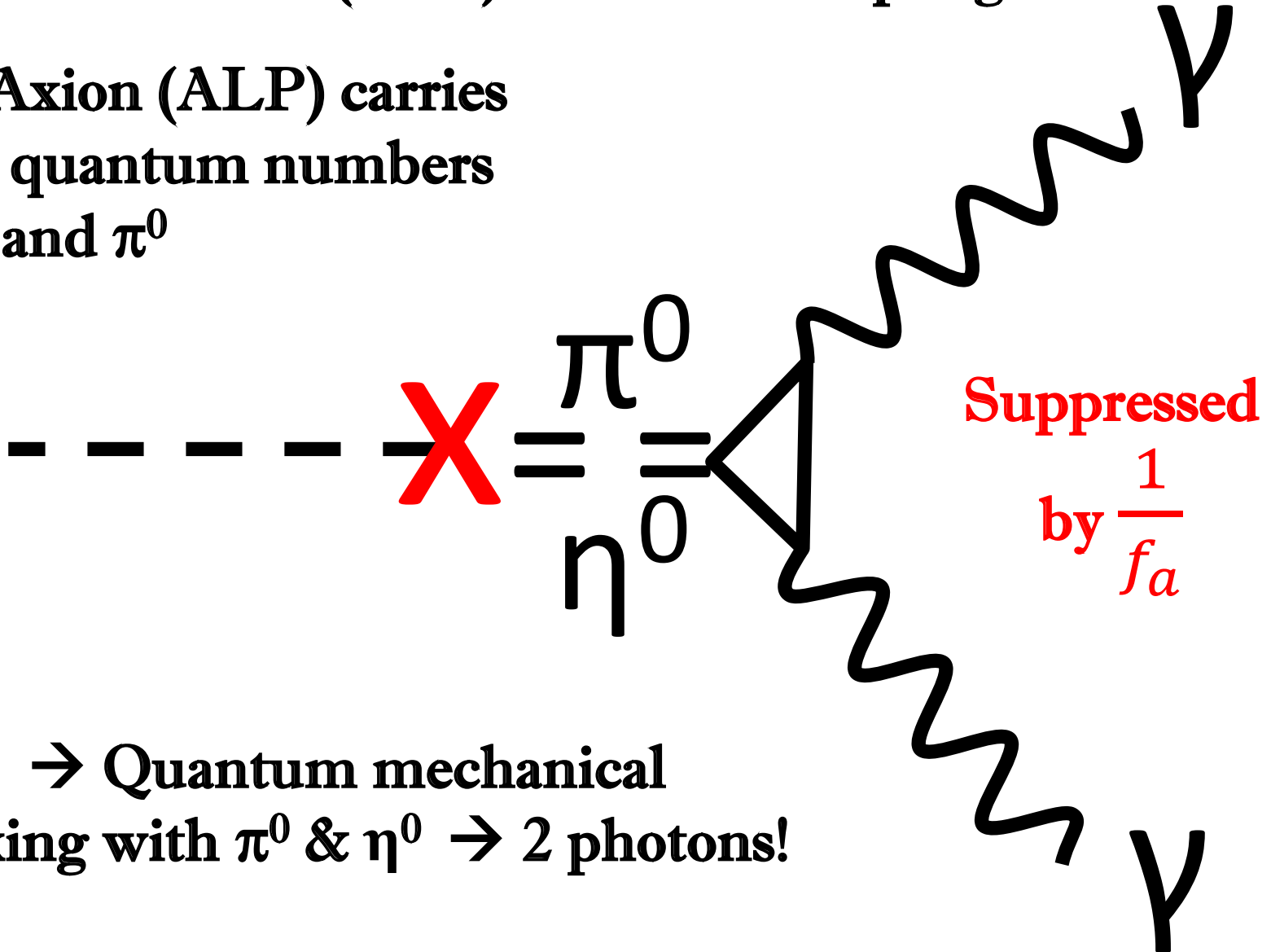
Some astrophysical inconsistencies:

- Transparency hint
- Cooling anomalies

Could be explained by ALPs

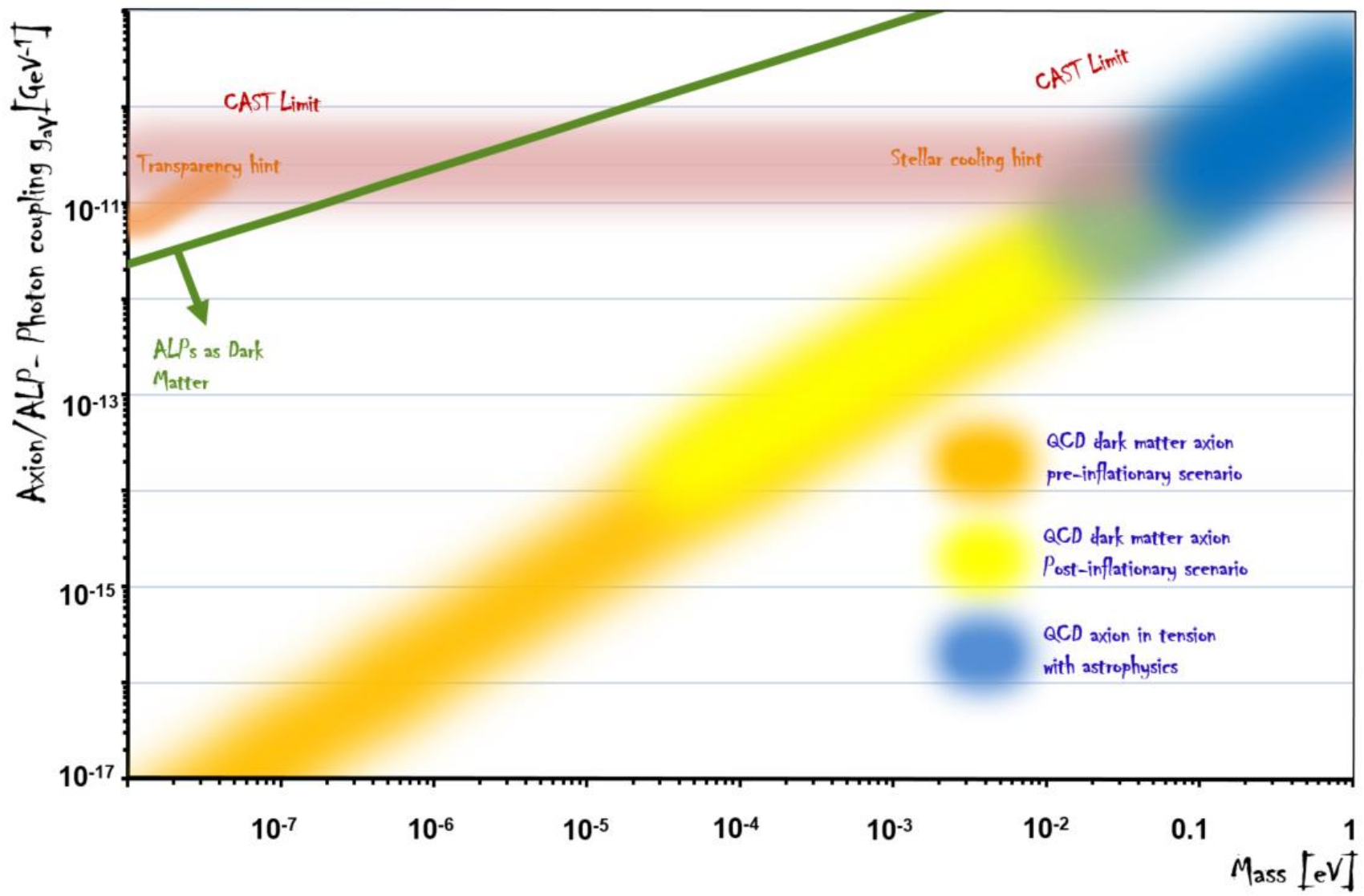
Axion (ALP) - Photon Coupling:

The Axion (ALP) carries same quantum numbers as η^0 and π^0

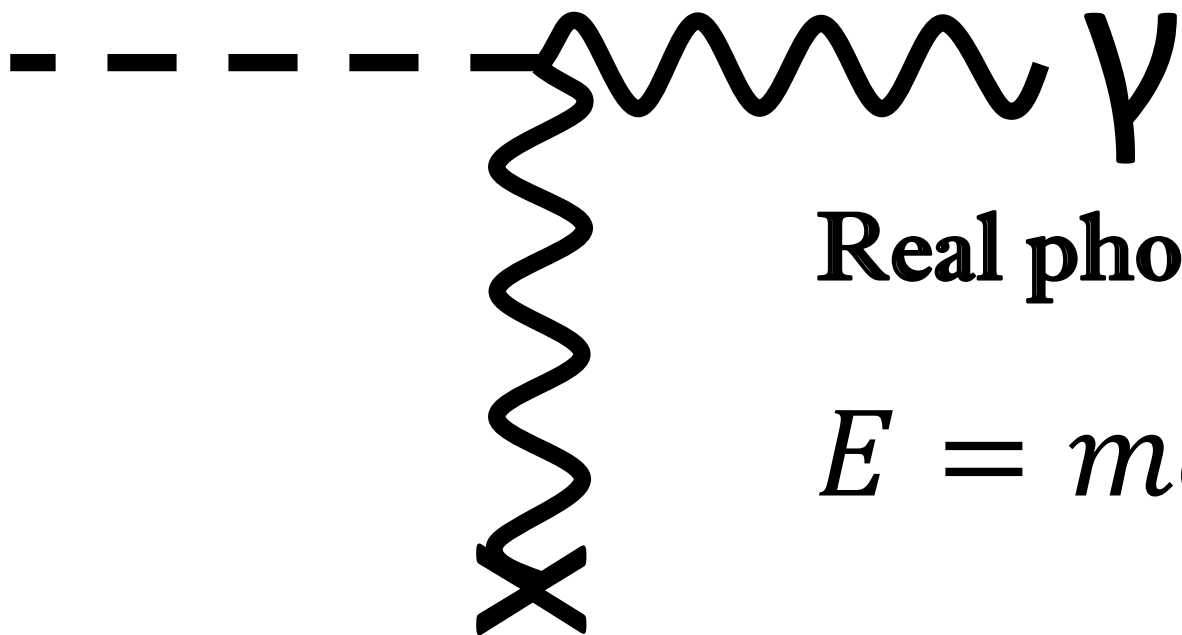


→ Quantum mechanical mixing with π^0 & η^0 → 2 photons!

Axion (ALP) - Landscape:



Axion detection: Primakoff Effect:



Real photon

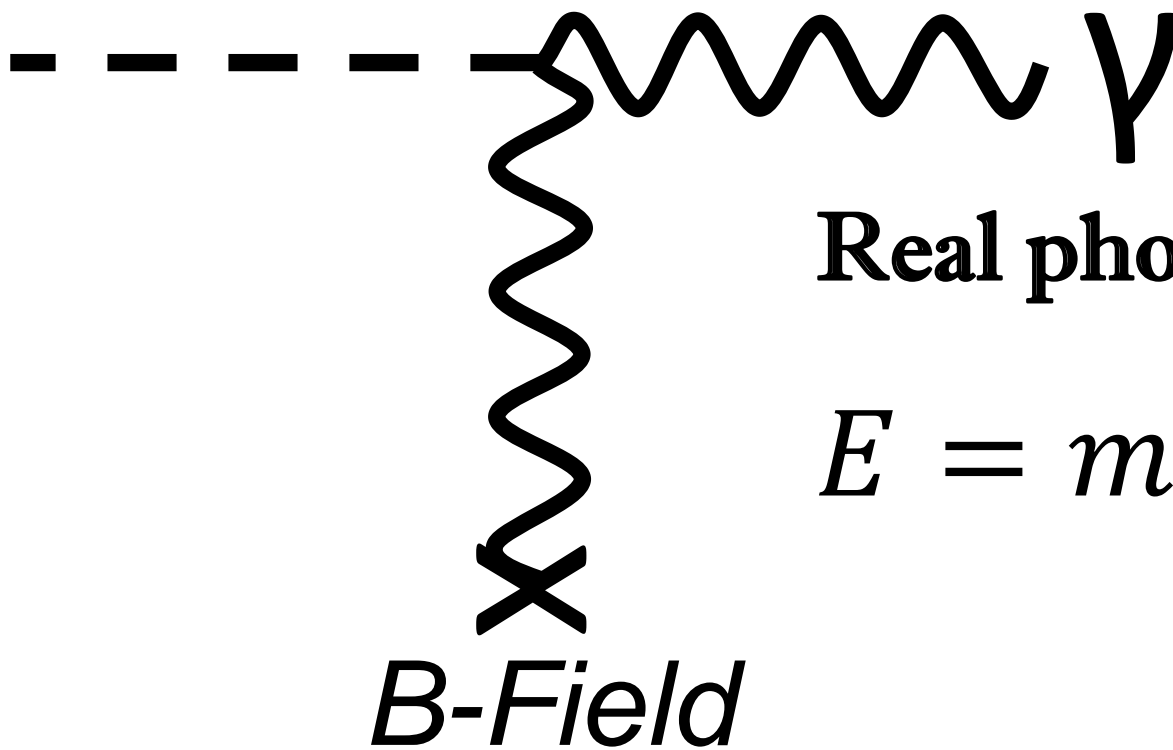
$$E = mc^2 = h\nu$$

B-Field

Suppressed
by $\frac{1}{f_a}$

$$\mathcal{L}_{a\gamma} = \frac{\alpha}{2\pi} C_{a\gamma} \frac{a(t)}{f_a} \mathbf{E} \cdot \mathbf{B}$$

Axion detection: Primakoff Effect:

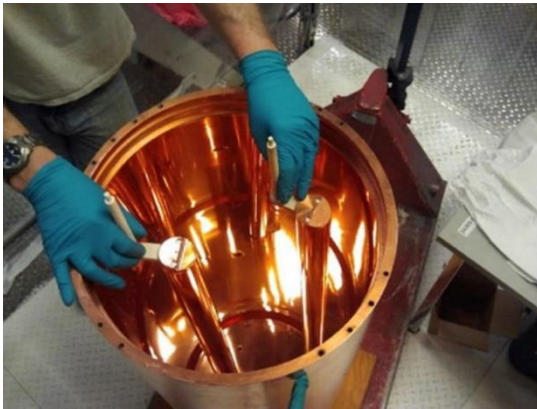


→ Axion in B-field sources E-field oscillations!
 Suppressed by $\frac{1}{f_a}$

Axion detection: Cavities in B-Field:

→ Use resonator to "pump cavity"

Adjusting resonance frequency: "Tuning Rod"



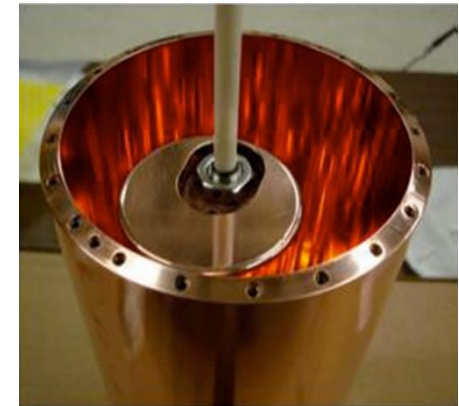
ADMX

U Washington, USA



CAPP

IBS, S. Korea

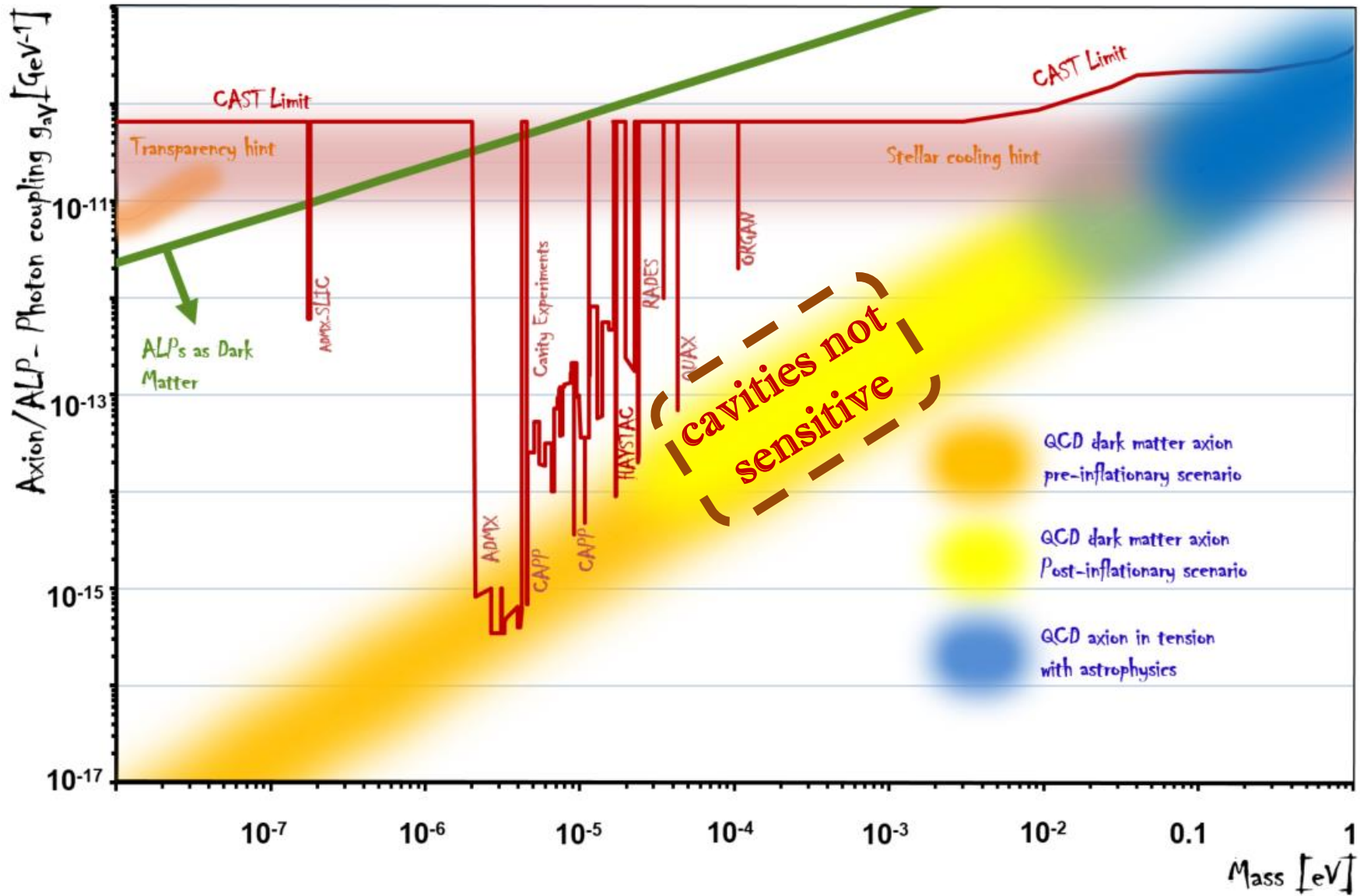


HAYSTAC

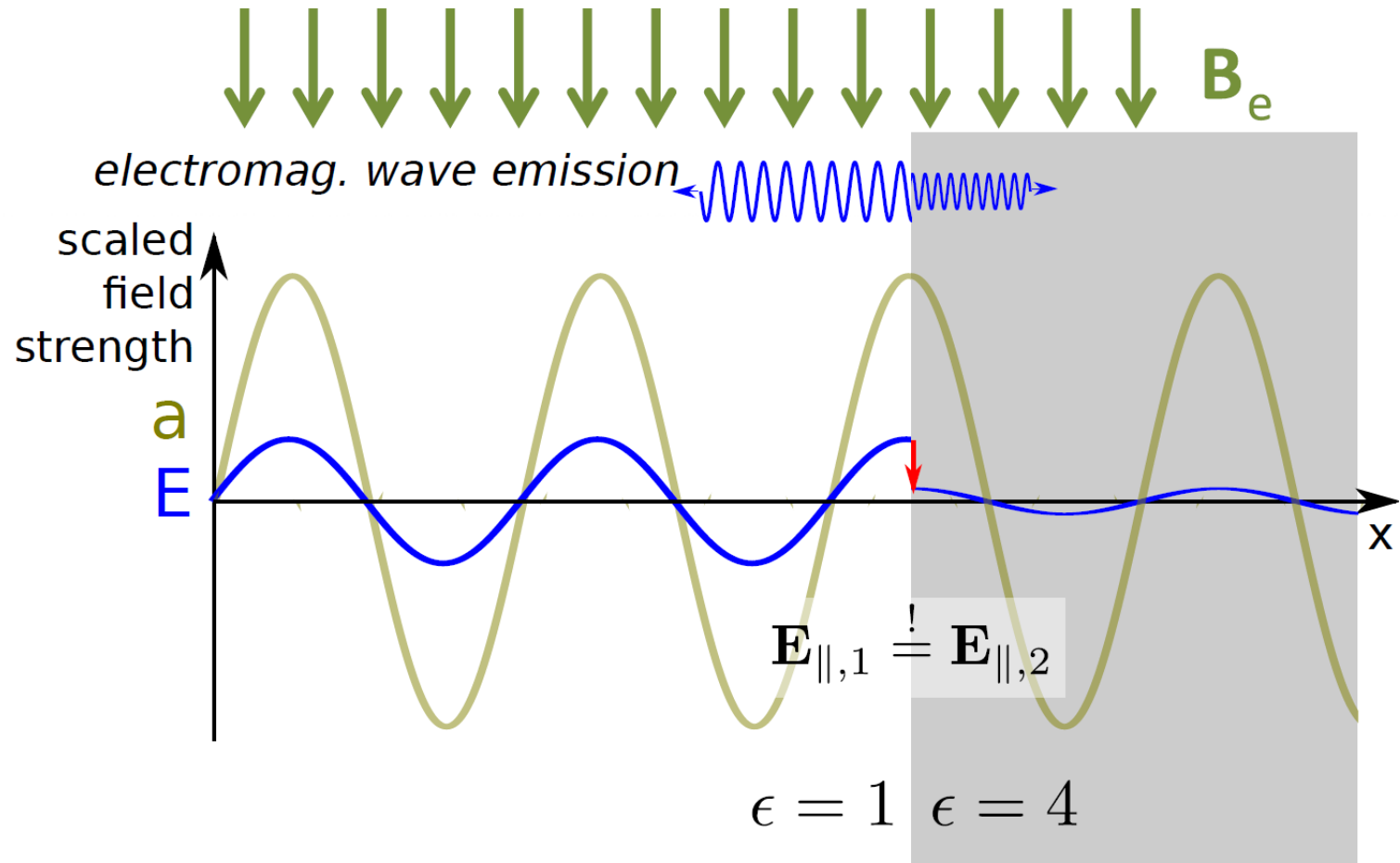
Yale University, USA

$$P_{sig} \propto B^2 V Q_{cav}$$

$$P_{sig}(B=6.8 T, V=136 l, Q=10^5) \sim 2 \cdot 10^{-22} W$$

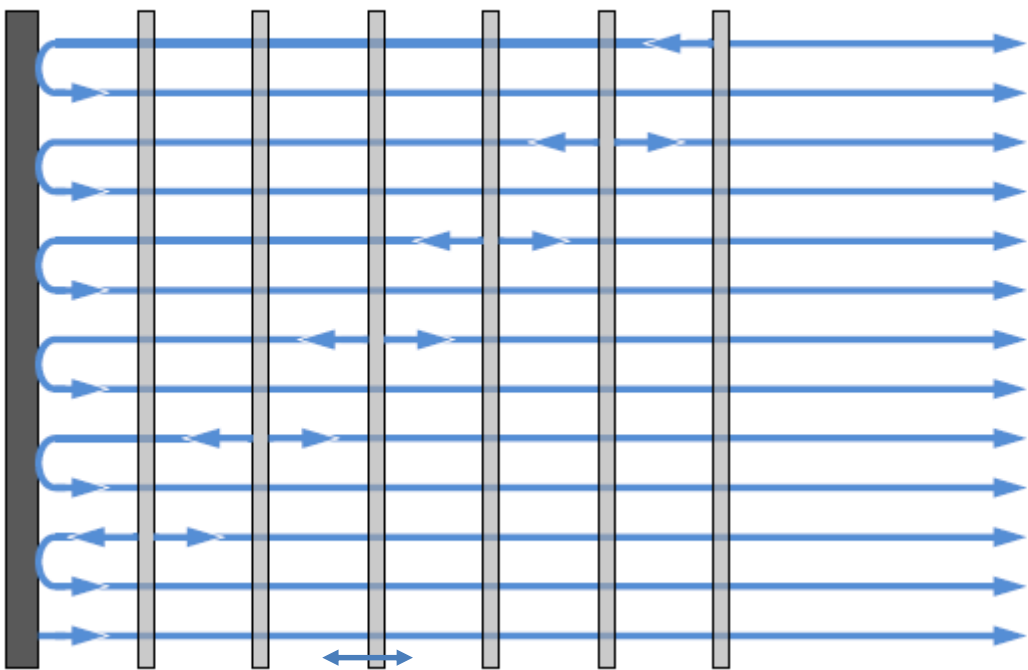


Dielectric Haloscope

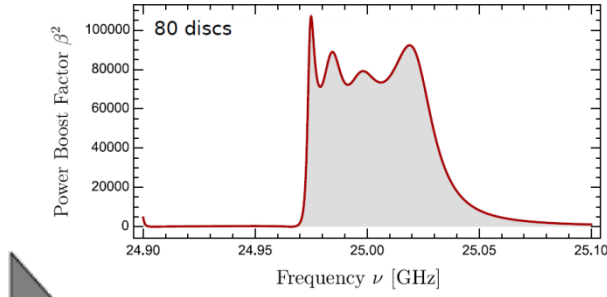


$$\left(\frac{P}{A}\right)_{\text{mirror}} \sim 2 \cdot 10^{-27} \frac{\text{W}}{\text{m}^2} \left(\frac{B_{\parallel}}{10 \text{ T}}\right)^2 (g_{a\gamma\gamma} m_a)^2$$

Dielectric Haloscope



Mirror $\sim \lambda/2$ Dielectric Disks



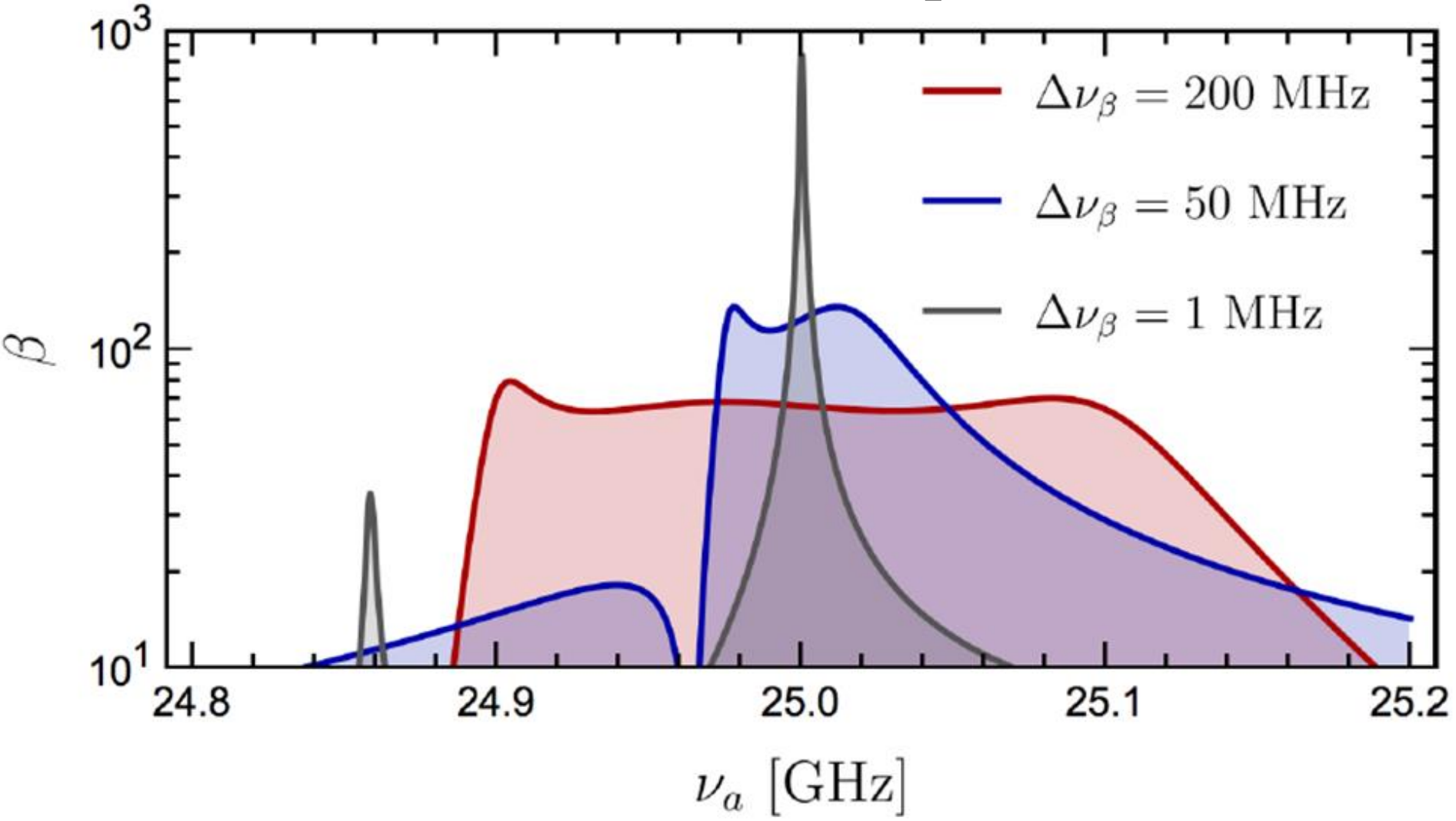
$$\beta^2 = \frac{P_{cavity}}{P_{mirror}}$$

Receiver

$$\left(\frac{P}{A}\right)_{cavity} \sim 2 \cdot 10^{-27} \frac{W}{m^2} \left(\frac{B_{||}}{10 T}\right)^2 (g_{a\gamma\gamma} m_a)^2 \beta^2$$



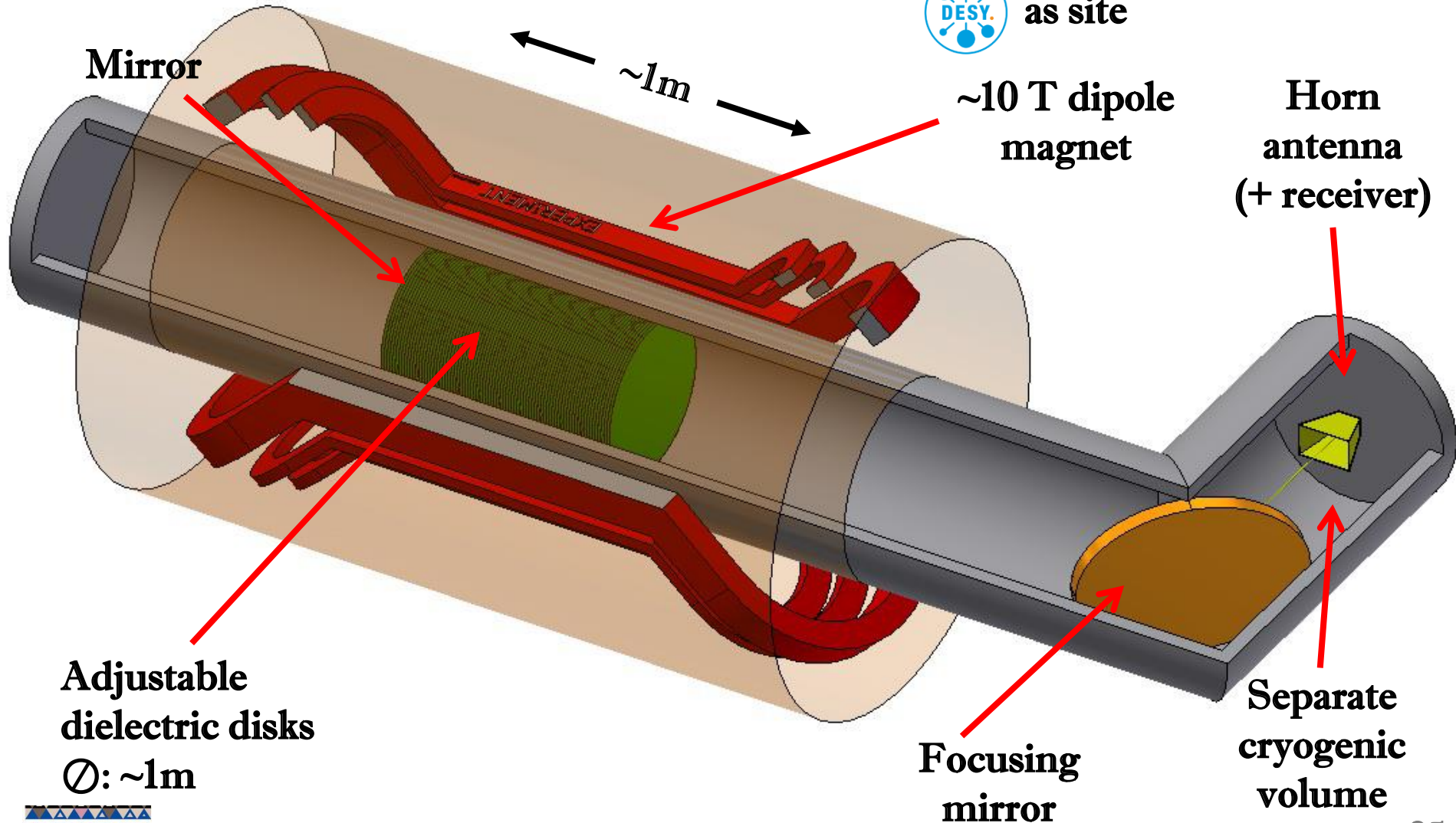
Dielectric Haloscope



Magnetized disk and Mirror Axion eXperiment



as site



Magnetized disk and Mirror Axion eXperiment

Collaboration forming on 18th Oct. 2017



MAD MAX institutes:



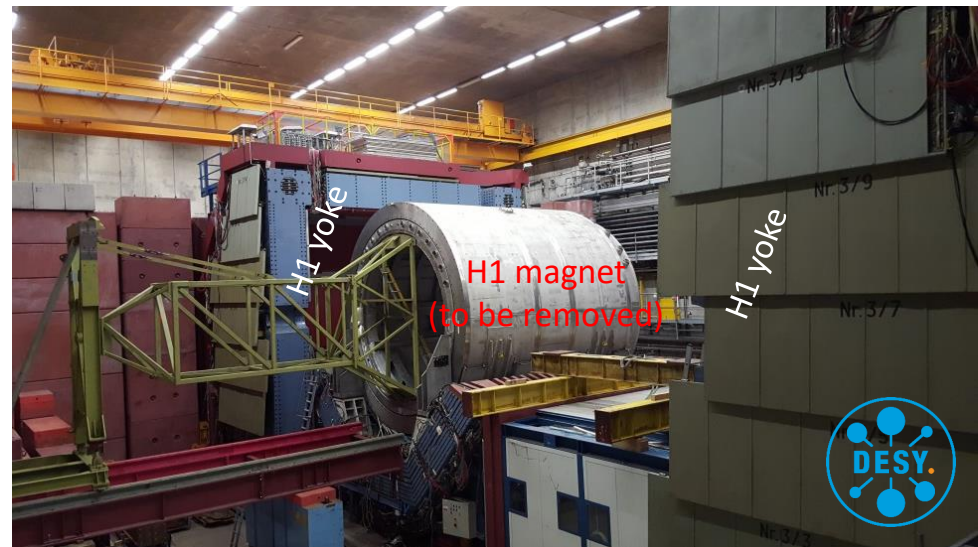
Designated Experimental Site



MADMAX to be operated at HERA Hall North

Make use of DESY infrastructure

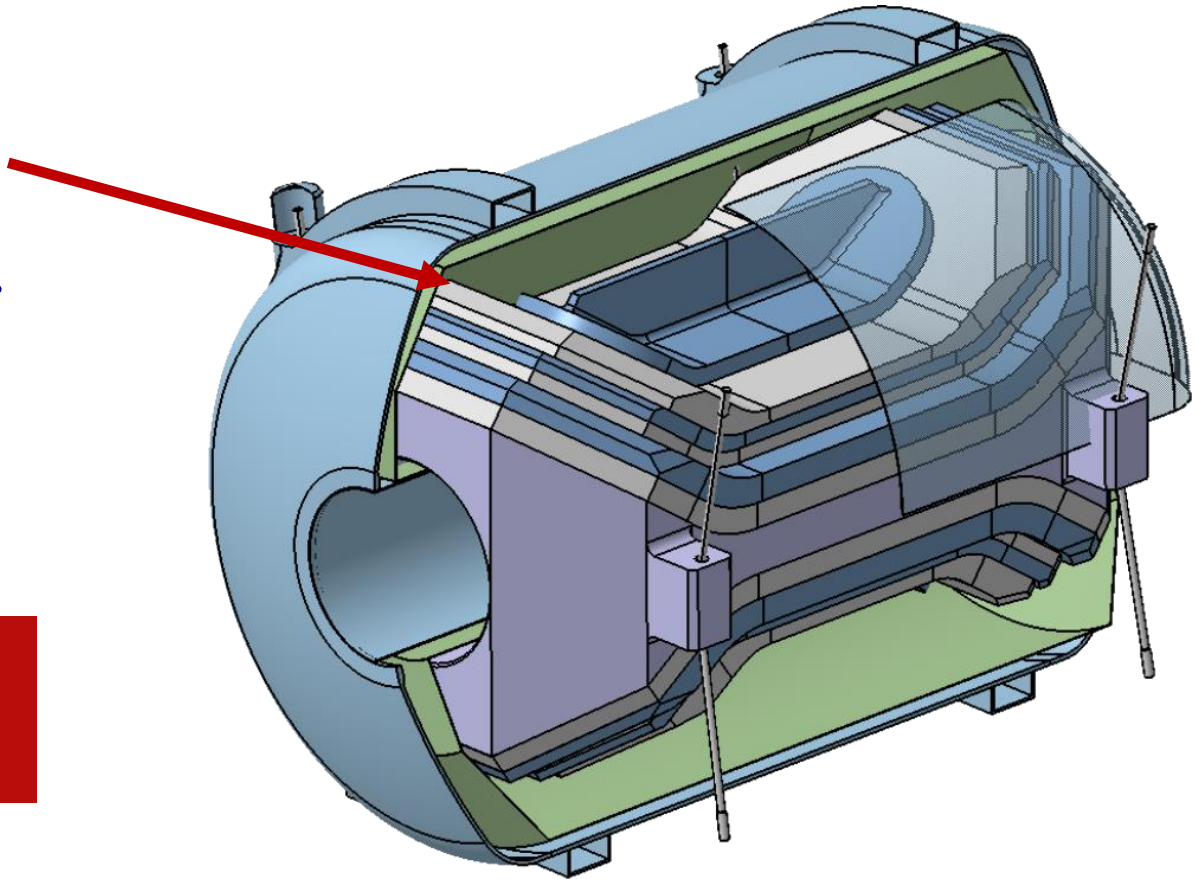
Benefit: re-use H1 yoke as magnetic shielding to reduce fringe field and increase B field



The magnet

Design and R&D on 9T large bore dipole magnet
FoM: $100 \text{ T}^2\text{m}^2$ stored energy $\sim 500\text{MJ}$

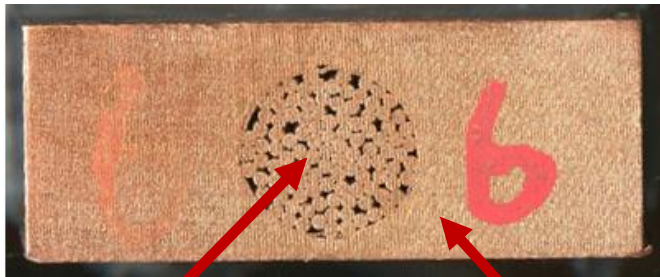
Baseline design:
2 * 9 "skateboard coils"
with novel
copper CICC conductor



The magnet

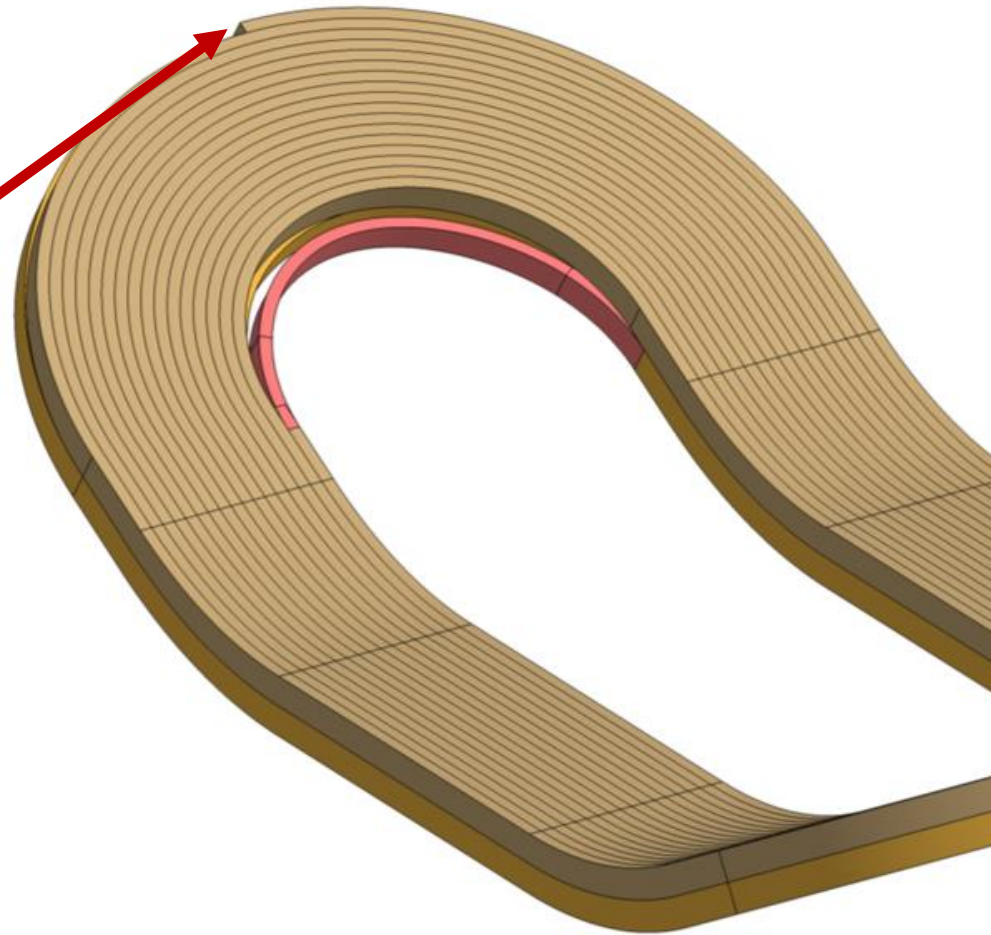
Design and R&D on 9T large bore dipole magnet
FoM: $100 \text{ T}^2\text{m}^2$ stored energy $\sim 500\text{MJ}$

Baseline design:
2 * 9 "skateboard coils"
novel **copper CICC conductor**



NbTi and Cu strands Copper doncuit

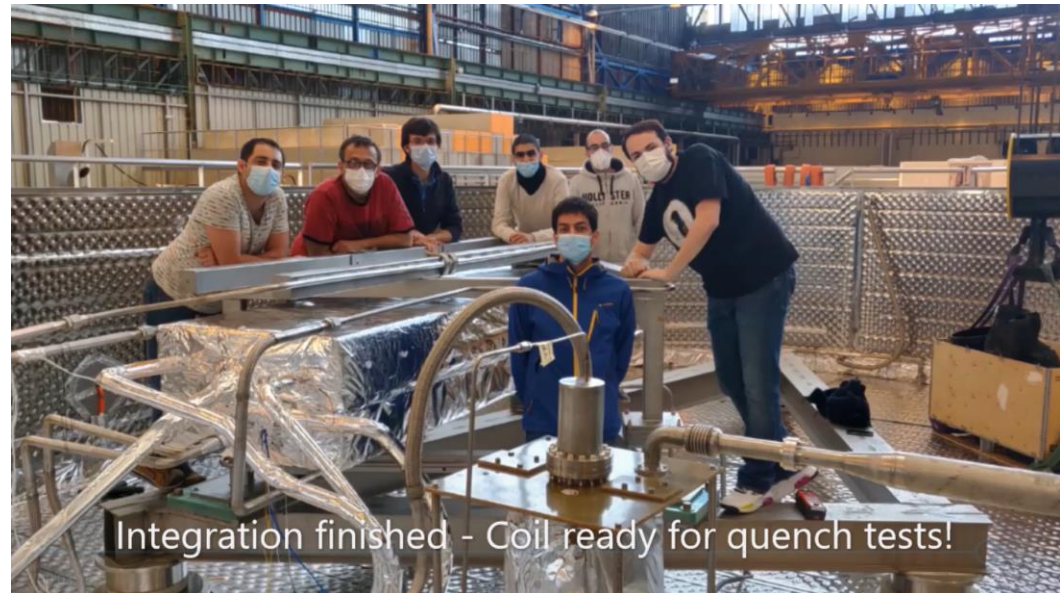
- ✓ Insertion tested
- ✓ Compaction \rightarrow right fill ratio
- ✓ Yield strength after cold work ok



The MAD MAX magnet

Main project risk: Quench detection

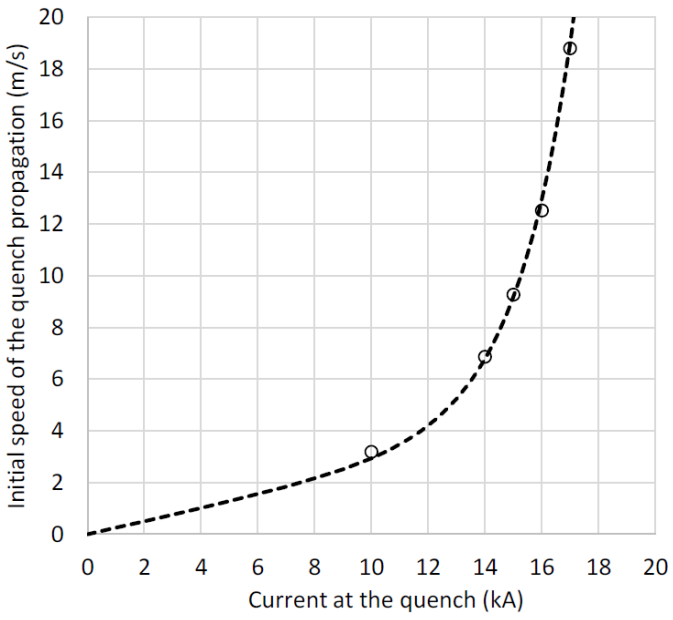
Dedicated magnet:
→ characterize quench propagation



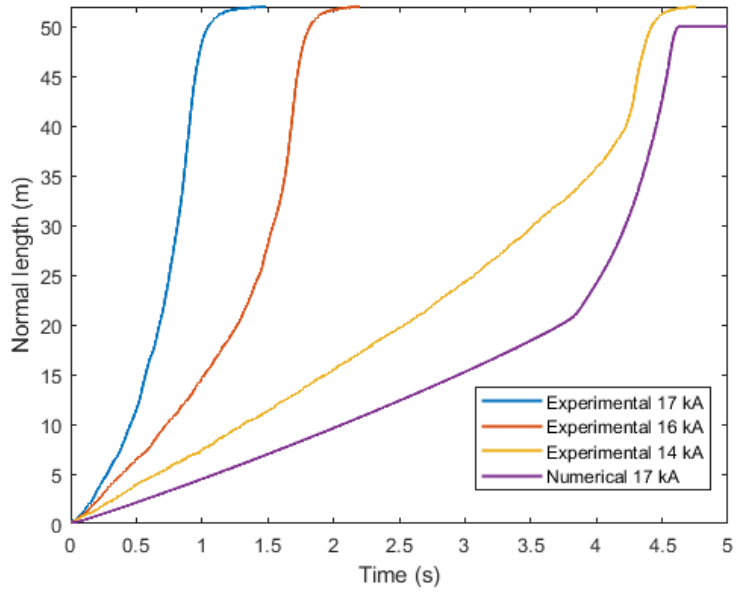
The magnet

Main project risk: Quench detection

Quench propagation velocity:



Thermo hydraulic quenchback:



**Quench propagation understood
→ Safe to operate magnet!**



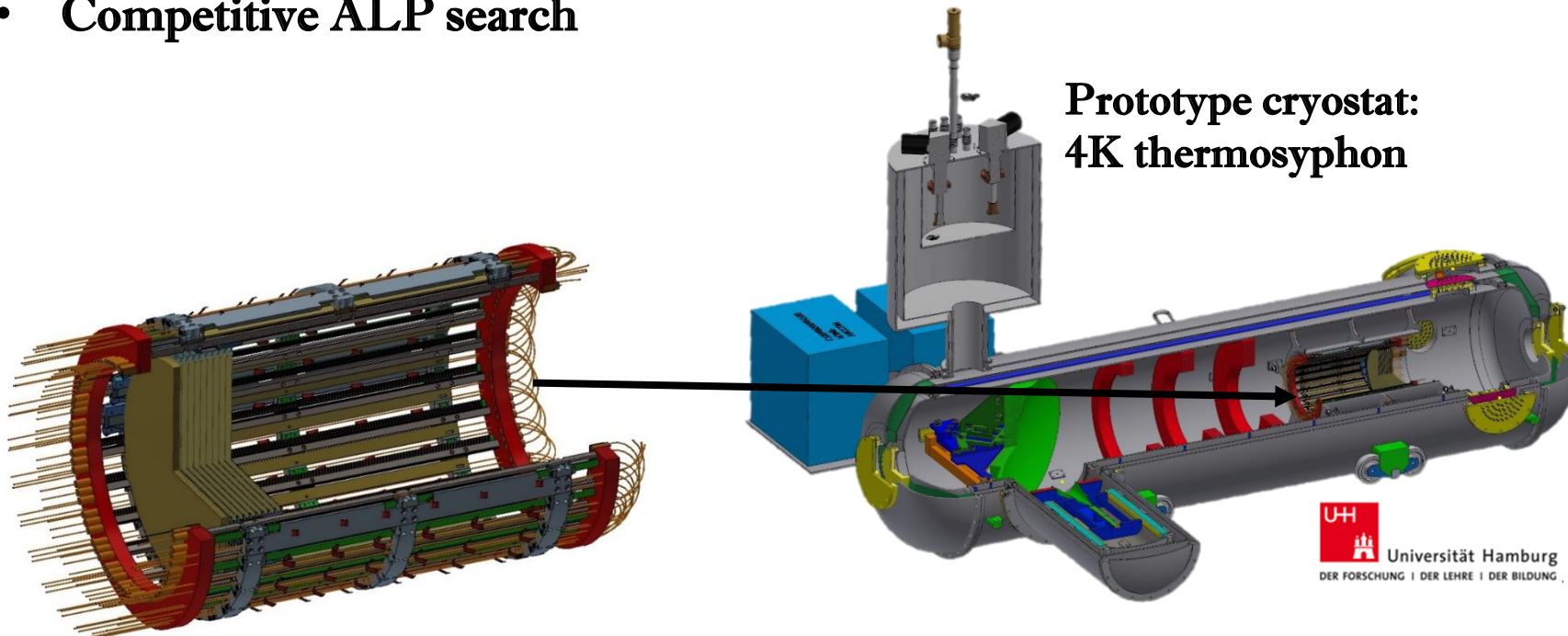
MADMAX prototype

Scaled down version of MADMAX:

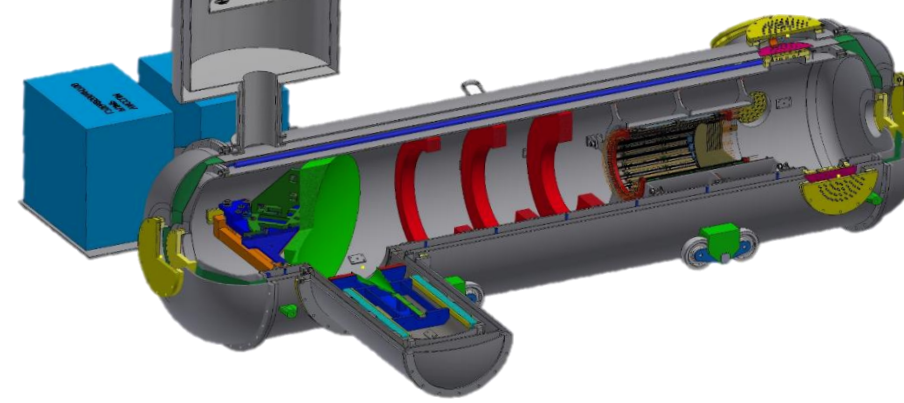
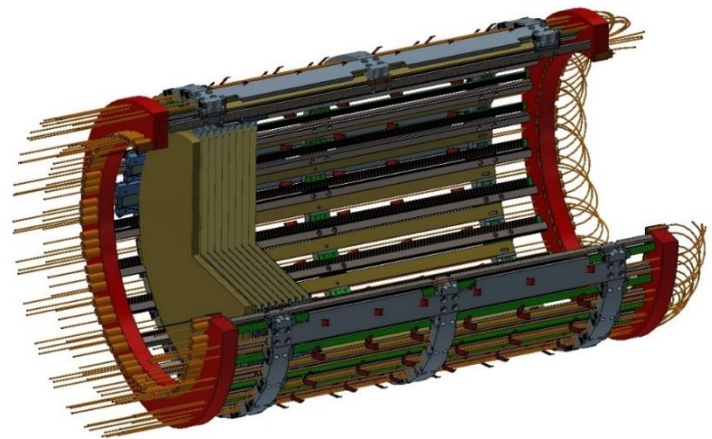
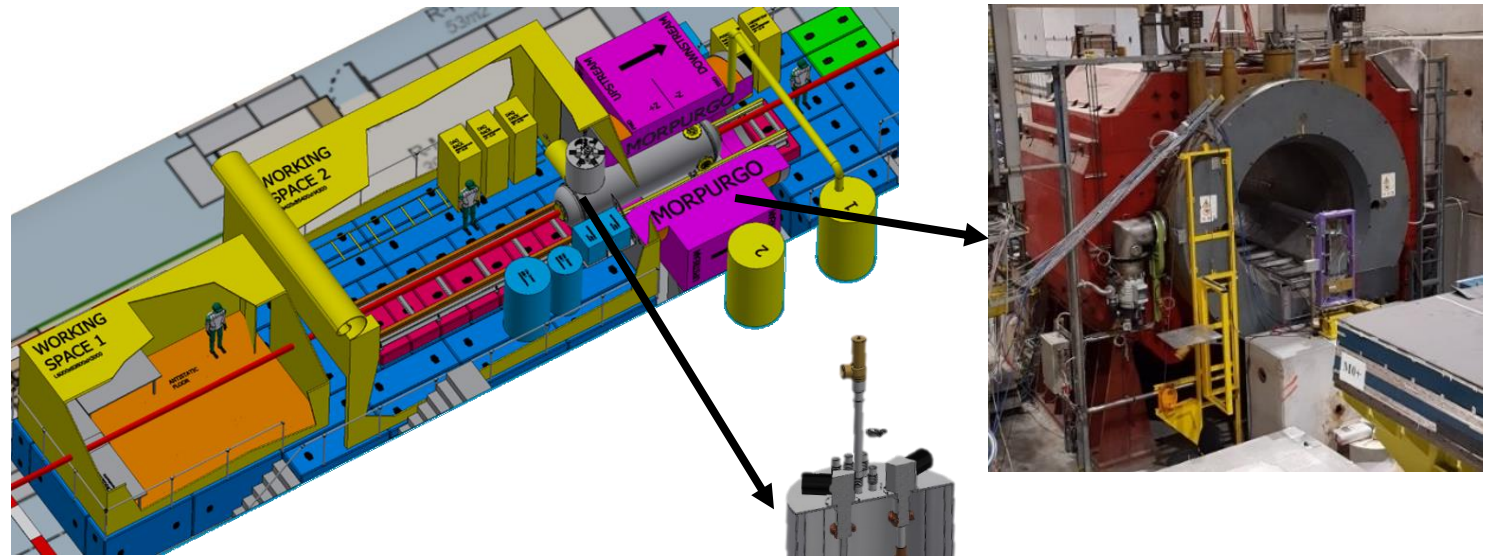
- Reduced number of disks, 1/16 disc area ($\varnothing:300\text{mm}$)
- 1/5 magnetic field (1.6 T, MORPURGO @ CERN)

Main goals

- Demonstrating and prototyping key technologies
- Competitive ALP search



MAD MAX prototype



The MADMAX mechanics

Feasibility of disc motors at 4K in strong B-field

✓ Developed Piezo motor drive unit

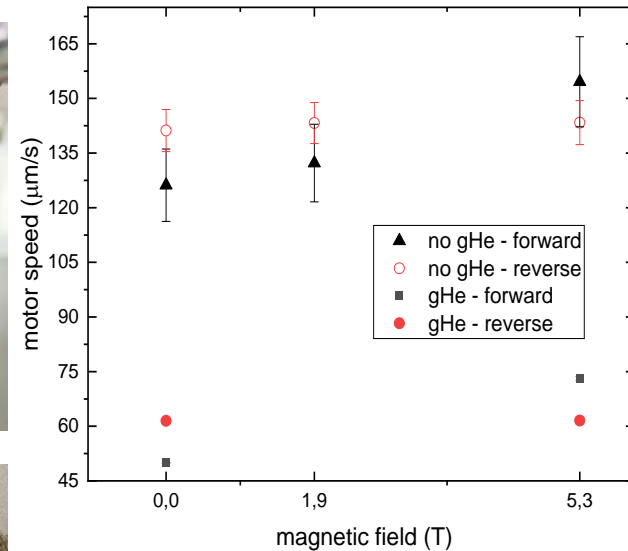
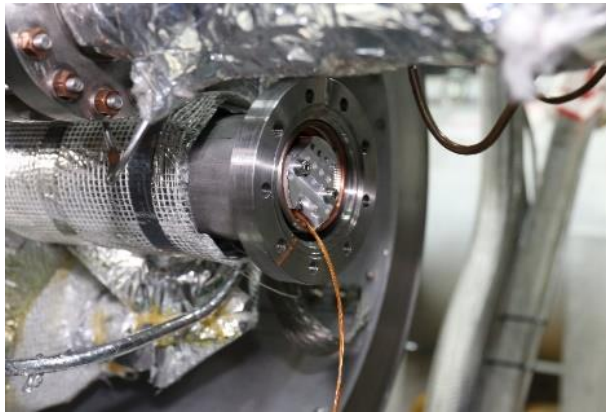
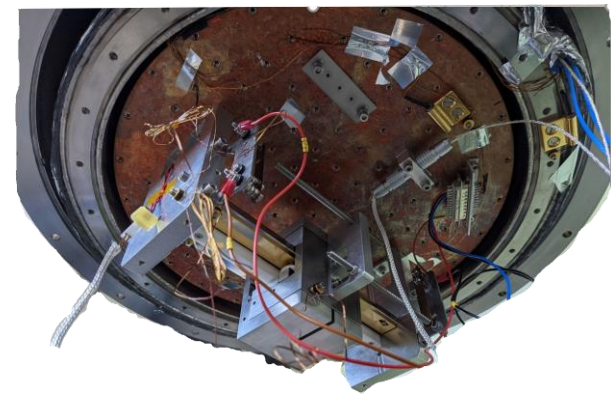
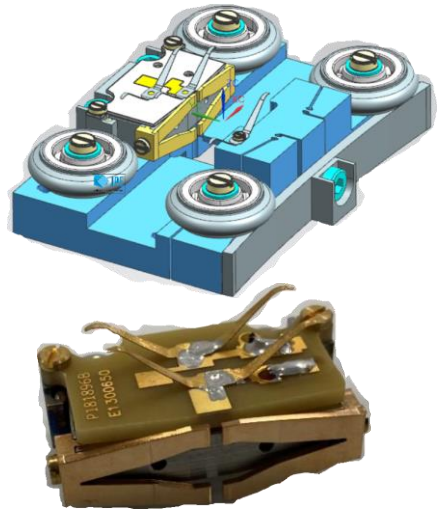
Repeatability < 1μm

✓ Characterized at 4.2 K ambient temperature

Speed > 0.1mm/sec

✓ Tested at 5 K in 5.3T magnet at DESY:

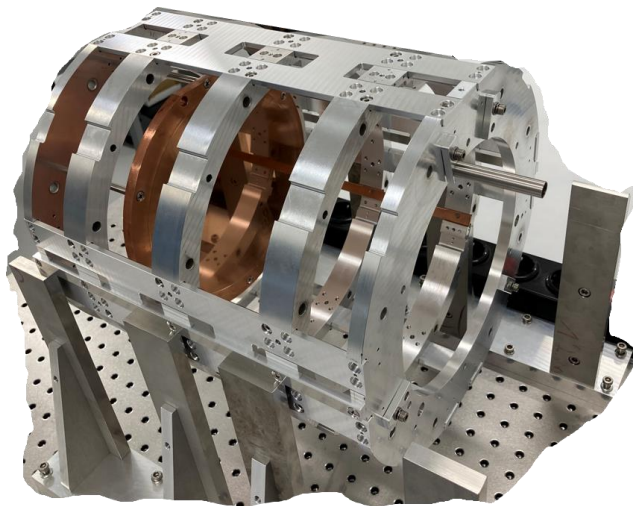
Moves reliably



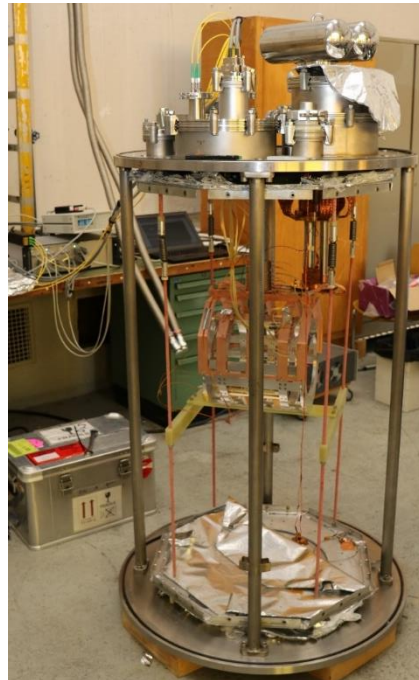
The MAD MAX mechanics

**Mechanical test bed:
Verify mechanical feasibility of
Baseline design**

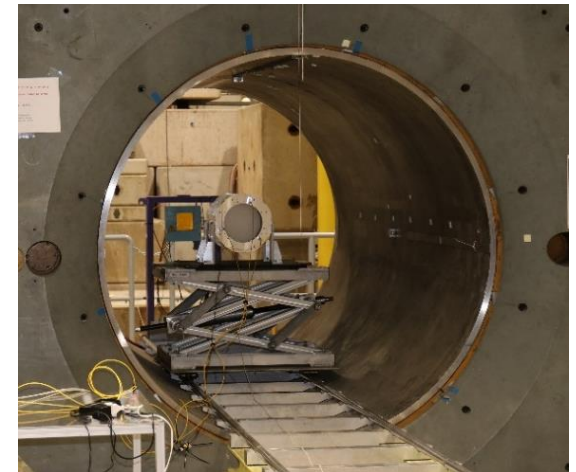
**Disc movement with required
accuracy at cold in high B-field**



**Project 200:
One disc (200mm)
In front of mirror**



Operated in cryostat



**Operated in 1.6T MORPURGO
magnet at**



The RF response

Closed booster 100

- Three sapphire discs 100m
- Resonant configuration
- Coupling to LNA via taper



parabolic taper

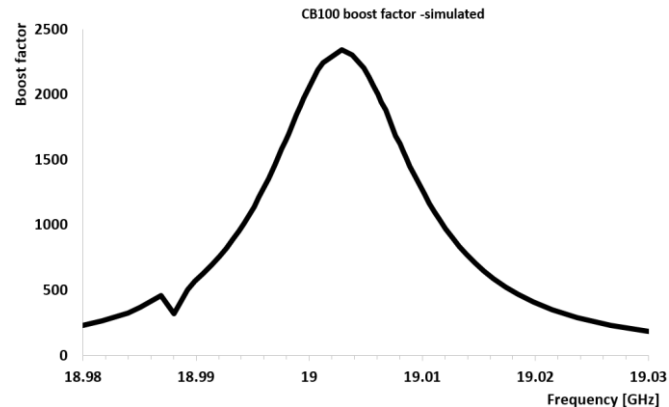
J. Doane, Int. J. Infrared
Milli. Waves 5 (1984)



spacing ring

sapphire

copper mirror



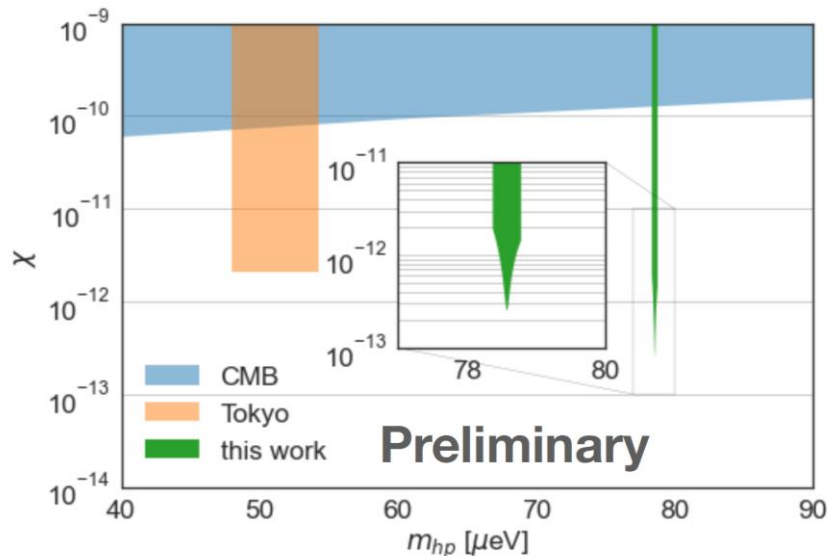
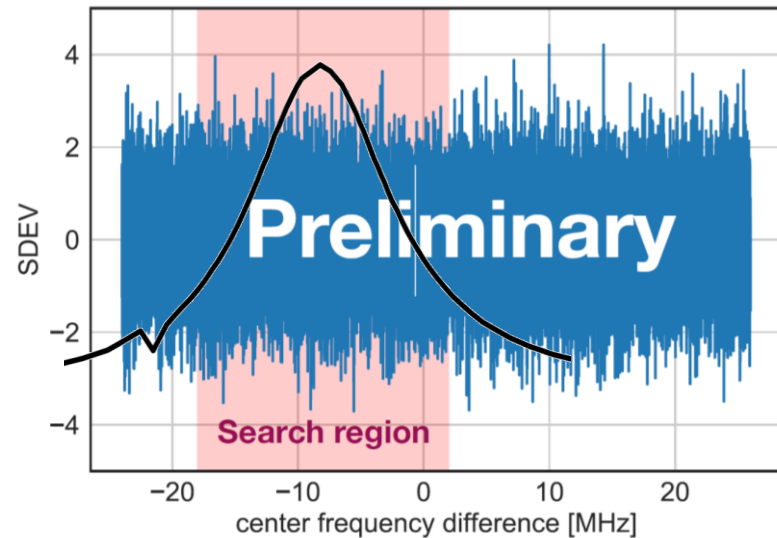
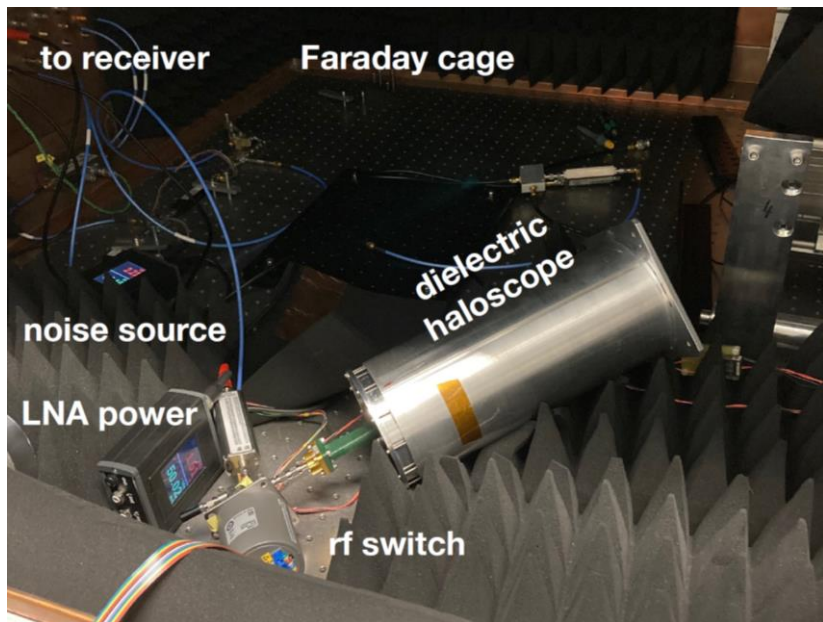
- Understand RF behavior in well defined boundary conditions
- Develop calibration procedure: reflectivity and system temperature
- First ALPS and HP measurement

First $\text{AD} \text{---} \text{MAX}$ HP measurement

Faraday cage at lab



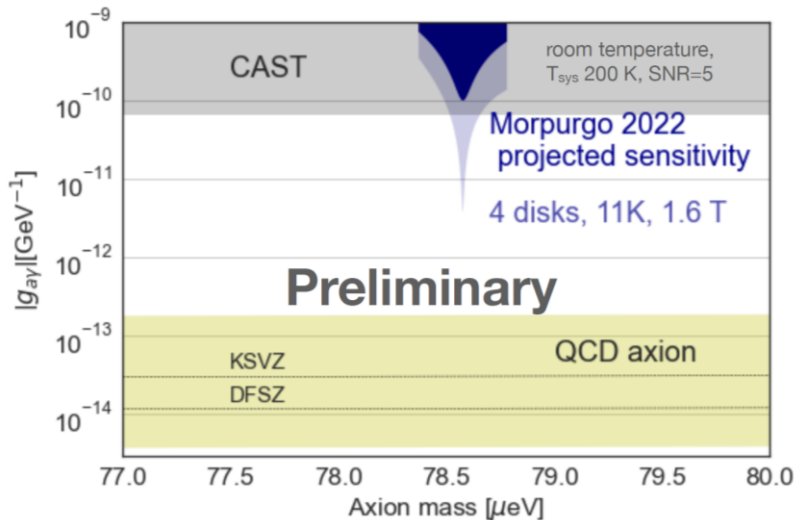
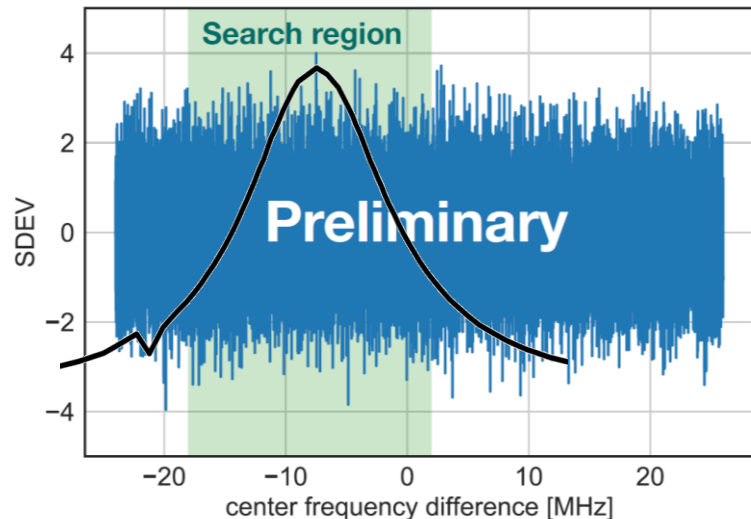
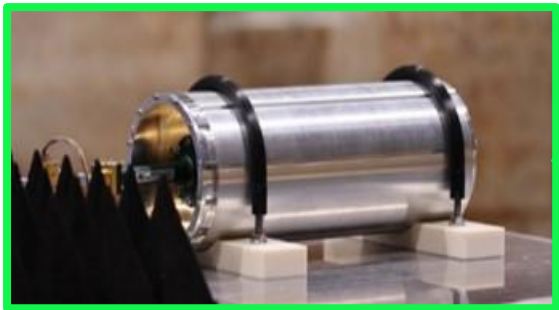
$B = 0 \text{ T}$, $t = 32 \text{ days}$



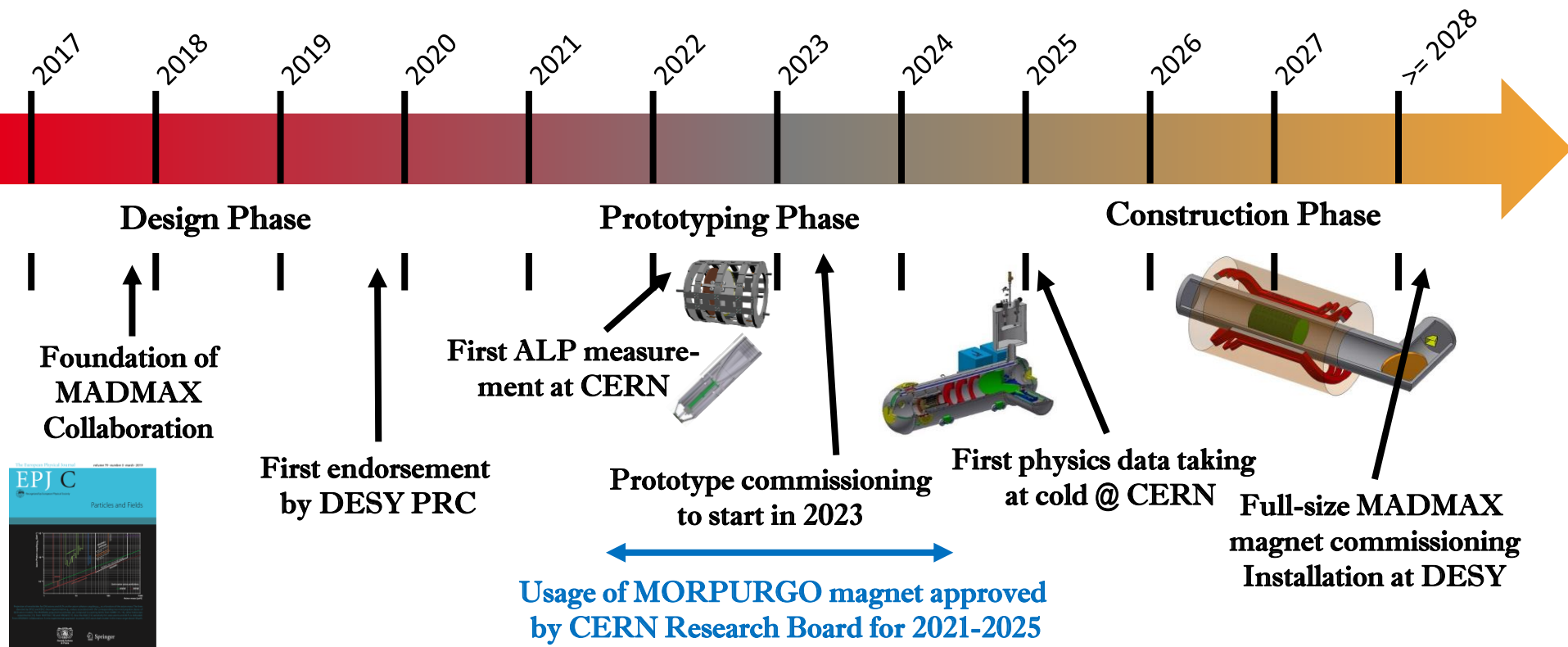
First $\text{AD} \text{---} \text{MAX}$ ALP measurement

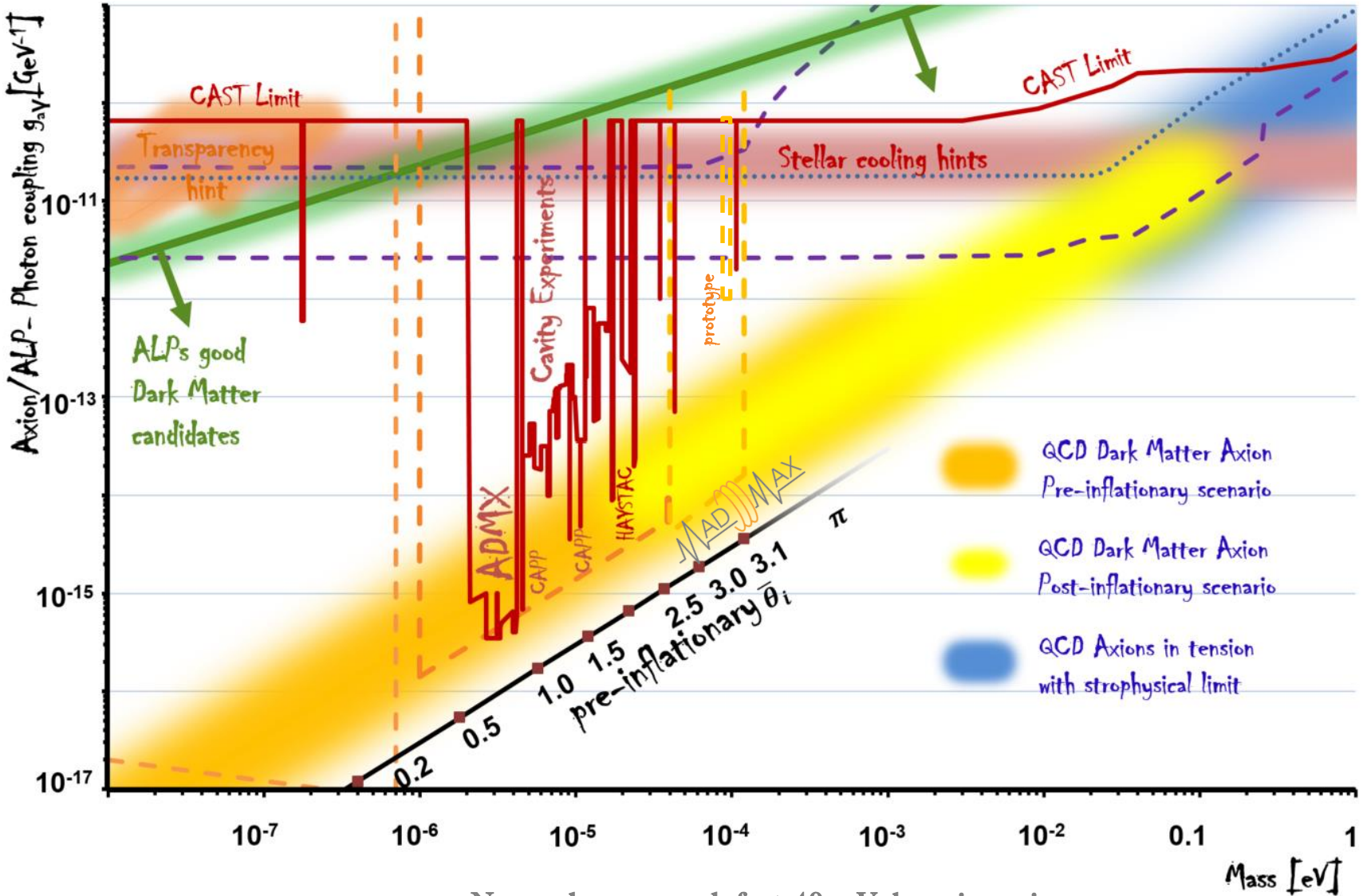
$B = 1.6 \text{ T}$, $t = 10.5 \text{ hours}$

MORPURGO magnet at 

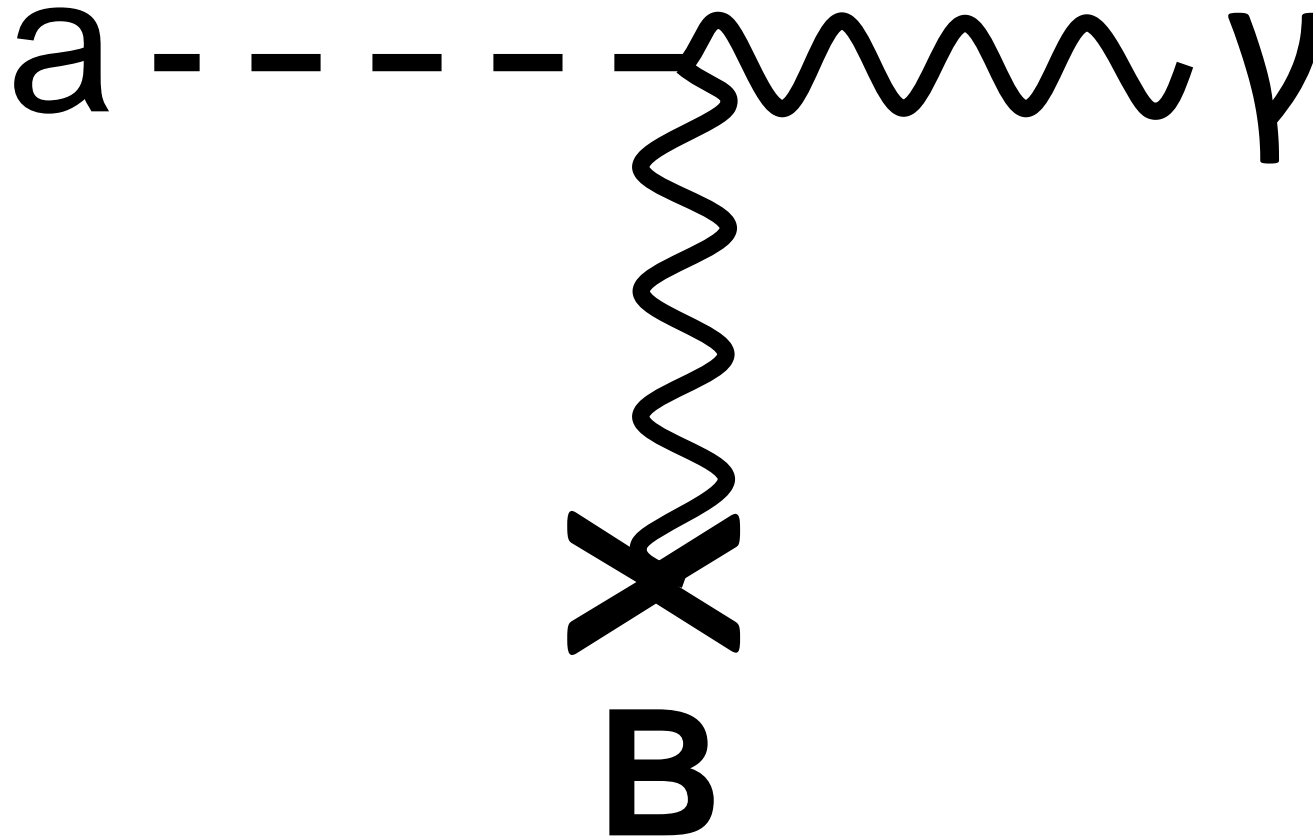


MADMAX timescale

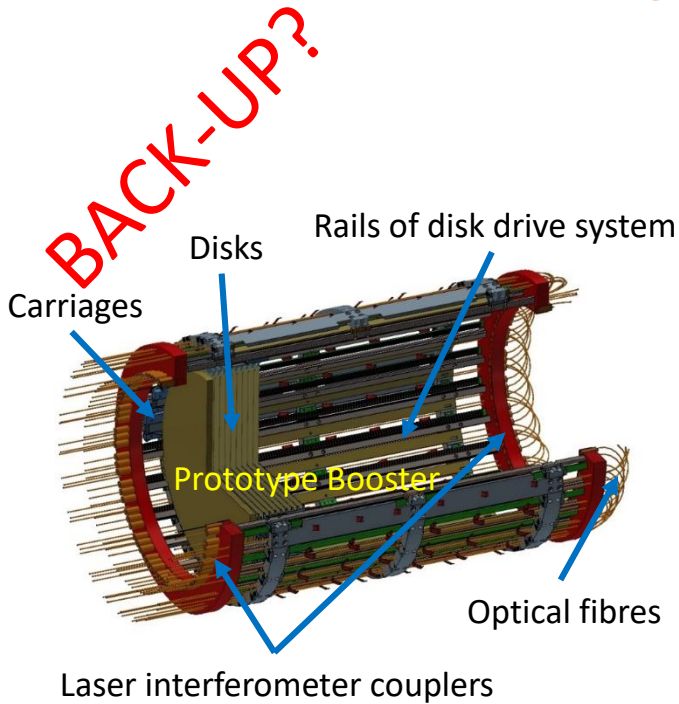




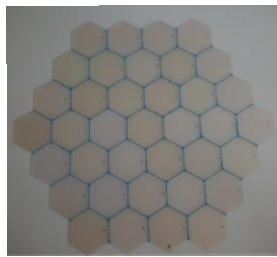
Note: other proposals for $>40 \mu\text{eV}$ detection exist:
plasma haloscope, dish antenna, ...



The booster



- Booster is the heart of MADMAX : a mirror and several adjustable dielectric discs
- Operating conditions:
 - Cryogenic temperatures: 4 K
 - High magnetic field: up to ~ 10 T
 - Vacuum or cold gehe exchange gas
- Disk weight: 600 g for $\varnothing 300$ mm
- **Piezo-driven actuator system with feedback from laser interferometer with absolute precision**
- Candidate disk materials:
 - LaAlO_3 ($\epsilon \approx 24$, $\tan\delta \approx \text{a few } 10^{-5}$)
 - Sapphire ($\epsilon \approx 9$, $\tan\delta \approx 10^{-5}$)
- LaAlO_3 available as 3" wafers at maximum
- **Tiling necessary → Semi-automatic gluing machine**



Simulations

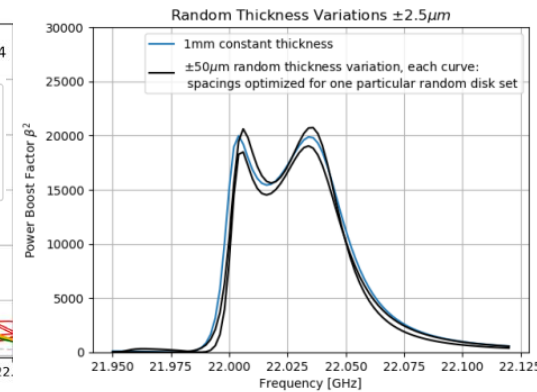
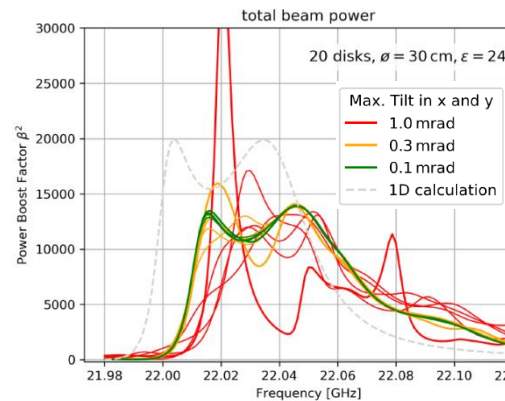
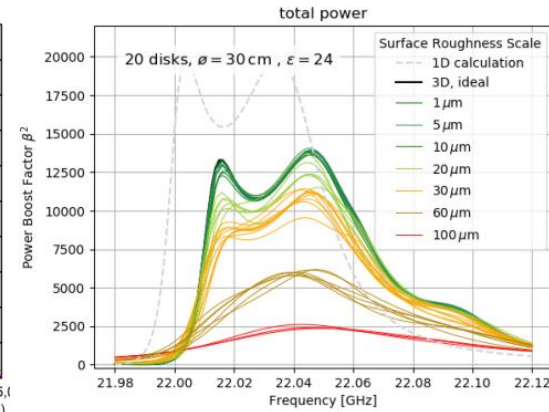
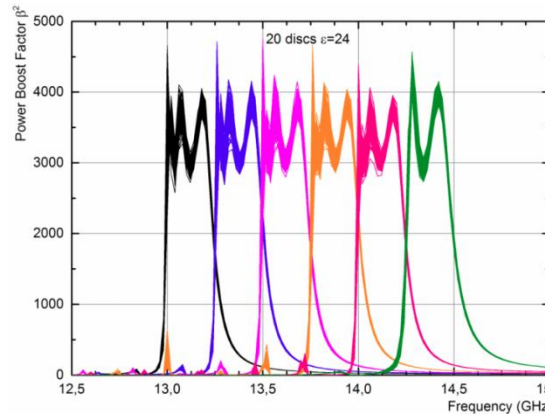
Calculations of achievable boost

3D FEM (COMSOL+Elmer), 2D3D FEM (radial symmetry), Beam propagation

Investigated effects:

- ✓ 3D effects (diffraction)
- ✓ Dielectric loss
- ✓ Inaccuracies (positioning, roughness, tilts, thickness,...)
- ✓ DM velocity dispersion
- Tiling of disks
- Coupling to antenna, receiver
- Calibration of boost factor

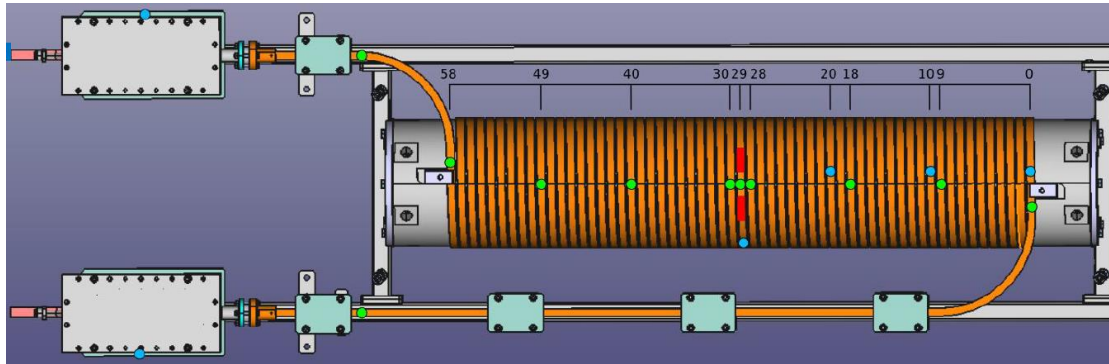
Prototype booster



The MAD MAX magnet

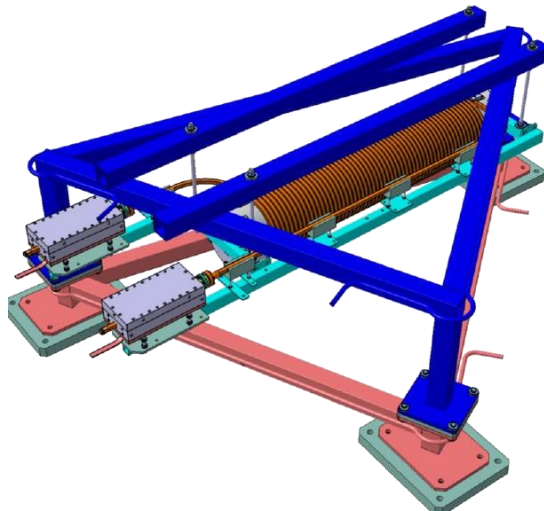


Design and R&D on 9T large bore dipole magnet
Quench protection: **demostrator magnet "MACQU"**



Production at NOELL,
integration at CEA

Design in cooperation between CEA and NOELL



R&D on mechanical feasibility

Cryostat for prototype (2024 MORPURGO run)

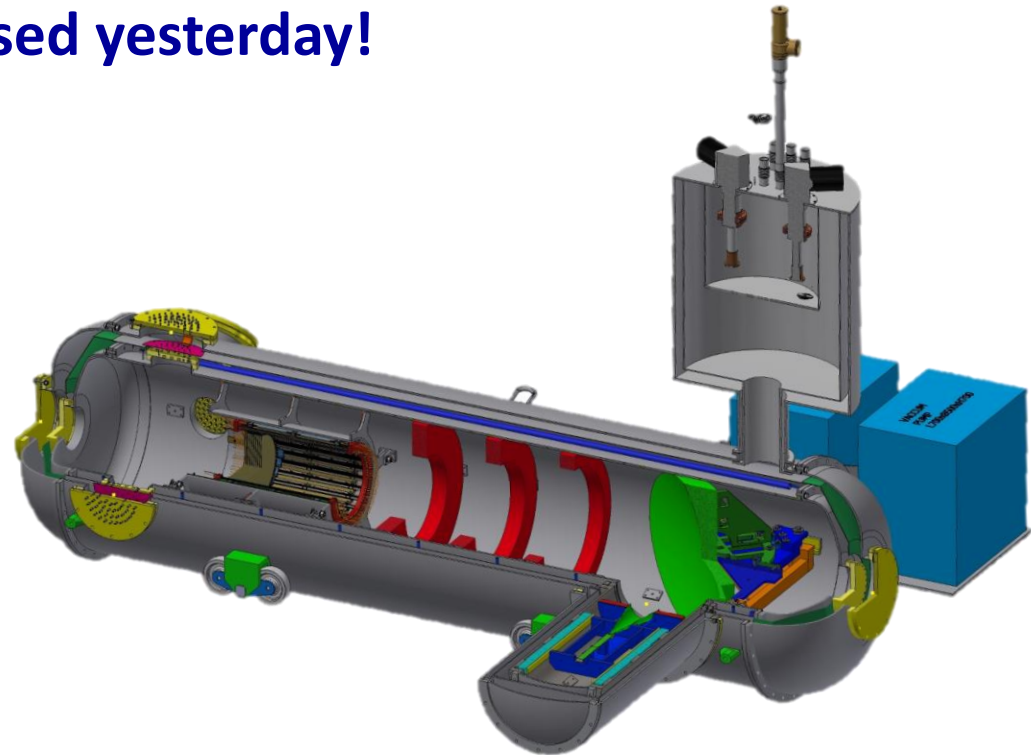
Financing of cryostat secured!

Tender ongoing:

application deadline passed yesterday!



→ More news from C. Krieger



Closed Booster

Name	acronym	disc diameter [mm]	Nr. of discs	Availability
Closed booster 100	CB100	100	3	2021
Closed booster 200	CB200	200	≥ 3	2022
Project 200	P200	200	1	2021
Reduced booster	r-booster	300	≥ 3	2023
Prototype booster	P-booster	300	20	2024
Final booster		1250	80	>2025

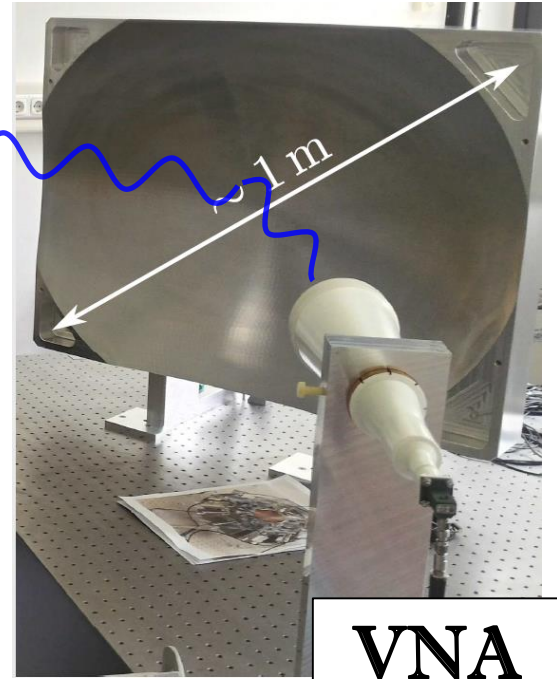
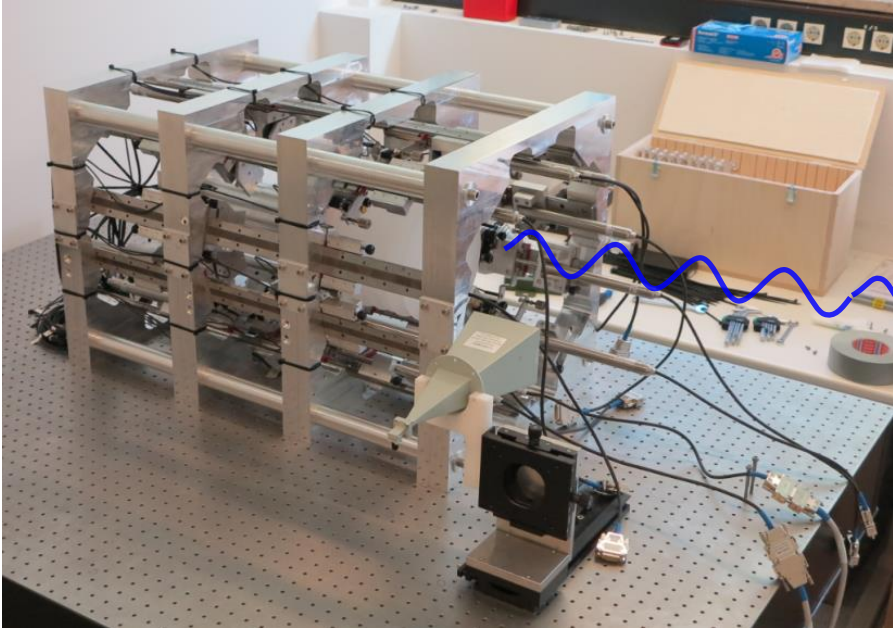
CB100 RT tests with HEMT preamp in 2022:
Quantification of system temperature
Work in progress!

CB100 4K tests with VNA
Reflectivity behavior at 4K
according to specs
Very good time stability
Work in progress!



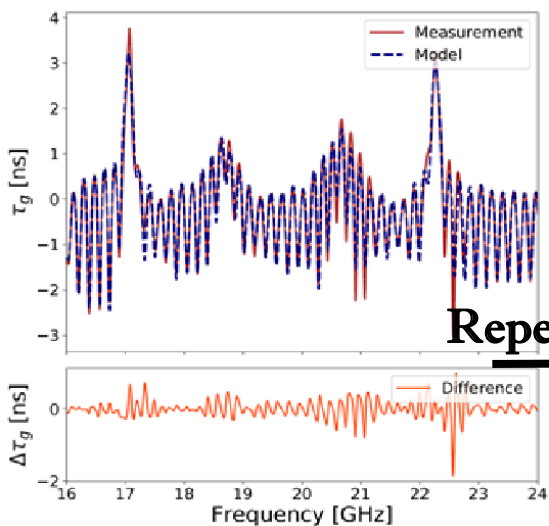
Proof of Principle Setup

Verification of experimental feasibility:



VNA

- Reproducibility
- Correlations
reflectivity vs. boost
- Positioning algorithm



Repetitions →

